

Reynolds-Averaged Two-Fluid Model prediction of moderately dilute fluid-particle flow over a backward-facing step

M. Riella*, R. Kahraman, G.R. Tabor

College of Engineering, Mathematics and Physical Sciences, University of Exeter, North Park Road, Exeter EX4 4QF, UK

Abstract

In this work a Reynolds-Averaged two-fluid fully coupled model (RA-TFM) for modelling of turbulent fluid-particle flow is implemented in OpenFOAM and applied to a vertically orientated backward-facing step. Three particle classes with varying mass loadings (10-40%) and different Stokes number are investigated. Details of the implementation and solution procedure are provided with special attention given to challenging terms. The prediction of mean flow statistics are in good agreement with the data from literature and show a distinct improvement over current model predictions. This improvement was due to the separation of the particle turbulent kinetic energy k_p , and the granular temperature Θ_p , in which the large scale correlated motion and small scale uncorrelated motion are governed by separate transport equations. For each case simulated in this work, turbulence attenuation was accurately predicted, a finding that is attributed to separate coupling terms in both transport equations of k_p and ε_p .

Keywords: Turbulence attenuation, Backward-facing step, RA-TFM, OpenFOAM

*Corresponding author

Email address: mjr214@exeter.ac.uk (M. Riella)

1 1. Introduction

2 Modulation of turbulence is a complex two-way coupled phenomenon [14] and
3 can be caused by fluid-particle interaction and/or particle-wall interaction. Modu-
4 lation can result in an increase in the fluid-phase fluctuating velocities [24] due to
5 particle vortex shedding [40], which is caused by a large particle Reynolds number,
6 Re_p . Conversely, modulation of turbulence can result in the reduction of fluid-phase
7 fluctuating velocities, i.e. attenuation. This behaviour is prevalent in fluid-particle
8 flows due to high density ratios ($\rho_p \gg \rho_f$). This leads to the mean-feedback ef-
9 fect through drag - their primary coupling mechanism. Turbulence attenuation as
10 a result of small heavy particles in the carrier flow is well established in the litera-
11 ture [15, 24, 26, 33, 50, 54, 57] and has been shown to be further influenced by the
12 inhomogeneity of wall-bounded flow [53].

13 To date there have been numerous experimental studies investigating small heavy
14 particles in wall-bounded, high Reynolds number flow [4, 9, 33, 50]. One valuable
15 study is that of Fessler and Eaton [19] in which mean and turbulence statistics
16 of dilute [14] fluid-particle flow were recorded in a vertically orientated backward-
17 facing step. They report turbulence attenuation across three particle classifications
18 (different Stokes number and mass loadings) and provide valuable insights into the
19 particle behaviour in the free shear layer. Traditionally, the backward-facing step
20 has been used as a benchmark for validation of single-phase turbulence models, as
21 flow separation, reattachment and redevelopment are rife in engineering applications.
22 Due to the complex nature of turbulence attenuation and the challenging physics in
23 a backward-facing step configuration, its successful prediction has proven difficult
24 [10, 36, 38, 51, 58].

25 There are two main approaches for predicting macroscale fluid-particle systems:
26 the Eulerian-Lagrangian (E-L) method in which the fluid-phase is solved in an Eu-
27 lerian frame and the particle-phase is solved with Lagrangian equations. Typically,
28 all scales of motion are resolved except the boundary layers on the particle surfaces
29 resulting in a high resolution of the flow field. It follows that E-L simulations are
30 used for understanding fundamental phenomenon e.g. clustering [6, 49], and verify-
31 ing experimental observations [25, 34]. Due to their expensive cost as each particle
32 requires its own momentum equation, large particulate systems with high Reynolds
33 number become inviable. This leads to the second approach: the Eulerian-Eulerian
34 (E-E) methodology models both the fluid- and particle-phase as interpenetrating
35 continua resulting in both phases acting as “fluids”. This reduces the computational
36 cost considerably with the fully resolved scales of E-L being modelled. This approach
37 then relies on constitutive relations to close the governing equations.

38 Numerous two-fluid (E-E) models have been derived using a one-step averaging

39 process [1, 13, 28, 47] e.g. volume- or time-averaging. Within this methodology,
40 kinetic theory is used to close the particle-phase stress that appears in its momen-
41 tum equation. This approach has been applied by many authors [2, 11, 12, 16, 27,
42 40, 52, 60] with varying degrees of success. Recently, Fox [21] has shown that a
43 two-step process is required in order to derive a Reynolds-Averaged multiphase tur-
44 bulence model. In the aforementioned models, the multiphase models are derived
45 analogously to a single-phase model using time- or volume-averaging techniques that
46 can lead to ill-formed equations e.g. time averaging results in a diffusive term in the
47 continuity equation.

48 In addition to this, a conceptual error has been highlighted. The statement
49 $k_p = 3\langle\Theta\rangle_p$ is often found in these models which is inaccurate. This is due to the
50 particle turbulent kinetic energy k_p , and the phase averaged (PA) granular tem-
51 perature $\langle\Theta\rangle_p$, belonging to two different realisations of the flow. This distinction
52 was first highlighted by Février et al. [20] in which particle velocities are decomposed
53 into correlated k_p large-scale motion and uncorrelated $\langle\Theta\rangle_p$ small-scale motion. Both
54 quantities are a result of separate models. It was shown that the correlated motion
55 k_p arises due to the hydrodynamic (macroscale) model and the uncorrelated motion
56 $\langle\Theta\rangle_p$ arises due to the kinetic (mesoscale) model.

57 The two-step derivation of Fox [21] has been shown to circumvent these issues.
58 Beginning at the kinetic (mesoscale) equation [22], phase-space integration is applied
59 to find the hydrodynamic (macroscale) moment equation which is then Reynolds-
60 Averaged to form the Reynolds-Averaged Two-Fluid model (RA-TFM) after the ap-
61 propriate closure modelling has been applied. This approach results in separate
62 transport equations for the particle turbulence kinetic energy k_p and the PA gran-
63 ular temperature $\langle\Theta\rangle_p$. Through the derivation of k_p the particle turbulent kinetic
64 energy dissipation ε_p , is defined which appears as a source term in the transport
65 equation of the PA granular temperature, $\langle\Theta\rangle_p$. This cascade of energy from corre-
66 lated motion to uncorrelated motion is a crucial distinction. This leads to the particle
67 fluctuation energy being written as $\kappa_p = k_p + 1.5\langle\Theta\rangle_p$. Février et al. [20] found that
68 even for non-collisional flow, separate transport equations for k_p and $\langle\Theta\rangle_p$ were es-
69 sential, a direct result of the energy cascade outlined previously. Given these recent
70 advances in the field, the modelling of previously challenging turbulent fluid-particle
71 interactions in the Eulerian-Eulerian framework become clearer and their successful
72 prediction more likely.

73 The overarching motivation of the present work is to increase the current under-
74 standing of the modelling of turbulent fluid-particle interaction in a complex flow
75 field. We confine ourselves to turbulence attenuation of small heavy particles in
76 a vertically orientated backward-facing step. The particles have material densities

77 much larger than the fluid ($\rho_p \gg \rho_f$) and diameters smaller than the Kolmogorov
78 length scale over the moderately dilute range $\mathcal{O}(10^{-4})$.

79 The RA-TFM of Fox [21] is implemented in OpenFOAM and applied to the
80 aforementioned flow configuration. The model predictions are compared against
81 benchmark experimental data of Fessler and Eaton [18, 19]. In addition, predictions
82 are also compared against the model of Peirano and Leckner [40] to highlight the
83 importance of separating correlated and uncorrelated motion. Analysis is carried
84 out on the mean particle stream-wise velocities and the fluctuation intensity of both
85 the particle and fluid phases. Applying the RA-TFM to wall-bounded flow requires
86 physical wall boundary conditions for the particle turbulent quantities, k_p , ε_p and
87 $\langle \Theta \rangle_p$. To this end the Johnson and Jackson [31] boundary conditions were recently
88 extended for the RA-TFM by Capecelatro et al. [8] and are implemented and applied
89 here to describe the particle-wall interaction.

90 The paper is organised as follows; in the following section the numerical model
91 is presented. This contains the RA-TFM in which the governing equations for each
92 phase are presented along with the fully coupled turbulence models. Next, the wall
93 boundary conditions used in this work are presented in their implemented form. Fol-
94 lowing this the numerical implementation of the RA-TFM into OpenFOAM is pro-
95 vided along with a description of the turbulence modelling implementation. Section
96 3 presents the results and discussion with the final section providing the conclusions.

97 **2. Numerical Model**

98 Here we begin by presenting the Reynolds-averaged transport equations from
 99 Fox [21]. The RA transport equations are presented in their conservative form with
 100 closures found in Table 5. For clarity the PA notation has been dropped and the PA
 101 variables along with their definitions can be found in Table 6. The particle phase
 102 continuity equation reads:

$$\frac{\partial \alpha_p \rho_p}{\partial t} + \nabla \cdot \alpha_p \rho_p \mathbf{u}_p = 0 \quad (1)$$

103 Where α_p is the volume fraction of particles, ρ_p is the density of the particles and
 104 \mathbf{u}_p is the particle phase velocity.

105 The momentum balance equation for the RA particulate phase is given as:

$$\begin{aligned} \frac{\partial \alpha_p \rho_p \mathbf{u}_p}{\partial t} + \nabla \cdot \alpha_p \rho_p \mathbf{u}_p \mathbf{u}_p &= \nabla \cdot 2(\mu_p + \mu_{pt}) \bar{\mathbf{S}}_p + \beta \left[(\mathbf{u}_f - \mathbf{u}_p) - \frac{\nu_{ft}}{Sc_{fs} \alpha_p \alpha_f} \nabla \alpha_p \right] \\ &\quad - \nabla \left(p_p + \frac{2}{3} \alpha_p \rho_p k_p \right) - \alpha_p \nabla p_f + \alpha_p \rho_p \left[1 - C_p \alpha_f \left(1 - \frac{\rho_f}{\rho_p} \right) \right] \mathbf{g} \end{aligned} \quad (2)$$

106 Where the first term on the right hand side (RHS) contains the diffusive viscosity
 107 associated with the material viscosity and the particle turbulent viscosity. The
 108 forms for these are analogous to those of a fluid and can be found in Table 5. The
 109 second term is the momentum transfer term and contains both the slip velocity
 110 and a turbulent dispersion term. Through the denominator of the dispersion term
 111 $Sc_{fp} = (k_f/k_p)^{1/2}$ a Stokes number (St) dependency is introduced which accounts for
 112 dispersion for moderate to large St. The form of this equation enforces the correct
 113 behaviour, when there is a small St the particle turbulent kinetic energy $k_p \rightarrow k_f$
 114 thus returning 1; for a large St where k_p is small this represents a large value reducing
 115 the amount of dispersion i.e. the particles remain correlated with the fluid.

116 The third term is the pressure gradient along with the so-called turbulent pressure,
 117 with the fourth term being the covariance of the volume fraction and the fluid-
 118 pressure gradient. This term appears in both momentum equations and is typically
 119 assumed to be negligible. The last term contains the body forces (i.e. gravity)
 120 and the velocity-fluid-pressure-gradient covariance term. This term represents the
 121 correlations between the velocity and pressure gradients which arise from buoyancy.

122 The fluid phase continuity equation reads:

$$\frac{\partial \alpha_f \rho_f}{\partial t} + \nabla \cdot \alpha_f \rho_f \mathbf{u}_f = 0 \quad (3)$$

¹²³ The momentum balance equation for the RA fluid phase is given as:

$$\frac{\partial \alpha_f \rho_f \mathbf{u}_f}{\partial t} + \nabla \cdot \alpha_f \rho_f \mathbf{u}_f \mathbf{u}_f = \nabla \cdot 2(\mu_f + \mu_{ft}) \bar{\mathbf{S}}_f + \beta \left[(\mathbf{u}_p - \mathbf{u}_f) + \frac{\nu_{ft}}{Sc_{fs} \alpha_p \alpha_f} \nabla \alpha_p \right] - \nabla \left(\alpha_f p_f + \frac{2}{3} \alpha_f \rho_f k_f \right) + \alpha_p \nabla p_f + \alpha_f \rho_f \left[1 + C_p \alpha_p \left(\frac{\rho_p}{\rho_f} - 1 \right) \right] \mathbf{g} \quad (4)$$

¹²⁴ The momentum equations are symmetric with opposite pressure gradients, hence the
¹²⁵ corresponding term in the particle momentum equation (2) are defined in the same
¹²⁶ manner but with respect the fluid phase.

¹²⁷ The RA turbulent kinetic energy transport equation for the fluid-phase takes the
¹²⁸ form:

$$\begin{aligned} \frac{\partial \alpha_f \rho_f k_f}{\partial t} + \nabla \cdot \alpha_f \rho_f k_f \mathbf{u}_f &= \nabla \cdot \left(\mu_t + \frac{\mu_{ft}}{\sigma_{fk}} \right) \nabla k_f + \alpha_f \rho_f \Pi_f - \alpha_f \rho_f \varepsilon_f \\ &\quad + 2\beta(k_{fp} - k_f) + \alpha_p \rho_p \Pi_{fp} + \alpha_p \rho_p \Pi_{\rho f} \end{aligned} \quad (5)$$

¹²⁹ The first term on the RHS is the fluid-phase turbulent kinetic energy flux. The
¹³⁰ second term Π_f represents kinetic energy production due to mean shear with the
¹³¹ third term being the turbulent kinetic energy dissipation. The remaining three terms
¹³² are the coupling terms: velocity correlations, mean slip and volume-fraction-velocity
¹³³ correlations, respectively.

¹³⁴ The RA turbulent kinetic energy dissipation transport equation for the fluid-
¹³⁵ phase reads as:

$$\begin{aligned} \frac{\partial \alpha_f \rho_f \varepsilon_f}{\partial t} + \nabla \cdot \alpha_f \rho_f \varepsilon_f \mathbf{u}_f &= \nabla \cdot \left(\mu_t + \frac{\mu_{ft}}{\sigma_{fk}} \right) \nabla \varepsilon_f + \frac{\varepsilon_f}{k_f} \left[C_1 \alpha_f \Pi_f - C_2 \alpha_f \rho_f \varepsilon_f \right] \\ &\quad + 2C_3 \beta(\varepsilon_{fp} - \varepsilon_f) + C_4 \frac{\varepsilon_p}{k_p} \alpha_p \rho_p \Pi_{fp} + C_5 \frac{\varepsilon_p}{k_p} \alpha_p \rho_p \Pi_{\rho f} \end{aligned} \quad (6)$$

¹³⁶ The first term on the RHS is the fluid-phase turbulent kinetic dissipation energy
¹³⁷ flux. The second term Π_f is kinetic energy production due to mean shear with the
¹³⁸ third term is dissipation. The remaining three terms are the coupling terms: velocity
¹³⁹ correlations, mean slip and volume-fraction-velocity correlations, respectively. The
¹⁴⁰ forms of these are as follows:

$$\Pi_f = 2\nu_{ft} \bar{\mathbf{S}}_f : \bar{\mathbf{S}}_f + \frac{2}{3} k_f \nabla \cdot \mathbf{u}_f \quad (7)$$

$$\Pi_{fp} = \left[C_g (\mathbf{u}_p - \mathbf{u}_f) - \frac{\nu_{ft}}{Sc_{fp} \alpha_p \alpha_f} \nabla \alpha_p \right] \cdot \left[\beta (\mathbf{u}_p - \mathbf{u}_f) + \frac{1}{\rho_p} \nabla p_f \right] \quad (8)$$

$$\Pi_{\rho f} = C_\rho \left(1 - \frac{\rho_f}{\rho_p}\right) \left[C_g \alpha_p \alpha_f (\mathbf{u}_p - \mathbf{u}_f) - \frac{\nu_{ft}}{Sc_{fp}} \nabla \alpha_p \right] \cdot \mathbf{g} \quad (9)$$

141 Where Π_f is the production of the turbulent kinetic energy, Π_{fp} is due to mean
 142 slip and $\Pi_{\rho f}$ is due to volume-fraction-velocity correlations. The coupling terms take
 143 the form of $k_{fp} = \beta_k \sqrt{k_f k_p}$ and $\varepsilon_{fp} = \beta_\varepsilon \sqrt{\varepsilon_f \varepsilon_p}$. Where correlation coefficients are
 144 $0 < \beta_k, \beta_\varepsilon \leq 1$. These terms represent the fluid-velocity covariance and their exact
 145 closure is still uncertain, a detailed discussion on this point can be found in [21].
 146 This form is adopted as it shows correct limiting behaviour for large St as well as
 147 diminishing correctly in the absence of the particulate phase.

148 The RA particle-phase turbulent kinetic energy reads as:

$$\begin{aligned} \frac{\partial \alpha_p \rho_p k_p}{\partial t} + \nabla \cdot \alpha_p \rho_p k_p \mathbf{u}_p &= \nabla \cdot \left(\mu_p + \frac{\mu_{pt}}{\sigma_{pk}} \right) \nabla k_p + \alpha_p \rho_p \Pi_p - \alpha_p \rho_p \varepsilon_p \\ &\quad + 2\beta(k_{fp} - k_p) + \alpha_p \rho_p \Pi_{\rho p} \end{aligned} \quad (10)$$

149 The first term on the RHS is the particle-phase turbulent kinetic energy flux. The
 150 second term Π_p is kinetic energy production due to mean shear with the third term
 151 being the particle turbulent kinetic energy dissipation. The remaining two terms
 152 are the coupling terms: velocity correlations, and the combination of the buoyancy
 153 induced and mean slip terms.

154 The RA particle-phase turbulent kinetic energy dissipation transport equation reads:

$$\begin{aligned} \frac{\partial \alpha_p \rho_p \varepsilon_p}{\partial t} + \nabla \cdot \alpha_p \rho_p \varepsilon_p \mathbf{u}_p &= \nabla \cdot \left(\mu_p + \frac{\mu_{pt}}{\sigma_{pk}} \right) \nabla \varepsilon_p + \frac{\varepsilon_p}{k_p} (C_1 \alpha_p \rho_p \Pi_p - C_2 \alpha_p \rho_p \varepsilon_p) \\ &\quad + 2C_3 \beta (\varepsilon_{fp} - \varepsilon_p) + C_5 \frac{\varepsilon_p}{k_p} \alpha_p \rho_p \Pi_{\rho p} \end{aligned} \quad (11)$$

155 The first term on the RHS is the particle phase turbulent kinetic dissipation
 156 energy flux. The second term Π_p is kinetic energy production due to mean shear
 157 with the third term being its dissipation. The remaining two terms are the coupling
 158 terms: velocity correlations, and the combination of the buoyancy induced and mean
 159 slip terms. The second term contains, Π_p which is the production of the turbulent
 160 kinetic energy and is expressed as:

$$\Pi_p = 2\nu_{pt} \bar{\mathbf{S}}_p : \bar{\mathbf{S}}_p + \frac{2}{3} k_p \nabla \cdot \mathbf{u}_p \quad (12)$$

161 It should be noted here that the final term on the RHS is a compressive term
 162 that appears in compressible turbulence modelling and plays a similar role of the bulk
 163 viscosity found in the typical granular temperature formulations in the literature [40],

₁₆₄ [51]. Finally, the buoyancy-induced source term $\Pi_{\rho f}$ is added to the mean slip Π_{fp}
₁₆₅ to be reformulated as Π_{pp} which is read as:

$$\Pi_{pp} = C_p C_p \alpha_f \left(1 - \frac{\rho_f}{\rho_p} \right) (\mathbf{u}_p - \mathbf{u}_f) \cdot \mathbf{g} \quad (13)$$

₁₆₆ The granular temperature equation reads as:

$$\begin{aligned} \frac{3}{2} \left[\frac{\partial \alpha_p \rho_p \Theta_p}{\partial t} + \nabla \cdot \alpha_p \rho_p \Theta_p \mathbf{u}_p \right] &= \nabla \cdot \left(\kappa_\Theta + \frac{3\mu_{pt}}{2Pr_{pt}} \right) \nabla \Theta_p + 2\mu_p \bar{\mathbf{S}}_p : \bar{\mathbf{S}}_p \\ &\quad - p_p \nabla \cdot \mathbf{u}_p + \alpha_p \rho_p \varepsilon_p - 3\beta \Theta_p - \gamma \end{aligned} \quad (14)$$

₁₆₇ The first term on the RHS is the PA granular temperature flux which is made up
₁₆₈ of two contributions, the granular temperature flux and the turbulent granular flux.
₁₆₉ The former is the granular conductivity of which there are various formulations in the
₁₇₀ literature. Here the formulation of Syamlal and O'Brien [48] is used as it correctly
₁₇₁ tends to zero in the dilute limit [51]. The latter term is the turbulent flux and
₁₇₂ includes the particle turbulent viscosity. The second term is a laminar source term
₁₇₃ due to viscous stresses. The third term is a pressure dilation term which accounts
₁₇₄ for compressibility. The fourth term is of particular interest as it represents the
₁₇₅ turbulent particle kinetic energy dissipation which appears here as a positive source
₁₇₆ term. The physical interpretation of this means that as large scale particle turbulent
₁₇₇ kinetic energy is dissipated, small scale granular temperature is produced. The two
₁₇₈ remaining terms represent decrease of granular temperature due to drag and decrease
₁₇₉ of granular temperature due to inelastic collisions.

Table 1: Turbulence model parameters.

C_p	C_g	C_ρ	C_1	C_2	C_3	C_4	C_5	β_k	β_ε	$C_{f\mu}$	$C_{p\mu}$
1	0	1	1.44	1.92	1	1	1	1	1	0.09	0.09

₁₈₀ The full form of the equations have been presented here with no mention of their
₁₈₁ relevance to the flow regime being simulated in this work. As the density ratio is
₁₈₂ high the buoyancy induced terms are negligible. The coefficient $C_g \rightarrow 0$ due to
₁₈₃ the small mass loading used in this work, a more thorough discussion on this topic
₁₈₄ can be found in [6, 7, 21]. The buoyancy terms were indeed found to be negligible
₁₈₅ but are retained here to show their solution treatment. Similarly, the compressible
₁₈₆ turbulence correction terms depend on the particle-phase Mach number, which is
₁₈₇ expected to be large for large St. Given the St values used in this study this is not

188 expected to be the case but have been retained. A thorough discussion on this topic
189 is provided in Section 3.4.1.

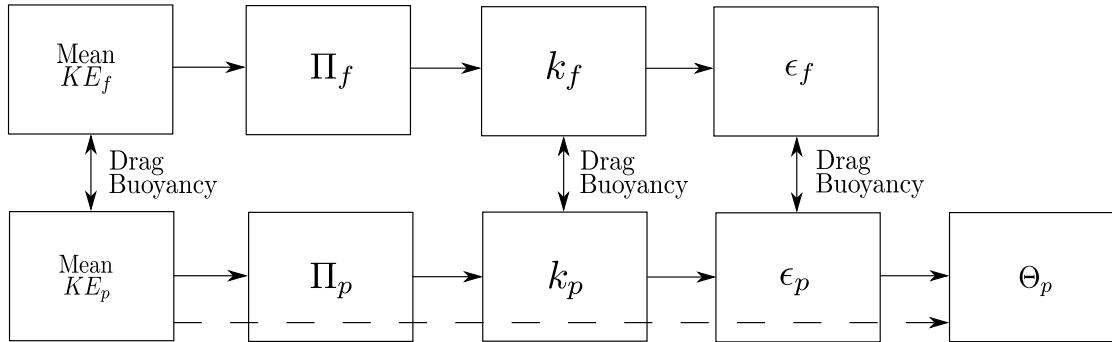


Figure 1: Schematic showing the energy cascade between each variable within the multiphase model [21]. The interaction between each quantity is shown along with their respective energy transfer mechanisms. The dashed line represents the energy flow in the (mesoscale) laminar model.

190 Fig. 1 shows an overall picture of the RA-TFM. As can be seen the energy cascade
191 is almost identical in both phases. Starting with the mean particle kinetic energy
192 $\frac{1}{2}\mathbf{u}_p \cdot \mathbf{u}_p$, the energy transfer to the turbulent particle kinetic energy is given by the
193 production term which is Π_p i.e. the shear. This then generates k_p and is dissipated
194 by the turbulent kinetic energy dissipation equation. Finally, this dissipation term
195 ϵ_p appears in the granular temperature Θ_p as a positive source term, meaning that
196 as the particle turbulent kinetic energy dissipates, granular energy is produced. As
197 can be seen both turbulent quantities interact via drag and buoyancy terms in the
198 same way the governing equations do. If there is dissipation due to collisions then
199 the granular temperature is reduced due to particle heating.

200 2.1. Particle wall boundary conditions

201 The wall boundary conditions for the particle phase require additional modelling.
 202 Recently, Capecelatro et al. [8] started from the Johnson and Jackson [31] wall bound-
 203 ary conditions and derived wall boundary conditions for the particle turbulence quan-
 204 tities k_p , ε_p and Θ_p . Here we present the boundary conditions in their implemented
 205 form applicable for Finite-Volume-Method codes and begin with the wall boundary
 206 condition for the particle velocity \mathbf{u}_p .

$$\mathbf{n} \cdot \bar{\boldsymbol{\sigma}}_p \nabla \mathbf{u}_{p,w} = -\frac{\pi}{6} \frac{\alpha_p}{\alpha_{p,max}} \phi \rho_p g_0 \sqrt{3\Theta_p} \mathbf{u}_{p,w} \quad (15)$$

207 Here we define $\mathbf{u}_{p,w}$ as the particle slip velocity parallel to the wall, $\mathbf{u}_{p,w} = \mathbf{u}_p - \mathbf{u}_w$
 208 with \mathbf{u}_w defined as the wall velocity. Where $\bar{\boldsymbol{\sigma}}_p$ is the particle viscous stress tensor
 209 and we denote \mathbf{n} as the unit vector normal to the wall. Then equation 15 is recast
 210 into a more compact form

$$\mathbf{n} \cdot \bar{\boldsymbol{\sigma}}_p \nabla \mathbf{u}_{p,w} = -\mathcal{D}_w \mathbf{u}_{p,w} \quad (16)$$

211 where the term $\mathcal{D}_w = \phi \mathcal{V}_w$ representing ϕ , the specularity coefficient and the term
 212 $\mathcal{V}_w = \frac{\pi}{6} \frac{\alpha_p}{\alpha_{p,max}} \rho_p \sqrt{3\Theta_p} g_0$ which contains the tangential momentum $\frac{\pi}{6} \frac{\alpha_p}{\alpha_{p,max}} \rho_p$ (omit-
 213 ting the particle slip velocity parallel to the wall $\mathbf{u}_{p,w}$) and the collisional frequency
 214 $\sqrt{3\Theta_p} g_0$. This boundary condition prescribes a particle partial-slip velocity at the
 215 wall. From this condition it follows that the components of the Reynolds stress tensor
 216 $\langle \mathbf{u}_p'' \mathbf{u}_p'' \rangle_p$ need not be zero at the wall unlike in the fluid phase. As we are interested
 217 in modelling the particle-wall interaction of the particle turbulent kinetic energy k_p
 218 we assume isotropy in the fluctuating components.

$$k_p = \frac{1}{2} (\mathbf{u}_{p,x}''^2 + \mathbf{u}_{p,y}''^2 + \mathbf{u}_{p,z}''^2) \quad (17)$$

219 Equating the principal Reynolds stress components ($\mathbf{u}_{p,x}''^2 \equiv \mathbf{u}_{p,y}''^2 \equiv \mathbf{u}_{p,z}''^2$) one
 220 arrives at $k_p = 1.5 \mathbf{u}_p''^2$ and substituting into Equation 16 by employing the PA
 221 decomposition (see Table 6) the wall boundary condition for k_p reads

$$\mathbf{n} \cdot \bar{\boldsymbol{\sigma}}_p \nabla k_p = -2\mathcal{D}_w k_p \quad (18)$$

222 Following on from this a condition for the particle turbulent kinetic energy dissipa-
 223 tion rate ε_p can be prescribed:

$$\mathbf{n} \cdot \bar{\boldsymbol{\sigma}}_p \nabla \varepsilon_p = -2\mathcal{D}_w \varepsilon_p \quad (19)$$

224 Finally, the wall boundary condition for the granular temperature can be found
225 by Reynolds averaging the Johnson and Jackson [31] which reads as

$$\mathbf{n} \cdot \mathbf{q}_\Theta \frac{3}{2} = \phi \mathcal{D}_w |\mathbf{u}_{p,w}|^2 - \frac{3}{2} \mathcal{D}_\kappa \Theta_p \quad (20)$$

226 where \mathbf{q}_Θ is the granular temperature flux and with $\mathcal{D}_\kappa = (1 - e_w)^2 \mathcal{D}_w$ and e_w
227 is the restitution of coefficient with the wall. It follows from this that this term
228 represents the energy loss through particle collisions with the wall. The first term
229 on the RHS represents the increase of the granular temperature due to the relative
230 velocity with the wall. This means that the slip condition at the wall is capable of
231 increasing the granular temperature. The particle-wall coefficient, e_w was set equal
232 to 0.9. The specularity coefficient used throughout was, $\phi = 0.001$ and the influence
233 of this parameter is discussed in Section 3.3.4.

234 2.2. Numerical implementation

235 Here we follow the phase intensive formulation of Rusche [43], Weller [55]. We
 236 rewrite the equations in their non-conservative form by expanding both the convective
 237 term and dividing by density and volume fraction.

238 Additionally, the phase respective Reynolds stress tensor is formulated by grouping
 239 the kinematic and turbulent viscosity into an effective viscosity, $\nu_{\text{eff},i} = \nu_i + \nu_{it}$
 240 and employing the Boussinesq hypothesis. Now, we can write the Reynolds stress
 241 tensor in the form:

$$\bar{\mathbf{R}}_{\text{eff,p}} = -2\nu_{\text{eff,p}}\bar{\mathbf{S}}_p + \frac{2}{3}\nu_{\text{eff,p}}\mathbf{I}\nabla \cdot \mathbf{u}_p + \frac{2}{3}\mathbf{I}k_p \quad (21)$$

$$\bar{\mathbf{R}}_{\text{eff,f}} = -2\nu_{\text{eff,f}}\bar{\mathbf{S}}_f + \frac{2}{3}\nu_{\text{eff,f}}\mathbf{I}\nabla \cdot \mathbf{u}_f + \frac{2}{3}\mathbf{I}k_f \quad (22)$$

242 For simplicity the turbulent dispersion term is now denoted as \mathcal{D} , separating the
 243 drag contributions into explicit and implicit terms and dividing by both the phase
 244 fraction and density we are left with:

$$\begin{aligned} \frac{\partial \mathbf{u}_p}{\partial t} + \nabla \cdot (\mathbf{u}_p \mathbf{u}_p) - \mathbf{u}_p \nabla \cdot \mathbf{u}_p + \frac{\nabla \alpha_p}{\alpha_p + \delta} \cdot \bar{\mathbf{R}}_{\text{eff,p}}^c + \nabla \cdot \bar{\mathbf{R}}_{\text{eff,p}}^c - \nabla \cdot (\nu_{\text{eff,p}} \nabla \mathbf{u}_p) + \frac{\beta \mathbf{u}_p}{\alpha_p \rho_p} \\ - \nabla \cdot \left(\nu_{\text{eff,p}} \frac{\nabla \alpha_p}{\alpha_p + \delta} \mathbf{u}_p \right) + \mathbf{u}_p \nabla \cdot \left(\nu_{\text{eff,p}} \frac{\nabla \alpha_p}{\alpha_p + \delta} \right) \\ = \frac{\beta \mathbf{u}_f}{\alpha_p \rho_p} - \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_p \rho_p} - \frac{\nabla p_p}{\alpha_p \rho_p} - \frac{\nabla p_f}{\rho_p} + \mathbf{g} - \alpha_f \left(1 - \frac{\rho_f}{\rho_p} \right) \mathbf{g} \end{aligned} \quad (23)$$

$$\begin{aligned} \frac{\partial \mathbf{u}_f}{\partial t} + \nabla \cdot (\mathbf{u}_f \mathbf{u}_f) - \mathbf{u}_f \nabla \cdot \mathbf{u}_f + \frac{\nabla \alpha_f}{\alpha_f + \delta} \cdot \bar{\mathbf{R}}_{\text{eff,f}}^c + \nabla \cdot \bar{\mathbf{R}}_{\text{eff,f}}^c - \nabla \cdot (\nu_{\text{eff,f}} \nabla \mathbf{u}_f) + \frac{\beta \mathbf{u}_f}{\alpha_f \rho_f} \\ - \nabla \cdot \left(\nu_{\text{eff,f}} \frac{\nabla \alpha_f}{\alpha_f + \delta} \mathbf{u}_f \right) + \mathbf{u}_p \nabla \cdot \left(\nu_{\text{eff,f}} \frac{\nabla \alpha_f}{\alpha_f + \delta} \right) \\ = \frac{\beta \mathbf{u}_p}{\alpha_f \rho_f} + \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_f \rho_f} - \frac{\nabla p_f}{\rho_f} - \frac{\alpha_p \nabla p_f}{\alpha_f \rho_f} + \mathbf{g} + \alpha_p \left(\frac{\rho_p}{\rho_f} - 1 \right) \mathbf{g} \end{aligned} \quad (24)$$

245 where δ is introduced to avoid a division by zero. As it can be seen from the system
 246 of equations in Eqs. 2 & 4 no diffusive flux exists that can be treated implicitly. This
 247 can have advantages when solving the equations i.e enhanced matrix positivity and
 248 diagonal dominance. Therefore, following Weller [55], Rusche [43] the Reynolds stress
 249 term can be rewritten into a diffusive and corrective component:

$$\begin{aligned}
\bar{\mathbf{R}}_{\text{eff},i} &= \bar{\mathbf{R}}_{\text{eff},i} + \nu_{\text{eff},i} \nabla \mathbf{u}_i - \nu_{\text{eff},i} \nabla \mathbf{u}_i \\
&= -\nu_{\text{eff},i} (\nabla \mathbf{u}_i + \nabla^T \mathbf{u}_i) + \frac{2}{3} \nu_{\text{eff},i} \mathbf{I} \nabla \cdot \mathbf{u}_i \\
&\quad + \frac{2}{3} \mathbf{I} k_i + \nu_{\text{eff},i} \nabla \mathbf{u}_i - \nu_{\text{eff},i} \nabla \mathbf{u}_i \\
&= (-\nu_{\text{eff},i} \nabla^T \mathbf{u}_i + \frac{2}{3} \nu_{\text{eff},i} \mathbf{I} \nabla \cdot \mathbf{u}_i + \frac{2}{3} k_i \mathbf{I}) - \nu_{\text{eff},i} \nabla \mathbf{u}_i \\
&= \bar{\mathbf{R}}_{\text{eff},i}^c - \nu_{\text{eff},i} \nabla \mathbf{u}_i
\end{aligned} \tag{25}$$

250 It is important to clarify the behaviour of terms with the volume fraction in their
251 denominator. For the first contribution due to drag i.e. the third term on the RHS
252 the coefficient β contains $\alpha_p \alpha_f$ which ensures the correct behavior of the function
253 as $\alpha_p \rightarrow 0$. The second term containing turbulent dispersion contains the gradient
254 of volume fraction which in the limit $\alpha_p \rightarrow 0$ means that the ratio approaches zero.
255 Finally, the fluid velocity-pressure-covariance term contains the volume fraction in
256 its denominator, however due to the particle packing limit ensured by the particle's
257 structural properties the volume fraction of the fluid phase does not approach 0.

258 *2.3. Discretisation of the intensive momentum equations*

259 First, we discretise the left hand side of the equation which contains the convective
260 and diffusive transport terms.

$$\begin{aligned}
\mathcal{T}_p := & \left[\frac{\partial \mathbf{u}_p}{\partial t} \right] + \left[\nabla \cdot (\mathbf{u}_p [\mathbf{u}_p]) \right] - \left[(\nabla \cdot \mathbf{u}_p) [\mathbf{u}_p] \right] + \frac{\nabla \alpha_p}{\alpha_p + \delta} \cdot \bar{\mathbf{R}}_{\text{eff},p}^c + \nabla \cdot \bar{\mathbf{R}}_{\text{eff},p}^c \\
& - \left[\nabla \cdot (\nu_{\text{eff},p} \nabla [\mathbf{u}_p]) \right] - \left[\nabla \cdot (\nu_{\text{eff},p} \frac{\nabla \alpha_p}{\alpha_p + \delta} [\mathbf{u}_p]) \right] \\
& - \left[\nabla \cdot (\nu_{\text{eff},p} \frac{\nabla \alpha_p}{\alpha_p + \delta}) [\mathbf{u}_p] \right] + \left[\frac{\beta \mathbf{u}_p}{\alpha_p \rho_p} \right]
\end{aligned} \tag{26}$$

$$\begin{aligned}
\mathcal{T}_f := & \left[\frac{\partial \mathbf{u}_f}{\partial t} \right] + \left[\nabla \cdot (\mathbf{u}_f [\mathbf{u}_f]) \right] - \left[(\nabla \cdot \mathbf{u}_f) [\mathbf{u}_f] \right] + \frac{\nabla \alpha_f}{\alpha_f + \delta} \cdot \bar{\mathbf{R}}_{\text{eff},f}^c + \nabla \cdot \bar{\mathbf{R}}_{\text{eff},f}^c \\
& - \left[\nabla \cdot (\nu_{\text{eff},f} \nabla [\mathbf{u}_f]) \right] - \left[\nabla \cdot (\nu_{\text{eff},f} \frac{\nabla \alpha_f}{\alpha_f + \delta} [\mathbf{u}_f]) \right] \\
& - \left[\nabla \cdot (\nu_{\text{eff},f} \frac{\nabla \alpha_f}{\alpha_f + \delta}) [\mathbf{u}_f] \right] + \left[\frac{\beta \mathbf{u}_f}{\alpha_f \rho_f} \right]
\end{aligned} \tag{27}$$

261 where $[\cdot]$ is the implicit discretisation of the term, \mathcal{T}_p & \mathcal{T}_f represents the numerical
262 coefficients of each respective algebraic system given by the discretisation. The

263 second and third terms on the RHS represent convection and have been split up
 264 into a convection term minus a divergence terms as it enhances boundedness of the
 265 solution.

266 \mathcal{T}_p & \mathcal{T}_f represents the system of algebraic equations from the discretised Eqs.
 267 26 & 27 which appear in the form,

$$(\mathcal{T}_p)_{coefs} \mathbf{u}_p = (\mathcal{T}_p)_s \quad (28a)$$

$$(\mathcal{T}_f)_{coefs} \mathbf{u}_f = (\mathcal{T}_f)_s \quad (28b)$$

268 This discretised form of the momentum equations will be revisited once the source
 269 terms on the RHS are addressed.

270 Now addressing the RHS of Eq. 23 & 24 which reads as

$$\dots = \frac{\beta \mathbf{u}_f}{\alpha_p \rho_p} - \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_p \rho_p} - \frac{\nabla p_p}{\alpha_p \rho_p} - \frac{\nabla p_f}{\rho_p} + \mathbf{g} - \alpha_f \left(1 - \frac{\rho_f}{\rho_p}\right) \mathbf{g} \quad (29a)$$

$$\dots = \frac{\beta \mathbf{u}_p}{\alpha_f \rho_f} + \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_f \rho_f} - \frac{\nabla p_f}{\rho_f} - \frac{\alpha_p \nabla p_f}{\alpha_f \rho_f} + \mathbf{g} + \alpha_p \left(\frac{\rho_p}{\rho_f} - 1\right) \mathbf{g} \quad (29b)$$

271 denoting the buoyancy terms to be $\mathbf{g}_p^* = \mathbf{g}(1 - \alpha_p(1 - \frac{\rho_f}{\rho_p}))$ & $\mathbf{g}_f^* = \mathbf{g}(1 + \alpha_p(\frac{\rho_p}{\rho_f} - 1))$
 272 we can write

$$\dots = \frac{\beta \mathbf{u}_f}{\alpha_p \rho_p} - \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_p \rho_p} - \frac{\nabla p_f}{\rho_p} - \frac{\nabla p_p}{\alpha_p \rho_p} + \mathbf{g}_p^* \quad (30a)$$

$$\dots = \frac{\beta \mathbf{u}_p}{\alpha_f \rho_f} + \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_f \rho_f} - \frac{\nabla p_f}{\rho_f} - \frac{\alpha_p \nabla p_f}{\alpha_f \rho_f} + \mathbf{g}_f^* \quad (30b)$$

273 Following the solution procedure of Weller [55] all terms on the RHS are evaluated
 274 at cell faces. In order to avoid checker-boarding in the solution, which is a prevalent
 275 problem on collocated grids due to the storage of values at cell centres and interpo-
 276 lating onto the face, the group of terms on the RHS are treated in a Rhie-Chow like
 277 manner Rhie and Chow [42].

278 2.4. Phase momentum flux correction equations

279 Now a semi-discretised formulation of both the particle- and fluid-phase can be
 280 written. Invoking Eqs. 28 and splitting up the total coefficients appearing in each
 281 system into diagonal and explicit \mathbf{H} [30] coefficients. The latter consisting of two
 282 parts, the neighbouring coefficients (multiplied by its respective phase velocity) and
 283 the source terms, $\mathbf{H}_i = -(\mathbf{A}_i)_N \mathbf{u}_i + (\mathbf{A}_i)_S$. The equations can then be written as:

$$\mathbf{A}_p \mathbf{u}_p = \mathbf{H}_p + \frac{\beta \mathbf{u}_f}{\alpha_p \rho_p} - \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_p \rho_p} - \frac{\nabla p_f}{\rho_p} - \phi_{p,p} + \mathbf{g}_p^* \quad (31a)$$

$$\mathbf{A}_f \mathbf{u}_f = \mathbf{H}_f + \frac{\beta \mathbf{u}_p}{\alpha_f \rho_f} + \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_f \rho_f} - \frac{\nabla p_f}{\rho_f} - \phi_{f,g} + \mathbf{g}_f^* \quad (31b)$$

284 Rearranging Eqs. 31 gives the phase momentum correction equations, note these
 285 equations are not used in the solution algorithm, but are required to derive a flux
 286 predictor and corrector:

$$\mathbf{u}_p = \frac{\mathbf{H}_p}{\mathbf{A}_p} + \frac{\beta \mathbf{u}_f}{\alpha_p \rho_p \mathbf{A}_p} - \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_p \rho_p \mathbf{A}_p} - \frac{\nabla p_f}{\rho_p \mathbf{A}_p} - \frac{\phi_{p,p}}{\mathbf{A}_p} + \frac{\mathbf{g}_p^*}{\mathbf{A}_p} \quad (32a)$$

$$\mathbf{u}_f = \frac{\mathbf{H}_f}{\mathbf{A}_f} + \frac{\beta \mathbf{u}_p}{\alpha_f \rho_f \mathbf{A}_f} + \frac{\beta \mathcal{D} \nabla \alpha_p}{\alpha_f \rho_f \mathbf{A}_f} - \frac{\nabla p_f}{\rho_f \mathbf{A}_f} - \frac{\phi_{f,g}}{\mathbf{A}_f} + \frac{\mathbf{g}_f^*}{\mathbf{A}_f} \quad (32b)$$

287 *2.5. Construction of the pressure equation*

288 In order to derive a pressure equation the continuity equation is enforced globally.
 289 The global continuity equation thus reads:

$$\nabla \cdot [(\alpha_p)_f \phi_p + (\alpha_f)_f \phi_f] = 0 \quad (33)$$

290 where the subscript $(\cdot)_f$ denotes the face value and $\phi_i = (\mathbf{u}_i)_f \cdot \mathbf{S}_f$ is the volumetric
 291 face flux i.e. the sum of all the fluxes over a control volume. From here the phase
 292 fluxes for each phase are found by interpolating the momentum correction equation
 293 (Eqs. 32) onto face centres. Using central differencing, we can write

$$\phi_p = \phi_p^* - \left(\frac{1}{\rho_p \mathbf{A}_p} \right)_f \nabla_f^\perp p_f \cdot \mathbf{S}_f \quad (34a)$$

$$\phi_f = \phi_f^* - \left(\frac{1}{\rho_f \mathbf{A}_f} \right)_f \nabla_f^\perp p_f \cdot \mathbf{S}_f \quad (34b)$$

294 where the flux predictions ϕ_p^* & ϕ_f^* are given by

$$\begin{aligned} \phi_p^* = & \left(\frac{\mathbf{H}_p}{\mathbf{A}_p} \right)_f \cdot \mathbf{S}_f + \left(\frac{\beta}{\alpha_p \rho_p \mathbf{A}_p} \right)_f \phi_f - \left(\frac{\beta \mathcal{D}}{\alpha_p \rho_p \mathbf{A}_p} \right)_f \nabla_f^\perp \alpha_p \cdot \mathbf{S}_f \\ & - \left(\frac{1}{\mathbf{A}_p} \right)_f \nabla_f^\perp p_p \cdot \mathbf{S}_f + \left(\frac{1}{\mathbf{A}_p} \right)_f \mathbf{g}_p^* \cdot \mathbf{S}_f \end{aligned} \quad (35)$$

$$\begin{aligned}\phi_f^* = & \left(\frac{\mathbf{H}_f}{\mathbf{A}_f} \right)_f \cdot \mathbf{S}_f + \left(\frac{\beta}{\alpha_f \rho_f \mathbf{A}_f} \right)_f \phi_p - \left(\frac{\beta \mathcal{D}}{\alpha_f \rho_f \mathbf{A}_f} \right)_f \nabla_f^\perp \alpha_p \cdot \mathbf{S}_f \\ & - \left(\frac{\alpha_p}{\alpha_f \rho_f \mathbf{A}_f} \right)_f \nabla_f^\perp p_f \cdot \mathbf{S}_f + \left(\frac{1}{\mathbf{A}_f} \right)_f \mathbf{g}_f^* \cdot \mathbf{S}_f\end{aligned}\quad (36)$$

Now the pressure equation can be constructed by substituting Eq. 34 into Eq. 33 which reads:

$$\left[\nabla \cdot \left(D_p \nabla_f^\perp [p_f] \cdot \mathbf{S}_f \right) \right] = \nabla \cdot \left((\alpha_p)_f \phi_p^* + (\alpha_f)_f \phi_f^* \right) \quad (37)$$

where

$$D_p = \left(\frac{\alpha_p}{\rho_p \mathbf{A}_p} + \frac{\alpha_f}{\rho_f \mathbf{A}_f} \right)_f, \quad (38)$$

and the pressure gradient has been discretised implicitly on the LHS as a diffusion term i.e. Laplacian. Essentially a shared or mixture pressure field is solved for, this ensures that continuity is obeyed throughout as the coupling is provided through the pressure equation. Once this equation has been solved the phase fluxes need to be updated to satisfy continuity, as in the predictor step the influence of the pressure gradient is removed, this can be achieved by invoking Eq. 34. From this stage the solution does not completely satisfy continuity as the velocities, which are stored at the cell centres, need to be updated with the influence of the pressure gradient.

This is achieved by invoking:

$$\mathbf{u}_p = \frac{\mathbf{H}_p}{\mathbf{A}_p} + \left[\phi_p^* - \left(\frac{1}{\rho_p \mathbf{A}_p} \right)_f \nabla_f^\perp p_f \cdot \mathbf{S}_f \right]_{f \rightarrow c} \quad (39a)$$

$$\mathbf{u}_f = \frac{\mathbf{H}_f}{\mathbf{A}_f} + \left[\phi_f^* - \left(\frac{1}{\rho_f \mathbf{A}_f} \right)_f \nabla_f^\perp p_f \cdot \mathbf{S}_f \right]_{f \rightarrow c} \quad (39b)$$

where the subscript $f \rightarrow c$ denotes a vector field reconstruction from face flux values to cell centre values. The influence of the gradient of pressure is incorporated into the reconstruction of the phase velocity - this ensures the phase velocity obeys continuity. Once this is completed the PISO loop is complete.

2.6. Solution of the phase-mixed continuity equation

In practice the phase-mixed continuity equation is solved first based on the initial conditions but for the sake of logical progression is presented now. Following Weller

³¹³ [55] the particle phase continuity equation ?? can be reformulated as:

$$\frac{\partial \alpha_p}{\partial t} + \nabla \cdot (\mathbf{u}_T \alpha_p) + \nabla \cdot (\mathbf{u}_r \alpha_p \alpha_f) = 0 \quad (40)$$

³¹⁴ where $\mathbf{u}_T = \alpha_p \mathbf{u}_p + \alpha_f \mathbf{u}_f$ is the mixture velocity and $\mathbf{u}_r = \mathbf{u}_p - \mathbf{u}_f$ is the relative
³¹⁵ velocity. This equation can then be discretised as

$$\left[\frac{\partial [\alpha_p]}{\partial t} \right] + \left[\nabla \cdot (\phi [\alpha_p]_f) \right] + \left[\nabla \cdot (\phi_{r,p} [\alpha_p]_f) \right] = 0 \quad (41)$$

³¹⁶ where $\phi_{r,p} = (\alpha_f)_f \phi_r$ and $\phi_r = \phi_p - \phi_f$.

³¹⁷ The second term on the LHS is ensured to be bounded between 0 and 1 due to the
³¹⁸ mixture flux, $\phi = (\mathbf{u}_p)_f \cdot \mathbf{S}_f + (\mathbf{u}_f)_f \cdot \mathbf{S}_f$ satisfying the mixture continuity equation.
³¹⁹ The third term is bounded by employing ϕ_r in the convective scheme by interpolating
³²⁰ α_p to the face and $-\phi_r$ in the face interpolation of α_f . This system is solved using
³²¹ the the Multi-dimensional Universal Limiter with Explicit Solution (MULES) [59]
³²² which is a flux-corrected transport algorithm which enhances robustness, stability
³²³ and convergence.

The numerical procedure adopted in the segregated algorithm:

1. Solve the volume fraction (Eq. 41).
2. Construct \mathbf{A}_i in each phase (Eqs. 28).
3. Enter PISO-Loop:
 - (a) Predict fluxes using Eqs. 35 & 36.
 - (b) Construct and solve the pressure equation (Eq. 37).
 - (c) Correct the phase fluxes using Eqs. 34.
 - (d) Reconstruct the phase velocities using Eqs. 39.
4. Solve the system of phase energy equations

325 The $k_p - \varepsilon_p$ transport equation is implemented in a similar manner to the $k_f - \varepsilon_f$
326 transport equation. A thorough discussion of its implementation in OpenFOAM can
327 be found in [37, 30, 43]. The coupling terms found in each equation are handled in a
328 segregated manner due to the velocity-pressure solution treatment. From the solution
329 of the aforementioned transport equations the granular temperature is then solved
330 as the particle turbulence dissipation ε_p appears as a source term in the transport
331 equation of granular temperature Θ_p .

332 *2.7. Numerical solution*

333 The open-source toolbox OpenFOAM [56] is used to solve the RA-TFM equa-
334 tions. To handle the pressure-velocity coupling the Pressure Implicit with Splitting
335 Operators (PISO) algorithm [17, 29] is used. The volume fraction is solved using a
336 Multi-dimensional Universal Limiter with Explicit Solution (MULES) [59] which is
337 a flux-corrected transport algorithm which ensures robustness, stability and conver-
338 gence.

339 2.8. Simulation cases

340 The computational domain is a two-dimensional channel section as seen in Fig.
 341 2 which starts at $60h$ upstream of the step to allow the flow to fully develop and
 342 extends $30H$ downstream. The material constants for each respective case can be
 343 found in Table 2. As reported in the experiments the centreline velocity, U_0 is 10.5
 344 ms^{-1} at the step ($x/H = 0$) and this corresponded to an inlet value of 9.3 ms^{-1} . Mass
 345 loading is given by a uniform particle volume fraction across the inlet, this is achieved
 346 by assuming a constant particle-to-fluid velocity ratio. Wall functions for the fluid
 347 phase are used throughout and the effect of the particles on the boundary layer is
 348 not considered here. For the particle phase the boundary conditions described in
 349 Section 2.1 are used for the turbulent quantities. At the inlet a first estimate of the
 350 two turbulent quantities is determined as follows; $k_p = 1/3k_f$ and $\varepsilon_p = 1/3\varepsilon_f$ (for
 351 a more elaborate approach see [21]). For the granular temperature a small value is
 352 specified; $\Theta_p = 1.0 \times 10^{-8} \text{ m}^2 \text{s}^{-2}$ [51]. Calculations are carried out on a fully structured
 353 hexahedral mesh consisting of 11,253 cells with $y^+ > 30$ along both walls. Refinement
 354 was introduced around the step resulting in the mesh cells sizes of 0.5 mm in the x
 355 and y direction, respectively.

Table 2: Table of simulated cases.

Case	Material	$d_p [\mu\text{m}]$	$\rho_p [\text{kg m}^{-3}]$	Mass loading	St	Re_p
1	glass	150	2500	20% and 40%	7.9	10.1
2	glass	90	2500	20%	3.8	2.9
3	copper	70	8800	10% and 40%	7.4	4.4

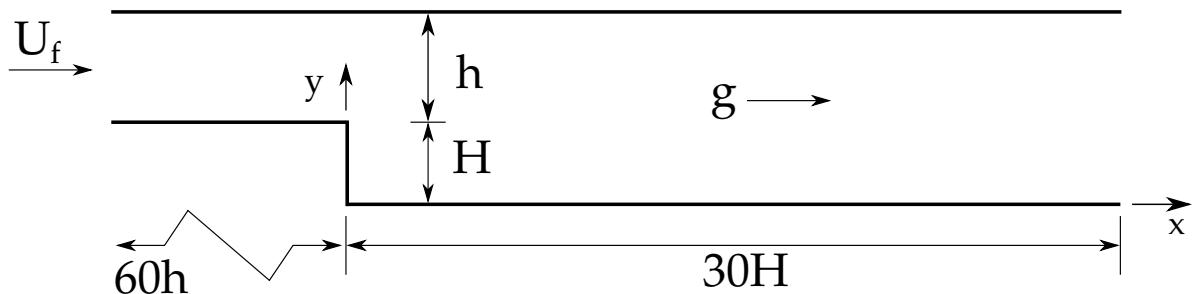


Figure 2: Schematic detailing the geometry used throughout, with $h = 0.02\text{m}$ and $H = 0.0267\text{m}$.

356 **3. Results and Discussion**

357 The simulated results from the RA-TFM and the modified Peirano model (MPM)
 358 are compared against two sets of experimental data given by Fessler and Eaton [19]
 359 & Fessler and Eaton [18]. The MPM equations can be found in the Appendix. Mean
 360 quantities of particle velocity, fluid turbulence intensity and particle turbulence in-
 361 tensity are presented across three cases focusing on three particle classes (see Table
 362 1). The measured velocity profiles start at the recirculation region ($x/H = 2$), con-
 363 tinue through to the reattachment zone ($x/H = 5$), and finally the redevelopment
 364 region ($x/H = 14$) with measurements being taken in between.

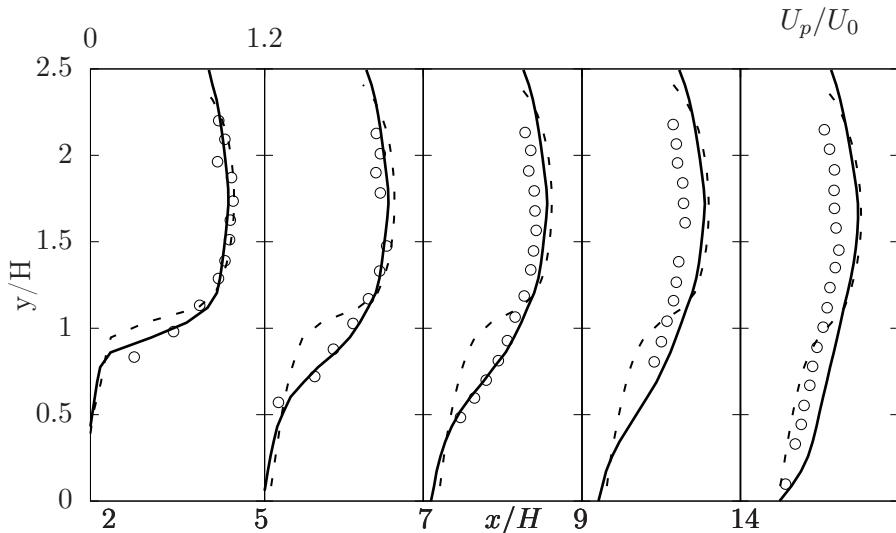


Figure 3: Stream-wise particle mean velocity for case 1. Solid line showing the RA-TFM and the dashed line showing the MPM. Data from Fessler and Eaton [18] with a mass-loading of 40%.

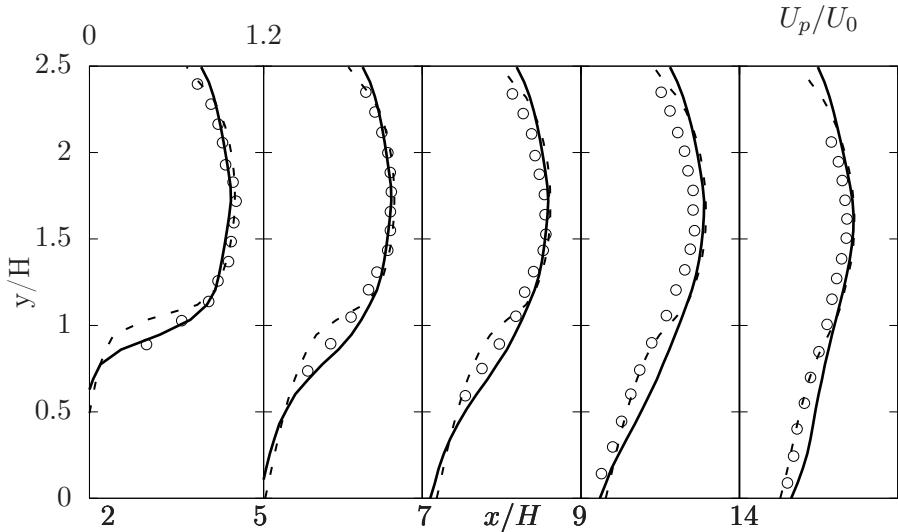


Figure 4: Stream-wise particle mean velocity for case 2. Solid line showing the RA-TFM and the dashed line showing the MPM. Data from Fessler and Eaton [18] with a mass-loading of 20%.

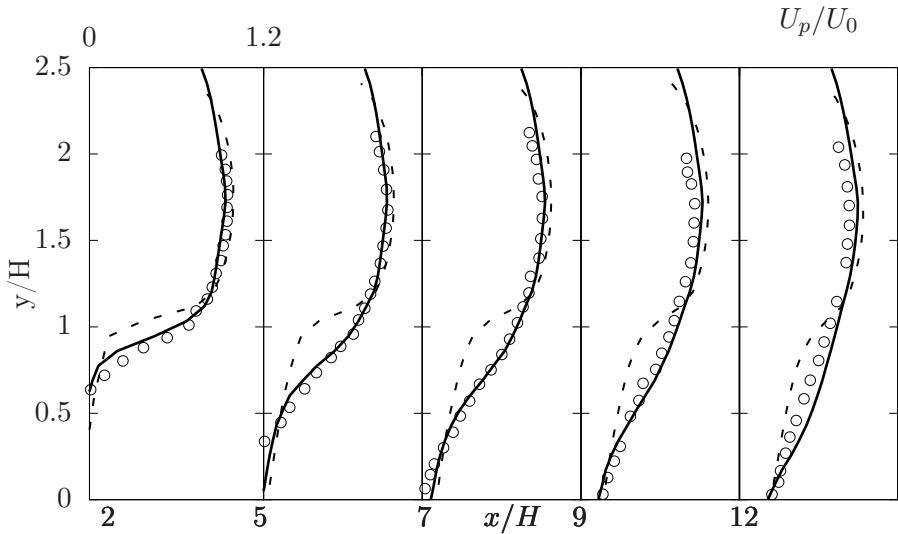


Figure 5: Stream-wise particle mean velocity for case 3. Solid line showing the RA-TFM and the dashed line showing the MPM. Data from Fessler and Eaton [19] with a mass-loading of 10%.

365 *3.1. Mean particle stream-wise velocity*

366 It is evident that across all three Figs. 3, 4 and 5 the prediction of the RA-TFM
367 mean particle velocity profiles are in good agreement with the measured results. The

368 model captures the affects of varying St and mass loading on the mean velocity pro-
369 file, especially in the recirculation region corresponding to locations ($x/H = 2$) and
370 ($x/H = 5$). The MPM shows a marked difference around the step ($0.5 < y/H \leq 1$)
371 as it does not include the correlated particle turbulent kinetic fluctuations. These
372 stresses are responsible for forming the shear layer and leading to the production of
373 turbulent viscosity.

374 All three Figs. 3, 4 and 5 show particle velocities with a flatter gradient across
375 the depth of the pipe when compared to the fluid velocities, a feature that is not
376 predicted by the MPM. This is attributed to the calculation of the particle phase
377 viscosity. In the RA-TFM the calculation of the turbulent viscosity μ_{pt} appearing in
378 the momentum equation is given by the $k_p - \varepsilon_p$ turbulence model, which accounts
379 for the correlated turbulent kinetic particle fluctuations that are dominant due to
380 the shear layer. In the MPM the viscosity is calculated directly from the granular
381 temperature equation which relies on constitutive closures of thermal conductivity,
382 shear viscosity and bulk viscosity [40]. As a result, a small value of both is predicted
383 due to the dilute nature of the flow and this leads to a gross under-prediction of the
384 particle viscosity. This then allows the momentum coupling term β to dominate in
385 this region, which is why the mean velocity profiles tend to closely follow the fluid
386 phase mean velocity profile.

387 Figs. 3 & 5 reveal the largest variation between the predicted mean particle ve-
388 locity profiles in the shear layer. This is attributed to the particles St, which varies
389 considerably over the shear layer as shown in both Figs ($y/H < 1$). When the par-
390 ticles $St \gg 1$ the particles tend to escape from the eddy they are in and ignore the
391 influence of external eddies. This can either unite small eddies to create larger more
392 energetic eddies or it can destroy large eddies which dissipate to smaller eddies. As
393 a consequence of this for $St \gg 1$ we can expect the particle to take longer to react
394 to the fluid. Therefore, when considering the shear layer the fluid response time, τ_f
395 will be small in comparison with the channel flow resulting in a much higher local
396 St. As a result the particle mean velocity profile does not show the sharp gradient
397 across the ($y/H > 1$) and becomes much flatter.

398 Fig. 3 shows the case denoting both a high St (7.9) and a large mass loading
399 (40%). It also corresponds to the largest over-prediction in the mean particle ve-
400 locities at locations ($x/H = 9$ & 14) for the RA-TFM. These locations correspond
401 to the redevelopment region which indicate that the energy in the particle phase is
402 recovering too quickly in comparison to the measured data. This overestimation is
403 difficult to explain as the predictions for case 3 with a large St are in good agree-
404 ment. One potential source of error could be due to the distribution of the particles
405 across the width of the pipe. As the particles pass the step they are redistributed

406 inhomogeneously (clustered) which reduces the slip velocity and as a result the drag.
 407 As the particles reach the redevelopment region they begin to redistribute homoge-
 408 neously which increases the drag in this region. However neither model considers the
 409 effects of clustering in their drag model and are only representative of one particle.
 410 This can cause the observed over-estimation of the mean stream-wise velocities in
 411 the redevelopment region.

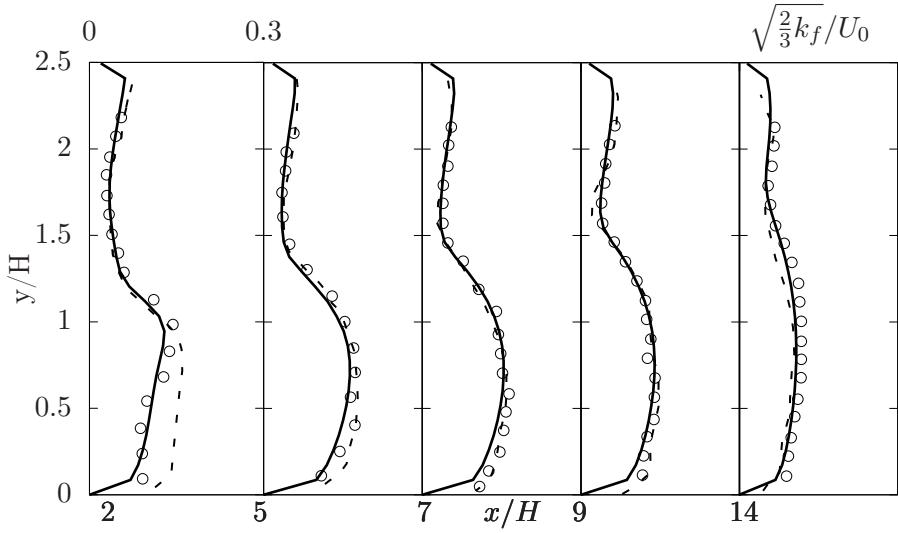


Figure 6: Fluid turbulent intensity for case 1. Solid line showing the RA-TFM and the dashed line showing the MPM. Data from Fessler and Eaton [19] with a mass-loading of 40%.

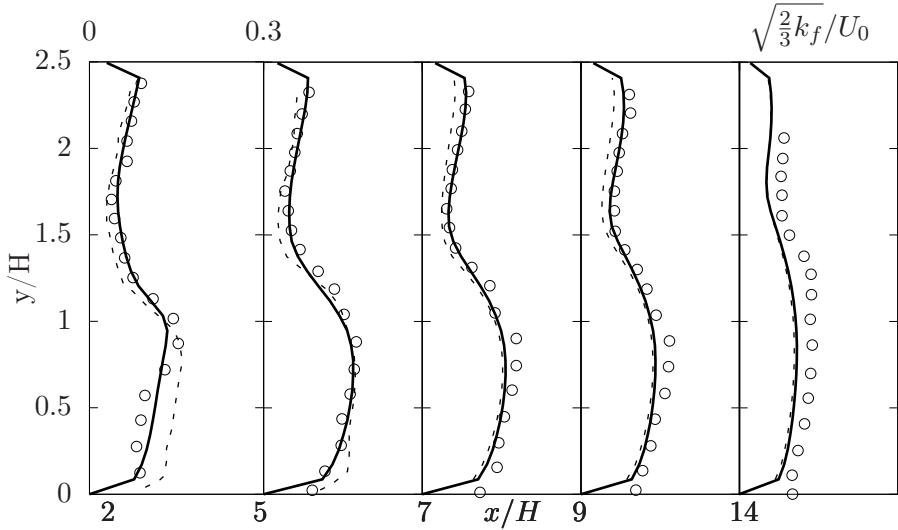


Figure 7: Fluid turbulent intensity for case 2. Solid line showing the RA-TFM and the dashed line showing the MPM. Data from Fessler and Eaton [19] with a mass-loading of 20%.

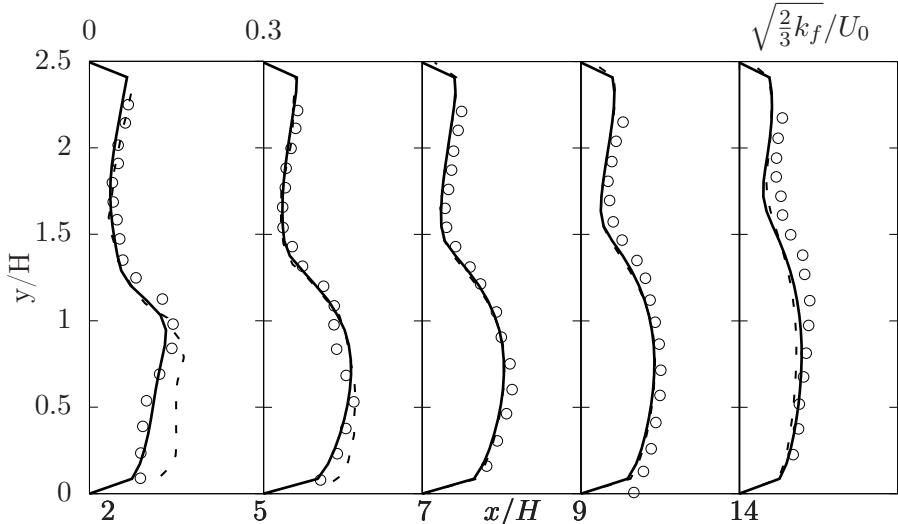


Figure 8: Fluid turbulent intensity for case 3. Solid line showing the RA-TFM and the dashed line showing the MPM. Data from Fessler and Eaton [19] with a mass-loading of 40%.

412 3.2. Fluid phase turbulence

413 As shown in the experiment of Fessler and Eaton [19], distinct turbulence attenu-
414 ation was found for the two larger St cases (1 & 3). Over the region of ($y/H > 1$)

extensive turbulence attenuation is shown across all five plots (corresponds to Fig. 9 in Fessler and Eaton [19]). Across locations ($x/H = 2 \& 5$) there is attenuation across the range ($y/H > 1$) which shows that as mass loading is increased the turbulence is suppressed and below this range ($y/H < 1$) the turbulence is unaffected. This behaviour of turbulence attenuation was accurately predicted by the RA-TFM and the corresponding plots to those in [19] are Figs 6 & 8. The predictions are in good agreement with the experimental measurements. The turbulence attenuation for case 1 is as much as 35% showing a large reduction over the region of ($y/H > 1$) at ($x/H = 2$) on Fig. 8.

Below ($y/H < 1$) very little turbulence attenuation was observed, this corresponds to the shear layer and recirculation zone. From the simulations carried out, the particle turbulent quantities k_p & ε_p are produced and dissipated primarily near and at the wall and step (shear layer), with the contribution in the recirculation zone ($y/H < 1$) being several orders of magnitudes smaller. When considering the form of the coupling terms (see Section 2) it is evident why the turbulence attenuation is small in this region. This also follows from the lack of particles within the recirculation region due to the large St of all cases ($St > 1$) as the particles are not dragged into the eddy in the same way a particle of ($St < 1$) would.

The turbulence attenuation was accurately captured across all three particle classes for the RA-TFM. For the MPM an over-prediction of the fluid turbulent kinetic energy was observed. It was found that the form of the velocity correlation coupling term posed two major problems, the first being that as the term $k_{fp} = \sqrt{k_f \Theta_p}$ contains Θ_p directly the evolution of the term is adversely affected as the granular temperature equation evolves too quickly. This behaviour is recognised in two fluid model codes, and typically an upper limit is employed to constrain the initial stages of the solution to increase robustness.

Secondly, the source term of this form exists in both the k & ε transport equations, this leads to the formulation of the turbulent kinetic energy dissipation equation given by Elghobashi and Abou-Arab [16], which has been shown to yield incorrect behaviour [21], mainly as a consequence of failing to differentiate between the correlated and uncorrelated turbulent kinetic energy. Conversely, two coupling source terms are used in this work k_{fp} and ε_{fp} respectively. This allows the coupling of the turbulent kinetic energy k_i and dissipation ε_i equations of both phases to contain source terms that are of the same physical attribute i.e. particle turbulent kinetic energy and particle turbulence kinetic energy dissipation both contain separate coupling terms, which ensures conservation of energy between the two phases. In addition to this no numerical limiting was needed as the evolution of the granular temperature was controlled by the production of turbulent kinetic energy dissipation.

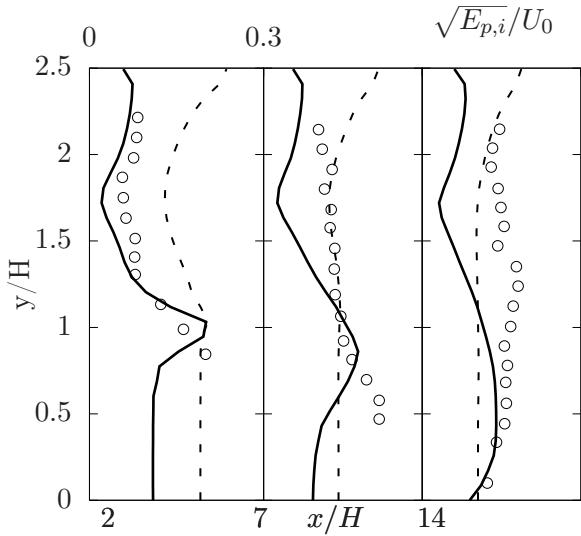


Figure 9: Particle turbulent intensity for case 1. Solid line showing the RA-TFM and the dashed line showing the MPM. Data from Fessler and Eaton [18] with a mass-loading of 40%.

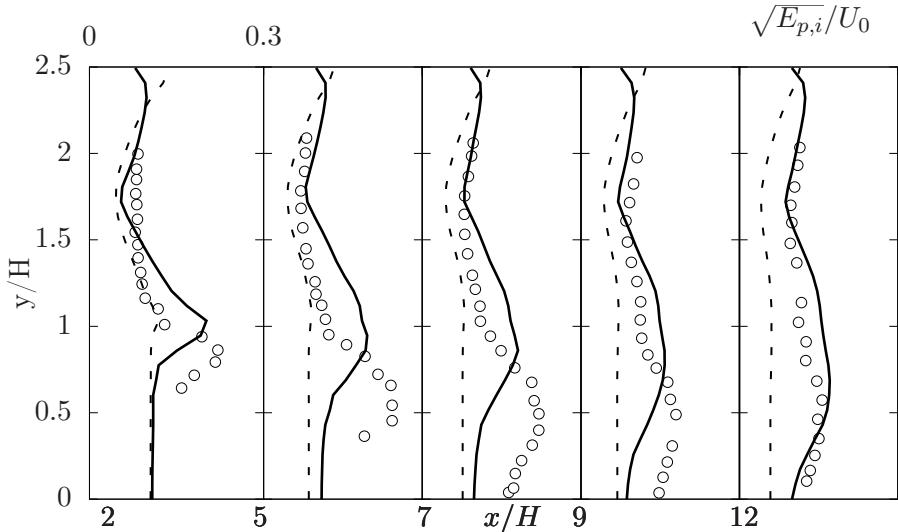


Figure 10: Particle turbulent intensity for case 3. Solid line showing the RA-TFM and the dashed line showing the MPM. Data from Fessler and Eaton [19] with a mass-loading of 10%.

453 3.3. *Particle phase turbulence*

454 From the experimental measurements it can be seen that the particles are being
 455 heavily influenced by the fluid phase's shear layer. This is true for both cases in-
 456 volving large St as shown in Fig 9 & 10. Looking at the results from the MPM the

prediction of the shear layer can not be seen across each location. Conversely, the RA-TFM is able to predict the presence of the shear layer and crucially convect it downstream. This feature is difficult to predict as the particles disperse and their fluctuating energy becomes more uniform across the profile. This result was almost exclusively attributed to the solution of the $k_p - \varepsilon_p$ transport equation.

As shown in Février et al. [20] the decomposition of the particle fluctuation energy into two components which reads, $\kappa_p = k_p + 3/2\Theta_p$ (following the notation of [21]) was needed when accounting for the particle's overall motion. Due to the step, turbulent scales at the integral scale are dominating the flow and as a result the large scale motions are the most relevant. This is reflected in the predictions of this model and highlighted when contrasted with the predictions of the MPM. Without the $k_p - \varepsilon_p$ transport model, the influence of the step is not captured and an under-prediction of the turbulent particle kinetic energy is seen. Table 3 shows the integral time scales associated with both flow regimes. This characteristic time scale associated with the particle turbulent kinetic energy enables the prediction of the shear layer and allows for the successful prediction of the predominant turbulent behaviour found in experiment across both Figs. 9 & 10.

Table 3: Table of integral time scales for each phase.

Flow regime	T_p	T_f
Channel flow centre line	0.04s	0.04s
Shear layer	4.09ms	5.12ms

In the MPM this definition of the particle-phase time scale is not present. The calculation of turbulent kinetic energy (granular temperature Θ_p) is heavily reliant on the constitutive relation of thermal conductivity, shear viscosity and bulk viscosity. As shown in both Figs. 9 & 10 the absence of the shear layer is demonstrated. The profiles are within the correct order of magnitude but the profile remains flat and largely unaffected by the step.

The $k_p - \varepsilon_p$ transport equation is modelled in an analogous manner to the single-phase $k - \varepsilon$ turbulence model using similar closure relations [41]. As a result some of the models well-known limitations are directly inherited. The Boussinesq approximation is one such limitation of RANS models of this form and introduces isotropy into the model; specifically the Reynolds stresses are assumed to be a scalar function of the mean velocity gradients. This introduction of isotropy has quite clear implications for the prediction of turbulent structures. The shear layer simulated in this studied is dominated by both the production term, Π_p and the turbulent dissipation

488 term, ε_p ; the former is where the isotropy is introduced and that is why it is so influ-
 489 ential. It has been shown by Simonin [44, 45] that the particle turbulence Reynold
 490 stresses are highly anisotropic and require transport equations for each term. This
 491 is a clear limitation of the current model and from the performance of the current
 492 RA-TFM an introduction of anisotropy for at least the particle phase is vital in an
 493 accurate prediction of the particle phase energy behaviour. Recently, Capecelatro
 494 et al. [7, 8] has derived a multiphase Reynolds-stress model which could fill this gap.
 495 Its application to a flow configuration similar to the one used in this work would be
 496 particularly interesting.

497 *3.4. Particle wall boundary conditions*

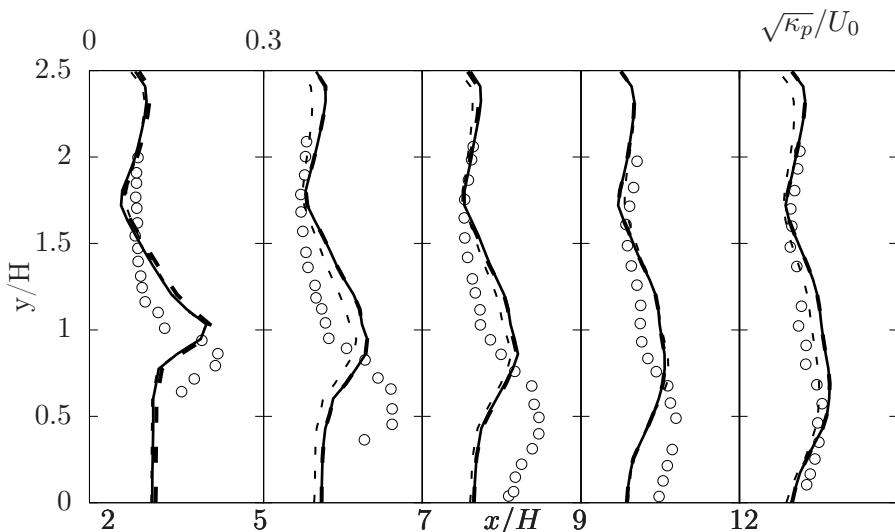


Figure 11: Particle turbulent intensity for case 3. Solid line showing the $\phi = 0.001$, thick dashed line showing $\phi = 0.01$ and the thin dashed line showing $\phi = 0$. Data from Fessler and Eaton [19] with a mass-loading of 10%.

498 As noted in Fessler and Eaton [19], the particles tend to conserve almost all
 499 their energy when interacting with the wall and consequently spend very little time
 500 interacting with it. As a consequence of this observation the specularity coefficient
 501 was varied from 0-0.01 in order to ascertain its effect on the numerical predictions.

502 Fig. 11 shows the particle turbulent intensity prediction of the RA-TFM with
 503 varying specularity coefficients. Immediately a general observation can be made;
 504 the particle phase wall boundary conditions have a relatively small impact on the
 505 prediction of particle fluctuation energy. This is to be expected as the particles

506 spend very little time interacting with the wall and the particle fluctuation energy
507 is dominated by the production in the shear layer.

508 The biggest difference can be seen by comparing $\phi = 0$ with $\phi = 0.01 - 0.001$,
509 and looking at locations ($x/H = 5$ - $x/H = 12$). The free slip condition exerts its
510 influence on the prediction immediately downstream of the shear layer, this results
511 in an underestimation in comparison with the larger ϕ values. When comparing
512 this result with the experimental data it seems that the prediction lies closer to the
513 measured values, this is seen most clearly at location ($x/H = 5$) across ($y/H > 1$)
514 and across the whole profile at location ($x/H = 12$).

515 When comparing the near wall predictions of particle fluctuation energy it can
516 be seen that there is a slight under-prediction when comparing $\phi = 0$ and $\phi =$
517 $0.01 - 0.001$. This is to be expected as a higher specularity coefficient will result in
518 a higher value of particle fluctuation energy due to the production of mean particle
519 shear. All three simulations under-predict the near wall behaviour, this result is
520 attributed to the lack of particle phase fluctuation anisotropy, but put more explicitly
521 the experimental observations show that the particle fluctuation energy is stretched
522 in the wall-normal direction. This stretching continues up to the wall (at $x/H =$
523 7), the RA-TFM used in this work can not predict this behaviour due to the inherit
524 assumptions made throughout.

525 A specularity coefficient value of 0.1 was tested but yielded unphysical results.
526 [3, 61] also found that a low specularity coefficient was representative of high velocity,
527 dilute fluid-particle flow. The unphysical results were due to the lowering of the slip
528 velocity near the wall. The mean velocity profiles for the fluid and particle phase
529 tend to converge as the no slip condition ($\phi = 1$) is approached. At the relatively
530 high speed velocities used in this study this resulted in a gross overestimation of
531 the particle fluctuation energy. An explanation for this behavior is as follows, the
532 high specularity coefficient at the wall promotes "sticking" of the particles. As these
533 particles are stuck at the wall and then released they begin to produce mean shear
534 in the particle phase momentum equation. This shearing which is imposed by the
535 boundary condition results in an overestimation of turbulence production resulting
536 in excessively large values of the particle phase fluctuation energy.

537 3.5. RA-TFM discussion

538 Given the plethora of terms required to close the RA-TFM, a discussion of the
539 relevant and negligible terms is warranted. The discussion is confined to the closure
540 modelled terms with transport, convective and Laplacian terms omitted. We begin
541 our discussion by focusing on the turbulent dispersion/drag/drift term which can
542 be found in Eq. 2 & 4 of Section 2. Throughout the study this term remained

negligible and this was true for both the glass particles of case 1 & 2 and the copper particles of case 3. The reasons for this are two-fold: across the width of the pipe the gradient of the volume fraction remains small due to the uniformity imposed by gravity. This uniformity of the particle volume fraction across the pipe width of near identical conditions has also been reported by Vreman [54] & Yamamoto et al. [57]. This uniformity changes in the presence of the shear layer as there is a smaller particle volume fraction in the recirculating region than in the shear layer. This large particle volume fraction gradient is directly opposed by the increase of the St, as highlighted previously in Section 3.1 the St increases dramatically in the shear layer thus decreasing the amount of dispersion.

Next, the turbulent pressure term, which contains the term $2/3\rho_p\alpha_pk_p$, remains negligible throughout. This term arises due to compressibility effects which are expected to be high for large Mach numbers, which was not the case in this work. The covariance of the volume fraction and fluid pressure gradient remains very small throughout; this was to be expected as it represents the fluctuations in the buoyancy force. Due to the high density ratio, ρ_p/ρ_f the flow was not considered buoyancy driven and so this term was negligible.

The velocity-fluid-pressure-gradient covariance term was found to be negligible for glass particles but was vital in the successful prediction of the mean particle velocities of the copper particles. In fact, without this term the mean particle velocities were over-predicted by up to 10%. If we revisit this term it becomes clear as to why; the term takes the form $\alpha_p\rho_p\mathbf{g}$ (neglecting the constant density ratio) and when the particle volume fraction is present there is a static pressure gradient of the particles acting on the mean momentum of both the particle and fluid phase. This means that the mean slip velocity is affected by the imposed pressure gradient. In this work it resulted in a reduction of the mean particle velocities. Consequently, this term can not be neglected for particles with large densities or (put more generally) for large St.

For the multiphase turbulence modelling coupling terms found in Eq. 5, 6, 10 & 11, only the velocity correlation terms were found to be relevant for this flow. These terms are responsible for the turbulence attenuation, so it follows that they should remain. The volume-fraction-velocity correlation was found to be negligible due to the high density ratio used in this work, but for small density ratios this term may be relevant. The mean slip term was found to be non-negligible but its magnitude was not enough to affect the solution. Interestingly, in this work there was a large mean slip value near the wall which is where this term is expected to be at its largest. Due to the use of single-phase wall functions this behaviour may have been suppressed resulting in an underestimation.

581 Finally, a note on the limitations of the current approach. As shown in Fox [21] the
582 hydrodynamic model (volume fraction, particle velocity and granular temperature)
583 is derived from the mesoscale model i.e the Boltzmann equation, using a Chapman-
584 Enskog expansion (shown in Garzó et al. [22]) in powers of Knudsen number. It
585 is owing to this mathematical linearisation that the hydrodynamical model is not
586 valid for large values of Knudsen number ($\text{Kn} > 1$). The hydrodynamical model, as
587 hinted at, is therefore most accurate for small values of Knudsen number ($\text{Kn} \ll 1$),
588 a feature of collisionally dominated flows (see Table 4 for Kn).

589 The flow regime used in this study is characterized as moderately dilute where
590 collisions are expected. For the smallest mass loading (case 3), the Knudsen number
591 was found to be ($\text{Kn} < 10^{-4}$) across the pipe, whilst for higher mass loadings the
592 Knudsen number was much lower. As this constraint is one across the whole domain
593 this can not always be fulfilled. When considering a fluidised bed for example, the
594 upper region of the chamber will not fulfill this criteria as no collisions are present as
595 there are no particles. For this study the recirculating region posed a problem as only
596 a small number of particles were present in the region. This meant that in this region
597 the Knudsen number would fluctuate due to the vortex shedding and temporarily
598 be $\mathcal{O}(1)$, compromising the validity of the solution. It was found that the value of
599 granular temperature was relatively high which may have allowed for near equilibrium
600 conditions thus satisfying the hydrodynamic constraint. Unfortunately, this type of
601 constraint is unavoidable when employing a hydrodynamical description and this can
602 not necessarily be fulfilled in every region of the flow. For a more flexible approach,
603 kinetic-based equations can be formulated from moments (see [5, 32, 35, 39]).

604 **4. Conclusions**

605 The current work has investigated turbulent attenuation of moderately dilute flow
606 in a vertically orientated backward-facing step using a Reynolds-Averaged Two-Fluid
607 model. The RA-TFM model is implemented into OpenFOAM with the implementa-
608 tion procedure and the treatment of challenging terms provided. The model results
609 are compared against benchmark experimental data and also against the model of
610 Peirano and Leckner [40]. A summary of our findings are as follows:

- 611 1. Prediction of the mean particle velocities and both turbulence quantities (tur-
612 bulent kinetic energy of the fluid and particles) were in good agreement with
613 the benchmark experimental data.
- 614 2. The inclusion of the correlated k_p and uncorrelated Θ_p particle motion was
615 crucial in accurately predicting the behaviour of the turbulent shear layer.
616 This was further highlighted when compared to the turbulent particle kinetic
617 energy predictions from the modified Peirano model.
- 618 3. The form of the velocity correlation coupling terms, i.e. separate coupling
619 terms for turbulent kinetic energy and turbulent kinetic energy dissipation,
620 resulted in a successful prediction of turbulence attenuation.
- 621 4. A specularity coefficient value of 0.001 was found to be representative of the
622 particle-wall behaviour in this study. Changes in the specularity coefficient (ϕ
623 < 0.01) had very little effect on the particle fluctuation energy prediction.
- 624 5. It has been shown that the behaviour of the particles interacting in a shear layer
625 is highly anisotropic. The current predictions of this behaviour are limited due
626 to the reliance on the Boussinesq approximation which introduces isotropy into
627 the model.

628 **Nomenclature**

U_0	centreline velocity, [ms^{-1}]
C_D	drag coefficient, [-]
A_i	diagonal coefficients of the matrix
g	gravity, [ms^{-2}]
\mathbf{n}	unit vector normal to the wall, [-]
Re_p	particle Reynolds number, [-]
d_p	particle diameter, [m]
\mathbf{u}_i	velocity, [ms^{-1}]
\mathbf{u}_w	wall velocity, [ms^{-1}]
$\mathbf{u}_{p,w}$	particle slip velocity parallel to the wall, [ms^{-1}]
\mathbf{u}_p''	particle velocity fluctuation w.r.t PA velocity, [ms^{-1}]
$\mathbf{u}_{p,i}''^2$	particle Reynolds stress component in direction i, [ms^{-1}]
$\mathbf{u}_f'''^2$	fluid velocity fluctuation w.r.t PA velocity, [ms^{-1}]
h	pipe width, [m]
p_i	pressure, [Pa]
g_0	radial distribution coefficient, [-]
H	step height, [m]
t	time, [s]
k_i	turbulent kinetic energy, [m^2s^{-2}]

629 *Greek letters*

α_i	volume fraction, [−]
$\alpha_{p,max}$	maximum particle volume fraction, [−]
β	momentum exchange coefficient, [$\text{kgm}^{-3}\text{s}^{-1}$]
ε_i	turbulent kinetic energy dissipation, [m^2s^{-3}]
Θ_p	granular temperature, [m^2s^{-2}]
κ_p	particle fluctuation energy, [m^2s^{-2}]
$\kappa_{\Theta s}$	diffusion coefficient for granular energy, [$\text{kgm}^{-1}\text{s}^{-1}$]
μ_i	shear viscosity, [$\text{kgm}^{-1}\text{s}^{-1}$]
$\mu_{i,t}$	turbulent shear viscosity, [$\text{kgm}^{-1}\text{s}^{-1}$]
ν_i	kinematic viscosity, [m^2s^{-1}]
$\nu_{i,t}$	turbulent kinematic viscosity, [m^2s^{-1}]
ρ_i	density, [kgm^{-3}]
$\bar{\sigma}_f$	fluid phase stress tensor, [$\text{kgm}^{-1}\text{s}^{-2}$]
$\bar{\sigma}_p$	particle phase stress tensor, [$\text{kgm}^{-1}\text{s}^{-2}$]
τ_d	particle relaxation time, [s]

630 *Subscripts*

1	RA-TFM
2	MPM
f	fluid
i	general index
p	particle
x	x direction
y	y direction
z	z direction

631 *Superscripts*

"	PA particle velocity fluctuation
""	PA fluid velocity fluctuation

632 *Special notation*

$\langle \cdot \rangle$	Reynolds averaging operator
$\langle \cdot \rangle_i$	phase averaging operator associated with phase i

Table 4: Model characteristics.

$$\beta = \frac{\rho_p \alpha_p}{\tau_d} = \frac{3}{4} \frac{\alpha_p \alpha_f \rho_f \mathbf{u}_r}{d_p} C_d$$

$$C_d = \begin{cases} \frac{24}{Re_p} \left[1 + 0.15 Re_p^{0.287} \right] & \text{if } Re_p < 1000 \\ 0.44 & \text{if } Re_p \geq 1000 \end{cases}$$

$$Re_p = \frac{\rho_f d_p |\mathbf{u}_p - \mathbf{u}_f|}{\mu_f}$$

$$E_{p,1} \equiv \kappa_p = k_p + 3/2\Theta_p$$

$$E_{p,2} \equiv \kappa_p = 3/2\Theta_p$$

$$\tau_f \equiv T_f = k_f / \varepsilon_f$$

$$St = \tau_d / \tau_f$$

$$T_p = \frac{k_p}{\varepsilon_p}$$

$$Kn = \frac{\sqrt{\pi} d_p}{12 \alpha_p g_0 L}$$

$$\chi = \frac{\alpha_p \rho_p}{\alpha_f \rho_f}$$

$$e = 0.9$$

Table 5: Definition of variables.

$$\kappa_p = k_p + 1.5\Theta_p$$

$$\mu_f = \rho_f \nu_f$$

$$\mu_{ft} = \alpha_f \rho_f \nu_{ft} = \alpha_f \rho_f C_{f\mu} \frac{k_f^2}{\varepsilon_f}$$

$$\mu_p = \alpha_p \rho_p \nu_p = \frac{2\mu_{pdil}}{(1+e)g_0} \left[1 + \frac{4}{5}(1+e)g_0\alpha_p \right]^2 + \frac{4}{5}\alpha_p^2 \rho_p d_p g_0 (1+e) \left(\frac{\Theta_p}{\pi} \right)^{1/2}$$

$$\mu_{pdil} = \frac{5\sqrt{\pi}}{96} \rho_p d_p \Theta_p^{1/2}$$

$$\mu_{pt} = \alpha_p \rho_p \nu_{pt} = \alpha_p \rho_p C_{p\mu} \frac{k_p^2}{\varepsilon_p}$$

$$p_p = \rho_p \alpha_p \Theta_p + 2(1+e)\rho_p \alpha_p^2 g_0 \Theta_p$$

$$\gamma = \frac{12(1-e^2)g_o}{\sqrt{\pi}d_p} \alpha_p^2 \rho_p \Theta_p^{3/2}$$

$$\kappa_\Theta = \frac{2}{(1+e)g_0} \left[1 + \frac{6}{5}(1+e)g_0\alpha_p \right]^2 \kappa_{\Theta,dil} + 2\alpha_p^2 \rho_p d_p g_0 (1+e) \left(\frac{\Theta_p}{\pi} \right)^{\frac{1}{2}}$$

$$\kappa_{\Theta,dil} = \frac{75}{384} \sqrt{\pi} \rho_p d_p \Theta_p^{1/2}$$

$$g_0 = \left[1 - \left(\frac{\alpha_p}{\alpha_{p,max}} \right)^{\frac{1}{3}} \right]^{-1}$$

$$\bar{\mathbf{S}}_{\mathbf{p}} = \frac{1}{2} [\nabla \mathbf{u}_p + (\nabla \mathbf{u}_p)^T] - \frac{1}{3} \nabla \cdot \mathbf{u}_p \mathbf{I}$$

$$\bar{\mathbf{S}}_{\mathbf{f}} = \frac{1}{2} [\nabla \mathbf{u}_f + (\nabla \mathbf{u}_f)^T] - \frac{1}{3} \nabla \cdot \mathbf{u}_f \mathbf{I}$$

$$k_{fp} = \beta_k \sqrt{k_f k_p}$$

$$\varepsilon_{fp} = \beta_\varepsilon \sqrt{\varepsilon_f \varepsilon_p}$$

Table 6: Definition of phase-averaged variables.

$$\alpha_p = \langle \alpha_p \rangle$$

$$\alpha_f = \langle \alpha_f \rangle$$

$$\mathbf{u}_p = \langle \mathbf{u} \rangle_p$$

$$\mathbf{u}_f = \langle \mathbf{u} \rangle_f$$

$$\Theta_p = \langle \Theta \rangle_p$$

$$k_p = \frac{1}{2} \langle \mathbf{u}_p'' \cdot \mathbf{u}_p'' \rangle_p$$

$$k_f = \frac{1}{2} \langle \mathbf{u}_f''' \cdot \mathbf{u}_f''' \rangle_f$$

$$\varepsilon_p = \frac{1}{\rho_p \alpha_p} \langle \bar{\boldsymbol{\sigma}}_p : \nabla \mathbf{u}_p'' \rangle$$

$$\varepsilon_f = \frac{1}{\rho_f \alpha_f} \langle \bar{\boldsymbol{\sigma}}_f : \nabla \mathbf{u}_f''' \rangle$$

$$\bar{\boldsymbol{\sigma}}_p = \mu_p [\nabla \mathbf{u}_p + (\nabla \mathbf{u}_p)^T] - \frac{1}{3} \mu_p \nabla \cdot \mathbf{u}_p \mathbf{I}$$

$$\bar{\boldsymbol{\sigma}}_f = \mu_f [\nabla \mathbf{u}_f + (\nabla \mathbf{u}_f)^T] - \frac{1}{3} \mu_f \nabla \cdot \mathbf{u}_f \mathbf{I}$$

$$\mathbf{u}_p'' = \mathbf{u}_p - \langle \mathbf{u}_p \rangle_p$$

$$\mathbf{q}_\Theta = \langle \mathbf{q}_\Theta \rangle_p = \frac{\kappa_\Theta}{\alpha_p \rho_p} \nabla \Theta_p$$

$$\mathbf{u}_f''' = \mathbf{u}_f - \langle \mathbf{u}_f \rangle_f$$

$$\langle \mathbf{u}_p \rangle_p = \langle \alpha_p \mathbf{u}_p \rangle / \langle \alpha_p \rangle$$

$$\langle \mathbf{u}_f \rangle_f = \langle \alpha_f \mathbf{u}_f \rangle / \langle \alpha_f \rangle$$

$$\mathbf{u}_p'' \mathbf{u}_p'' = \langle \mathbf{u}_p'' \mathbf{u}_p'' \rangle_p$$

633 **5. Appendix**

634 Here the equations used in the MPM are presented. A full explanation of the
 635 equations can be found in Peirano and Leckner [40]. The modification of the model
 636 comes from the type of closure for the particle-fluid covariance in which the isotropic
 637 model of Sinclair and Mallo [46] is used.

638 The continuity equations for each phase read:

639

$$\frac{\partial \alpha_p \rho_p}{\partial t} + \nabla \cdot \alpha_p \rho_p \mathbf{u}_p = 0 \quad (42)$$

640

641

$$\frac{\partial \alpha_f \rho_f}{\partial t} + \nabla \cdot \alpha_f \rho_f \mathbf{u}_f = 0 \quad (43)$$

642

643 The momentum balance equation for each phase:

644

$$\frac{\partial \alpha_p \rho_p \mathbf{u}_p}{\partial t} + \nabla \cdot \alpha_p \rho_p \mathbf{u}_p \mathbf{u}_p = \alpha_p \nabla \cdot \boldsymbol{\tau}_p - \alpha_p \nabla p_f - \nabla p_p + \beta(\mathbf{u}_f - \mathbf{u}_p) + \alpha_p \rho_p \mathbf{g} \quad (44)$$

645

646

$$\frac{\partial \alpha_f \rho_f \mathbf{u}_f}{\partial t} + \nabla \cdot \alpha_f \rho_f \mathbf{u}_f \mathbf{u}_f = \alpha_f \nabla \cdot \boldsymbol{\tau}_f - \alpha_f \nabla p_f - \beta(\mathbf{u}_p - \mathbf{u}_f) + \alpha_f \rho_f \mathbf{g} \quad (45)$$

647 *5.1. Kinetic Theory of Granular Flow*

648 Following the kinetic theory of granular flow the closure of the particle pres-
 649 sure, shear and bulk viscosities can be provided. The granular temperature Θ_p is
 650 introduced as a measure of the particle velocity fluctuations [23].

$$\Theta_p = \frac{1}{3} \mathbf{u}_p''^2 \quad (46)$$

651 where \mathbf{u}_p'' is the particle fluctuation velocity. A balance equation is introduced for
 652 the granular energy ($\frac{3}{2}\Theta_p$) to satisfy the continuity and momentum balance for both
 653 phases. The fluctuation energy conservation for the particles is then given as:

$$\frac{3}{2} \left[\frac{\partial \alpha_p \rho_p \Theta_p}{\partial t} + \nabla \cdot \alpha_p \rho_p \Theta_p \mathbf{u}_p \right] = (p_p \bar{I} + \bar{\tau}) : \nabla \mathbf{u}_p + \nabla \cdot \kappa \nabla \Theta_p - \gamma_p + J_{vis} + J_{slip} \quad (47)$$

654 where the first term on the RHS is the fluctuation energy created by the shearing in
 655 the particle phase. The second term is associated with the diffusion of fluctuating
 656 energy, the third term is responsible for the dissipation due to inelastic collisions.

657 Finally, J_s is either dissipation of granular temperature due to viscous damping
 658 and/or creation of granular temperature from the slip between the fluid and particles.
 659 Both terms can be written more intuitively to read:

$$J_{vis} + J_{slip} = \beta(\mathbf{u}_p''\mathbf{u}_p'' - \mathbf{u}_p''\mathbf{u}_f'') \quad (48)$$

660 The first term can be modeled as $3\Theta_p$ according to Gidaspow [23] and the last term
 661 can be modeled as k_{pf} .

662 5.2. Turbulence modelling

663 The transport equations for the fluid phases turbulence model $k_f - \varepsilon_f$ reads as
 664 follows:

$$\frac{\partial \alpha_f \rho_f k_f}{\partial t} + \nabla \cdot \alpha_f \rho_f k_f \mathbf{u}_f = \nabla \cdot \alpha_f \rho_f \nu_t \nabla k_f + \alpha_f G - \alpha_f \rho_f \varepsilon_f + \Pi_{kf} \quad (49)$$

665

$$\frac{\partial \alpha_f \rho_f \varepsilon_f}{\partial t} + \nabla \cdot \alpha_f \rho_f \varepsilon_f \mathbf{u}_f = \nabla \cdot \alpha_f \rho_f \nu_t \nabla \varepsilon_f + \frac{\varepsilon_f}{k_f} [C_1 \alpha_f G - C_2 \alpha_f \rho_f \varepsilon_f + C_3 \Pi_{kf}] \quad (50)$$

666 Where G is the production of the turbulent kinetic energy, which is expressed as
 667 follows:

$$G = 2\nu \rho_f \left(\bar{\mathbf{S}}_f : \bar{\mathbf{S}}_f - \frac{2}{3} \text{tr}(\bar{\mathbf{S}}_f)^2 \cdot \mathbf{I} \right) + \frac{2}{3} \rho_f k_f \nabla \cdot (\mathbf{u}_f \mathbf{I}) \quad (51)$$

668 The term Π_{kf} accounts for turbulence modulation from particles and represents the
 669 velocity fluctuation correlation of each phase,

$$\Pi_{kf} = -\beta(2k_f - k_{pf} - \mathbf{u}_r \mathbf{u}_d) \quad (52)$$

670 The term \mathbf{u}_d accounts for turbulence dispersion and is also known as the drift velocity.
 671 Here it is given from the formulation of [44].

$$\mathbf{u}_d = -D_{sf} \left(\frac{1}{\alpha_p} \nabla \alpha_p - \frac{1}{\alpha_f} \nabla \alpha_f \right) \quad (53)$$

672 The fluid-particle covariance term is given by Sinclair and Mallo [46] and is also
 673 termed an isotropic model.

$$k_{pf} = c_{pf} \sqrt{k_f \Theta_p} \quad (54)$$

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