

Supergravity and $\text{IOSp}(3, 1|4)$ gauge theory

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Abstract

A new formulation of simple $D = 4$ supergravity in terms of the geometry of superspace is presented. The formulation is derived from the gauge theory of the inhomogeneous orthosymplectic group $\text{IOSp}(3, 1|4)$ on a $(4, 4)$ -dimensional base supermanifold by imposing constraints and taking a limit. Both the constraints and the limiting procedure have a clear *a priori* physical motivation, arising from the relationship between $\text{IOSp}(3, 1|4)$ and the super Poincaré group. The construction has similarities with the space-time formulation of Newtonian gravity.

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1 Introduction

The fact that general relativity is a theory of the dynamics of space-time geometry leads to the supposition that its supersymmetric version may also have a simple geometrical formulation, where the geometry involved is that of superspace. This does not appear to be the case, however. The analogue in superspace of general relativity, known as gauge supersymmetry [1], is certainly not supergravity, so one must consider a more complicated superspace geometry. Nath and Arnowitt [2] obtained supergravity from gauge supersymmetry by contracting the tangent-space group $\text{OSp}(3, 1|4)$ of the $(4, 4)$ -dimensional supermanifold to its $\text{SO}(3, 1)$ subgroup and taking a limit, while the standard Wess–Zumino formulation [3, 4] requires the imposition of constraints on the superspace torsion. A drawback of both these formalisms is that they have elements, viz the limiting procedure of the former and the constraints of the latter, that have no clear *a priori* physical motivation. In this article we seek a superspace formulation of supergravity that, in contrast to the Nath–Arnowitt and Wess–Zumino formulations, derives from an *a priori*, physically motivated principle.

We shall obtain simple $D = 4$ supergravity from the gauge theory of the inhomogeneous orthosymplectic group $\text{IOSp}(3, 1|4)$ on a $(4, 4)$ -dimensional base supermanifold. This is achieved by imposing constraints and taking a limit, however both the constraints and the limiting procedure are determined at the outset by physical

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considerations. We point out that the resulting superspace geometry has similarities with Cartan's space-time formulation of Newtonian gravity [14]. Indeed, the method used here to derive supergravity from $\text{IOSp}(3, 1|4)$ gauge theory can also be used to derive Cartan's picture of Newtonian gravity from general relativity, when the latter is formulated as Poincaré gauge theory.

As explained above, the work described here was undertaken because of drawbacks of the Nath–Arnowitt and Wess–Zumino approaches to supergravity. We should mention a constraint-free, superfield formulation of simple $D = 4$ supergravity due to Siegel and Gates [5], and Ogievetsky and Sokatchev (see [6] and references therein). This superfield approach has the advantage of being analogous to the superfield formulation of supersymmetric Yang–Mills theory.

Section 2 uses superspace to explore the relationship between $\text{IOSp}(3, 1|4)$ and the super Poincaré group; we find that the latter can be obtained from the former by the imposition of constraints and the taking of a limit. In Section 3 we derive the super Lie algebra of $\text{IOSp}(3, 1|4)$ and show that one obtains the super Poincaré algebra by applying the constraints and limit of Section 2. Section 4 provides the necessary background in $\text{IOSp}(3, 1|4)$ gauge theory. In Section 5 we apply the constraints and limit of Section 2 to the gauge potential and curvature of $\text{IOSp}(3, 1|4)$ gauge theory and thereby derive supergravity. We then point out that our results make clear a geometrical relationship between supergravity and gauge supersymmetry that has similarities with the geometrical relationship between Newtonian gravity and general relativity.

2 Superspace and the super Poincaré group

Superspace is remarkable as a supermanifold in that its coordinates¹ $Z^\Lambda = (x^\mu, \theta^\alpha, \theta^{\dot{\alpha}})$ have different mass dimensions $[Z^\Lambda]$:

$$[x^\mu] = -1, \quad [\theta^\alpha] = [\theta^{\dot{\alpha}}] = -\frac{1}{2}. \quad (1)$$

The reason for this is that the scale of the supersymmetry generators $Q_\alpha, Q_{\dot{\alpha}}$ is chosen to avoid introducing a physically irrelevant constant into the supersymmetry algebra [8]. Eqn. (1) has the consequence (though this is sometimes ignored) that the canonical metric in flat Riemannian superspace [7] cannot have dimensionless components; for distance to be real with units of x^μ the canonical metric is

$${}_\Lambda H_\Pi = \begin{pmatrix} \eta_{\mu\nu} & 0 & 0 \\ 0 & k^2 \varepsilon_{\alpha\beta} & 0 \\ 0 & 0 & -k^2 \varepsilon_{\dot{\alpha}\dot{\beta}} \end{pmatrix} \Rightarrow {}^\Lambda H^\Pi = \begin{pmatrix} \eta^{\mu\nu} & 0 & 0 \\ 0 & -\frac{1}{k^2} \varepsilon^{\alpha\beta} & 0 \\ 0 & 0 & \frac{1}{k^2} \varepsilon^{\dot{\alpha}\dot{\beta}} \end{pmatrix}, \quad (2)$$

where k is a real constant with $[k] = -\frac{1}{2}$. We then have

$$H(\Delta Z, \Delta Z) = \Delta Z^\Lambda {}_\Lambda H_\Pi \Delta Z^\Pi = \Delta x_\mu \Delta x^\mu + k^2 \Delta \theta_\alpha \Delta \theta^\alpha - k^2 \Delta \theta_{\dot{\alpha}} \Delta \theta^{\dot{\alpha}},$$

¹Superspace coordinate indices are denoted by $\Lambda = (\mu, \alpha, \dot{\alpha})$, orthosymplectic-frame indices by $A = (m, a, \dot{a})$; rules for super index positioning and manipulation are those of DeWitt [7]. The space-time metric has signature +2 and the Infeld–van der Waerden symbols are $\sigma^m_{a\dot{a}} = \frac{i}{\sqrt{2}}(I, \vec{\sigma})$, where $\vec{\sigma}$ are the Pauli matrices.

so distance squared is real with units $[x^\mu]^2 = -2$.

The group of coordinate transformations that leaves (2) unchanged is the analogue in superspace of the Poincaré group in space-time; it is the inhomogeneous orthosymplectic group $\text{IOSp}(3, 1|4)$, i.e. the orthosymplectic group $\text{OSp}(3, 1|4)$ plus translations. An $\text{IOSp}(3, 1|4)$ transformation in superspace has the infinitesimal form

$$Z'^\Lambda = Z^\Lambda + {}^\Lambda\Lambda_\Pi Z^\Pi + \Xi^\Lambda, \quad (3)$$

where ${}_\Lambda\Lambda_\Pi = -(-1)^{\Lambda+\Pi+\Lambda\Pi} {}_\Pi\Lambda_\Lambda$ (antisupersymmetric),

with constant parameters ${}^\Lambda\Lambda_\Pi$ and Ξ^Λ . Gauge supersymmetry, the analogue in superspace of general relativity, can be constructed by gauging $\text{IOSp}(3, 1|4)$ on superspace in complete analogy to the manner (discussed in [9], for example) in which general relativity is obtained by gauging the Poincaré group on space-time.

Consideration of the geometry of a (4,4)-dimensional supermanifold thus leads naturally to the group $\text{IOSp}(3, 1|4)$; however this group has nothing to do with physics. The group relevant to supersymmetric physics is the super Poincaré group, which gives the following infinitesimal transformation on superspace [8, 7]:

$$x'^\mu = x^\mu + a^\mu + \zeta^\mu{}_\nu x^\nu + \sigma^\mu{}_{\alpha\dot{\alpha}}(\xi^\alpha \theta^{\dot{\alpha}} - \theta^\alpha \xi^{\dot{\alpha}}), \quad (4)$$

$$\theta'^\alpha = \theta^\alpha + \xi^\alpha + \frac{1}{2}\zeta_{\mu\nu} \sigma^{\mu\nu\alpha}{}_\beta \theta^\beta, \quad (5)$$

$$\theta'^{\dot{\alpha}} = \theta^{\dot{\alpha}} + \xi^{\dot{\alpha}} + \frac{1}{2}\zeta_{\mu\nu} \sigma^{\mu\nu\dot{\alpha}}{}_{\dot{\beta}} \theta^{\dot{\beta}}. \quad (6)$$

Note that the absence of a dimensionful constant in the last term of (4) is a consequence of (1); this term is also remarkable because in it we have the translation parameter of θ appearing as a “rotation” parameter of x . The metric (2) is, of course, not invariant under a super Poincaré transformation; this becomes obvious when we write (4)–(6) in matrix form:

$$\begin{pmatrix} x'^\mu \\ \theta'^\alpha \\ \theta'^{\dot{\alpha}} \end{pmatrix} = \begin{pmatrix} x^\mu \\ \theta^\alpha \\ \theta^{\dot{\alpha}} \end{pmatrix} + \begin{pmatrix} \lambda^\mu{}_\nu & \sigma^\mu{}_{\beta\dot{\alpha}} \xi^{\dot{\alpha}} & \sigma^\mu{}_{\alpha\dot{\beta}} \xi^\alpha \\ 0 & \frac{1}{2}\lambda_{\mu\nu} \sigma^{\mu\nu\alpha}{}_\beta & 0 \\ 0 & 0 & \frac{1}{2}\lambda_{\mu\nu} \sigma^{\mu\nu\dot{\alpha}}{}_{\dot{\beta}} \end{pmatrix} \begin{pmatrix} x^\nu \\ \theta^\beta \\ \theta^{\dot{\beta}} \end{pmatrix} + \begin{pmatrix} a^\mu \\ \xi^\alpha \\ \xi^{\dot{\alpha}} \end{pmatrix}. \quad (7)$$

In the notation of (3) the “rotation” matrix in (7) is

$${}^\Lambda\Lambda_\Pi = \begin{pmatrix} \lambda^\mu{}_\nu & \sigma^\mu{}_{\beta\dot{\alpha}} \xi^{\dot{\alpha}} & \sigma^\mu{}_{\alpha\dot{\beta}} \xi^\alpha \\ 0 & \frac{1}{2}\lambda_{\mu\nu} \sigma^{\mu\nu\alpha}{}_\beta & 0 \\ 0 & 0 & \frac{1}{2}\lambda_{\mu\nu} \sigma^{\mu\nu\dot{\alpha}}{}_{\dot{\beta}} \end{pmatrix} \quad (8)$$

$$\Rightarrow {}_\Lambda\Lambda_\Pi = {}_\Lambda H_\Sigma {}^\Sigma\Lambda_\Pi = \begin{pmatrix} \lambda_{\mu\nu} & \sigma_{\mu\beta\dot{\alpha}} \xi^{\dot{\alpha}} & \sigma_{\mu\alpha\dot{\beta}} \xi^\alpha \\ 0 & -\frac{1}{2}k^2 \lambda_{\mu\nu} \sigma^{\mu\nu}{}_{\alpha\beta} & 0 \\ 0 & 0 & \frac{1}{2}k^2 \lambda_{\mu\nu} \sigma^{\mu\nu}{}_{\dot{\alpha}\dot{\beta}} \end{pmatrix} \quad (9)$$

and the latter is clearly not antisupersymmetric; hence (7) does not leave (2) unchanged. One can proceed to find geometrical objects such as a metric and a connection that *are* invariant under the super Poincaré group of transformations (7), as

is done in [7] and [8]. The approach here however will be to explore the relationship between the super Poincaré group and the “natural”, canonical structures on superspace.

We can construct an antisymmetric matrix from (9) by inserting additional elements; the least modification necessary to achieve this results in the matrix

$${}^{\Lambda}\Lambda_{\Pi} = \begin{pmatrix} \lambda_{\mu\nu} & \sigma_{\mu\beta\dot{\alpha}} \xi^{\dot{\alpha}} & \sigma_{\mu\alpha\dot{\beta}} \xi^{\alpha} \\ \sigma_{\nu\alpha\dot{\alpha}} \xi^{\dot{\alpha}} & -\frac{1}{2}k^2\lambda_{\mu\nu} \sigma^{\mu\nu}{}_{\alpha\beta} & 0 \\ \sigma_{\nu\alpha\dot{\alpha}} \xi^{\alpha} & 0 & \frac{1}{2}k^2\lambda_{\mu\nu} \sigma^{\mu\nu}{}_{\dot{\alpha}\dot{\beta}} \end{pmatrix} \quad (10)$$

$$\Rightarrow {}^{\Lambda}\Lambda_{\Pi} = \begin{pmatrix} \lambda^{\mu}{}_{\nu} & \sigma^{\mu}{}_{\beta\dot{\alpha}} \xi^{\dot{\alpha}} & \sigma^{\mu}{}_{\alpha\dot{\beta}} \xi^{\alpha} \\ -\frac{1}{k^2}\sigma_{\nu}{}^{\alpha}{}_{\dot{\alpha}} \xi^{\dot{\alpha}} & \frac{1}{2}\lambda_{\mu\nu} \sigma^{\mu\nu\alpha}{}_{\beta} & 0 \\ \frac{1}{k^2}\sigma_{\nu\alpha}{}^{\dot{\alpha}} \xi^{\alpha} & 0 & \frac{1}{2}\lambda_{\mu\nu} \sigma^{\mu\nu\dot{\alpha}}{}_{\dot{\beta}} \end{pmatrix}. \quad (11)$$

By replacing the “rotation” matrix in (7) with (11) we obtain a transformation that does leave the metric (2) unchanged:

$$\begin{pmatrix} x'^{\mu} \\ \theta'^{\alpha} \\ \theta'^{\dot{\alpha}} \end{pmatrix} = \begin{pmatrix} x^{\mu} \\ \theta^{\alpha} \\ \theta^{\dot{\alpha}} \end{pmatrix} + \begin{pmatrix} \lambda^{\mu}{}_{\nu} & \sigma^{\mu}{}_{\beta\dot{\alpha}} \xi^{\dot{\alpha}} & \sigma^{\mu}{}_{\alpha\dot{\beta}} \xi^{\alpha} \\ -\frac{1}{k^2}\sigma_{\nu}{}^{\alpha}{}_{\dot{\alpha}} \xi^{\dot{\alpha}} & \frac{1}{2}\lambda_{\mu\nu} \sigma^{\mu\nu\alpha}{}_{\beta} & 0 \\ \frac{1}{k^2}\sigma_{\nu\alpha}{}^{\dot{\alpha}} \xi^{\alpha} & 0 & \frac{1}{2}\lambda_{\mu\nu} \sigma^{\mu\nu\dot{\alpha}}{}_{\dot{\beta}} \end{pmatrix} \begin{pmatrix} x^{\nu} \\ \theta^{\beta} \\ \theta^{\dot{\beta}} \end{pmatrix} + \begin{pmatrix} a^{\mu} \\ \xi^{\alpha} \\ \xi^{\dot{\alpha}} \end{pmatrix}. \quad (12)$$

The transformations (12) are a subset of the $\text{IOSp}(3,1|4)$ transformations (3), but they do not form a group (as is verified by a tedious calculation). Thus, the transformations (12) are a subset of $\text{IOSp}(3,1|4)$, not a subgroup. Nevertheless the $k \rightarrow \infty$ limit of (12) does give a group—the group of super Poincaré transformations (7). To reiterate this point, in the limit $k \rightarrow \infty$ the transformations (12) are no longer a subset of $\text{IOSp}(3,1|4)$, but rather form a new group, the super Poincaré group.

These considerations suggest how supergravity may be related to $\text{IOSp}(3,1|4)$ gauge theory. The latter gives us super one-form gauge potentials $A^A{}_{B\Lambda}$ and ${}^A E_{\Lambda}$ (the latter chosen to be the vielbein²) with values in the super Lie algebra of $\text{IOSp}(3,1|4)$. Now we have seen how to obtain the infinitesimal super Poincaré group from infinitesimal $\text{IOSp}(3,1|4)$ —extract all elements of infinitesimal $\text{IOSp}(3,1|4)$ of the form (12) and take $k \rightarrow \infty$ —so we can perform a similar operation with the respective Lie algebras. In this manner the potentials $A^A{}_{B\Lambda}$ and ${}^A E_{\Lambda}$ are turned into super Poincaré-algebra-valued objects and it is at this point that one might expect to see some physics.

Note that when $k \rightarrow \infty$ we lose the metric (2), though we can preserve a metric $\eta_{\mu\nu}$ in the bosonic sector. This is reminiscent of Cartan’s space-time formulation of Newtonian gravity [14], wherein one also does not have a metric in the full space (space-time) but only in a subspace (3-space). This analogy will be pursued further below.

²We thus have an *affine connection* [11] on the principle bundle with fibre $\text{IOSp}(3,1|4)$.

3 Super Lie algebra of $\text{IOSp}(3, 1|4)$ and the super Poincaré algebra

In order to proceed we require the super Lie algebra of $\text{IOSp}(3, 1|4)$, which we will derive by a method used in [10] to obtain the Poincaré algebra. We shall then demonstrate explicitly that this super Lie algebra becomes the super Poincaré algebra when we select the elements corresponding to (12) and take the limit $k \rightarrow \infty$.

It will be convenient to write both indices of the $\text{OSp}(3, 1|4)$ parameters ${}^\Lambda \Lambda_\Pi$ on the left; the rule for moving the lower index to the left is the same as if the upper index were absent [7]:

$${}^\Lambda_\Pi \Lambda := (-1)^\Pi {}^\Lambda \Lambda_\Pi. \quad (13)$$

We write the infinitesimal $\text{IOSp}(3, 1|4)$ element as

$$G(1 + \Lambda, \Xi) = 1 + \frac{i}{2} J^\Lambda_\Pi {}^\Pi_\Lambda \Lambda - iP_\Lambda \Xi^\Lambda. \quad (14)$$

Eqn. (14) does not serve to define the generators J^Λ_Π and P_Λ completely; we must specify how the group element acts on a representation space. We define the element (14) to act on a pure vector X in the representation space according to

$${}^i X' = {}^i X + \frac{i}{2} (-1)^{X(\Lambda+\Pi)} {}^i [J^\Lambda_\Pi(X)] {}^\Pi_\Lambda \Lambda - i(-1)^{X\Lambda} {}^i [P_\Lambda(X)] \Xi^\Lambda.$$

We arrange matters in this way so as to avoid the appearance in the group element (14) of factors of (-1) that are dependent on representation-space indices, and so that the generators obey the super Lie algebra of the group.³

We denote a non-infinitesimal group element by $G(L, A)$ and consider the product

$$G(L, A)G(1 + \Lambda, \Xi)G(L, A)^{-1}. \quad (15)$$

Now a non-infinitesimal transformation (L, A) , given by

$$Z'^\Lambda = {}^\Lambda L_\Pi Z'^\Pi + A^\Lambda,$$

has as its inverse the transformation $(L^{-1}, -L^{-1}A)$:

$$Z''^\Lambda = {}^\Lambda L^{-1}_\Pi Z'^\Pi - {}^\Lambda L^{-1}_\Pi A^\Pi = Z^\Lambda.$$

Also, a transformation (L, A) followed by a transformation (\bar{L}, \bar{A}) is a transformation $(\bar{L}L, \bar{L}A + \bar{A})$. Hence the product (15) is

$$G(L, A)G(1 + \Lambda, \Xi)G(L^{-1}, -L^{-1}A) = G(1 + L\Lambda L^{-1}, -L\Lambda L^{-1}A + L\Xi).$$

To first order in Λ, Ξ we therefore have from (14)

$$\begin{aligned} G(L, A) \left[1 + \frac{i}{2} J^\Lambda_\Pi {}^\Pi_\Lambda \Lambda - iP_\Lambda \Xi^\Lambda \right] G(L, A)^{-1} \\ = 1 + \frac{i}{2} J^\Lambda_\Pi {}^\Pi_\Lambda (L\Lambda L^{-1}) - iP_\Lambda {}^\Lambda (L\Xi - L\Lambda L^{-1}A) \end{aligned}$$

³These complicated issues are discussed both in generality and with many examples (including $\text{OSp}(m|n)$, but not $\text{IOSp}(m|n)$) in Chapters 3 and 4 of deWitt [7]

$$[(13)] = 1 + \frac{i}{2}(-1)^\Lambda J^\Lambda_{\Pi} {}^\Pi L_{\Xi}{}^\Xi \Lambda_{\Sigma}{}^\Sigma L^{-1}_{\Lambda} - iP_{\Lambda} ({}^\Lambda L_{\Pi}{}^\Xi \Xi^{\Pi} - {}^\Lambda L_{\Pi}{}^\Pi \Lambda_{\Xi}{}^\Xi L^{-1}_{\Sigma} A^{\Sigma}) \quad (16)$$

We wish to equate coefficients of ${}^\Pi \Lambda$ and Ξ^Λ in (16). In doing this we must take account of the antisupersymmetry of ${}^\Pi \Lambda_{\Lambda}$, as expressed in (3); from the antisupersymmetry we obtain

$$\begin{aligned} {}^\Lambda \Lambda_{\Pi} &= -(-1)^{\Pi(\Lambda+1)} {}^\Pi \Lambda^{\Lambda} \\ \implies {}^\Lambda_{\Pi} \Lambda &= -(-1)^{\Pi\Xi} {}^\Lambda H^{\Xi} {}^\Pi H_{\Sigma}{}^{\Sigma} \Lambda \\ \implies {}^\Lambda_{\Pi} \Lambda &= \frac{1}{2} ({}^\Lambda_{\Pi} \Lambda - (-1)^{\Pi\Xi} {}^\Lambda H^{\Xi} {}^\Pi H_{\Sigma}{}^{\Sigma} \Lambda). \end{aligned} \quad (17)$$

Using (17) and the fact that ${}^\Lambda H_{\Pi}$ vanishes if Λ and Π are of opposite type, we obtain from equating coefficients of ${}^\Pi \Lambda$ in (16)

$$\begin{aligned} G(L, A) J^\Lambda_{\Pi} G(L, A)^{-1} &= (-1)^{\Sigma+\Pi(\Lambda+\Sigma)+\Lambda\Sigma} J^{\Sigma}_{\Xi}{}^\Xi L_{\Pi}{}^\Lambda L^{-1}_{\Sigma} \\ &+ (-1)^{\Pi\Lambda} P_{\Xi}{}^\Xi L_{\Pi}{}^\Lambda L^{-1}_{\Sigma} A^{\Sigma} - P_{\Xi}{}^\Xi L_{\Phi}{}^\Phi H^{\Lambda}{}^{\Upsilon} H_{\Pi}{}^{\Upsilon} L^{-1}_{\Sigma} A^{\Sigma}, \end{aligned} \quad (18)$$

while equating coefficients of Ξ^Λ in (16) gives

$$G(L, A) P_{\Lambda} G(L, A)^{-1} = P_{\Pi}{}^\Pi L_{\Lambda}. \quad (19)$$

We now apply (18) and (19) to the case where $G(L, A)$ is an infinitesimal element $G(1 + \Lambda, \Xi)$; using (14), (18) now gives for terms of order Λ and Ξ

$$\begin{aligned} i \left[\frac{1}{2} J^{\Sigma}_{\Xi}{}^\Xi \Lambda - P_{\Sigma}{}^{\Sigma} \Xi, J^\Lambda_{\Pi} \right] &= J^{\Lambda}_{\Xi}{}^\Xi \Lambda_{\Pi} - (-1)^{\Sigma+\Pi(\Lambda+\Sigma)+\Lambda\Sigma} J^{\Sigma}_{\Pi}{}^\Lambda \Lambda_{\Sigma} \\ &+ (-1)^{\Pi\Lambda} P_{\Pi}{}^\Xi \Lambda - P_{\Xi}{}^\Xi H^{\Lambda}{}_{\Sigma} H_{\Pi}{}^{\Sigma}, \end{aligned} \quad (20)$$

while (19) gives

$$i \left[\frac{1}{2} J^{\Pi}_{\Sigma}{}^{\Sigma} \Lambda - P_{\Pi}{}^{\Pi} \Xi, P_{\Lambda} \right] = P_{\Pi}{}^\Pi \Lambda_{\Lambda}. \quad (21)$$

Equating coefficients of ${}^\Xi \Lambda$ and Ξ^Σ in (20) and (21) we obtain

$$\begin{aligned} J^\Lambda_{\Pi} J^{\Sigma}_{\Xi} - (-1)^{(\Lambda+\Pi)(\Sigma+\Xi)} J^{\Sigma}_{\Xi} J^\Lambda_{\Pi} &= i \left[(-1)^{\Sigma} {}^{\Sigma} \delta_{\Pi} J^{\Lambda}_{\Xi} - (-1)^{\Pi(1+\Sigma)} {}^\Pi H_{\Xi}{}^{\Upsilon} H^{\Sigma}{}^{\Upsilon} J^{\Lambda}_{\Upsilon} \right. \\ &\left. - (-1)^{\Pi(\Lambda+\Sigma)+\Lambda\Sigma} {}^\Lambda \delta_{\Xi} J^{\Sigma}_{\Pi} + (-1)^{\Pi(\Lambda+\Upsilon)} {}^\Lambda H^{\Sigma}{}^{\Upsilon} H_{\Xi}{}^{\Upsilon} J^{\Upsilon}_{\Pi} \right], \end{aligned} \quad (22)$$

$$J^\Lambda_{\Pi} P_{\Sigma} - (-1)^{\Sigma(\Lambda+\Pi)} P_{\Sigma} J^\Lambda_{\Pi} = i \left[-(-1)^{\Pi\Lambda} {}^\Lambda \delta_{\Sigma} P_{\Pi} + {}^{\Sigma} H_{\Pi}{}^{\Xi} H^{\Lambda}{}^{\Xi} P_{\Xi} \right], \quad (23)$$

$$P_{\Lambda} P_{\Pi} - (-1)^{\Lambda\Pi} P_{\Pi} P_{\Lambda} = 0. \quad (24)$$

The supercommutation relations (22)–(24) constitute the super Lie algebra of $\text{IOSp}(3, 1|4)$.

Now consider the subset (12) of $\text{IOSp}(3, 1|4)$ transformations. In this subset the group parameters ${}^\Lambda \Lambda_{\Pi}$ and Ξ^Λ have the restrictions

$${}^\alpha \Lambda_{\beta} = \frac{1}{2} {}^{\mu} \Lambda_{\nu} \sigma^{\mu\nu\alpha}{}_{\beta}, \quad \dot{\alpha} \Lambda_{\dot{\beta}} = \frac{1}{2} {}^{\mu} \Lambda_{\nu} \sigma^{\mu\nu\dot{\alpha}}{}_{\dot{\beta}}, \quad {}^\alpha \Lambda_{\dot{\alpha}} = 0, \quad \dot{\alpha} \Lambda_{\alpha} = 0, \quad (25)$$

$${}^{\mu} \Lambda_{\alpha} = \sigma^{\mu}{}_{\alpha\dot{\alpha}} \Xi^{\dot{\alpha}}, \quad {}^{\mu} \Lambda_{\dot{\alpha}} = \sigma^{\mu}{}_{\alpha\dot{\alpha}} \Xi^{\alpha}, \quad {}^\alpha \Lambda_{\mu} = -\frac{1}{k^2} \sigma_{\mu}{}^{\alpha}{}_{\dot{\alpha}} \Xi^{\dot{\alpha}}, \quad \dot{\alpha} \Lambda_{\mu} = \frac{1}{k^2} \sigma_{\mu\alpha}{}^{\dot{\alpha}} \Xi^{\alpha}. \quad (26)$$

Thus the only independent parameters in (12) are

$${}^\mu\Lambda_\nu \text{ and } \Xi^\Lambda. \quad (27)$$

We see from (14) that with the constraints (25)–(26) the generators conjugate to ${}^\nu{}_\mu\Lambda$ are

$$J^\mu{}_\nu - \frac{1}{2}J^\alpha{}_\beta \sigma_\nu{}^{\mu\beta}{}_\alpha - \frac{1}{2}J^{\dot{\alpha}}{}_{\dot{\beta}} \sigma_\nu{}^{\mu\dot{\beta}}{}_{\dot{\alpha}}. \quad (28)$$

The generators conjugate to Ξ^μ are still P_μ , whereas the generators conjugate to Ξ^α are now⁴

$$P_\alpha + \frac{1}{2}J^{\dot{\alpha}}{}_\mu \sigma^\mu{}_{\alpha\dot{\alpha}} - \frac{1}{2k^2}J^\mu{}_{\dot{\alpha}} \sigma_{\mu\alpha}{}^{\dot{\alpha}} = P_\alpha + J^{\dot{\alpha}}{}_\mu \sigma^\mu{}_{\alpha\dot{\alpha}}, \quad (29)$$

and those conjugate to $\Xi^{\dot{\alpha}}$ are

$$P_{\dot{\alpha}} + \frac{1}{2}J^\alpha{}_\mu \sigma^\mu{}_{\alpha\dot{\alpha}} - \frac{1}{2k^2}J^\mu{}_\alpha \sigma_{\mu\dot{\alpha}}{}^\alpha = P_{\dot{\alpha}} + J^\alpha{}_\mu \sigma^\mu{}_{\alpha\dot{\alpha}}. \quad (30)$$

We now show that, in the limit $k \rightarrow \infty$, (27)–(30) (plus P_μ) are the parameters and generators of the super Poincaré group. It might appear that we obtain different generators depending on whether we take the $k \rightarrow \infty$ limit of the left-hand sides or right-hand sides of (29) and (30), but this is not the case: the two J -terms on the left-hand sides have a factor of one half because each represents the same independent generator of $\text{OSp}(3,1|4)$ which is double counted in the summation in (14); if we take $k \rightarrow \infty$ on the left-hand sides then we must drop the factor of one half in the remaining J -term since it is then no longer double counted. Thus the right-hand sides of (29) and (30) represent the correct $k \rightarrow \infty$ limit of these generators.

The $k \rightarrow \infty$ limit of the subset (12) of $\text{IOSp}(3,1|4)$ transformations thus has group parameters

$${}^\mu\Lambda_\nu =: \lambda^\mu{}_\nu, \quad \Xi^\alpha =: \xi^\alpha, \quad \Xi^{\dot{\alpha}} =: \xi^{\dot{\alpha}}, \quad \text{and} \quad \Xi^\mu =: a^\mu \quad (31)$$

with, respectively, group generators

$$J^\mu{}_\nu - \frac{1}{2}J^\alpha{}_\beta \sigma_\nu{}^{\mu\beta}{}_\alpha - \frac{1}{2}J^{\dot{\alpha}}{}_{\dot{\beta}} \sigma_\nu{}^{\mu\dot{\beta}}{}_{\dot{\alpha}} =: -\mathcal{J}^\mu{}_\nu, \quad (32)$$

$$P_\alpha + J^{\dot{\alpha}}{}_\mu \sigma^\mu{}_{\alpha\dot{\alpha}} =: -iQ_\alpha, \quad (33)$$

$$P_{\dot{\alpha}} + J^\alpha{}_\mu \sigma^\mu{}_{\alpha\dot{\alpha}} =: -iQ_{\dot{\alpha}}, \quad (34)$$

$$\text{and } P_\mu. \quad (35)$$

It is now a straightforward, though tedious, matter to calculate the algebra of the new generators using (22)–(24) and (2): from the definitions (32)–(34) we obtain, in the limit $k \rightarrow \infty$,

$$\begin{aligned} [P_\mu, P_\nu] &= 0 = [P_\mu, Q_\alpha] = [P_\mu, Q_{\dot{\alpha}}], \\ \{Q_\alpha, Q_\beta\} &= 0 = \{Q_{\dot{\alpha}}, Q_{\dot{\beta}}\}, \\ \{Q_\alpha, Q_{\dot{\alpha}}\} &= 2i\sigma^\mu{}_{\alpha\dot{\alpha}} P_\mu, \end{aligned}$$

⁴Recall that indices are raised and lowered by ${}^\Lambda H^\Pi$ and ${}_\Lambda H_\Pi$, so that $J^\mu{}_{\dot{\alpha}} = -J_{\dot{\alpha}}{}^\mu = -J^{\Lambda\mu}{}_\Lambda H_{\dot{\alpha}} = k^2 J^{\dot{\beta}\mu}{}_\beta \varepsilon_{\dot{\beta}\dot{\alpha}}$; but the definition of the Infeld-van der Waerden symbols requires us to raise and lower the spinor indices of these objects with ε s, e.g. $\sigma^\mu{}_{\alpha\dot{\beta}} = -\sigma_{\mu\alpha}{}^{\dot{\alpha}} \varepsilon_{\dot{\beta}\dot{\alpha}}$.

$$\begin{aligned}
[Q_\alpha, \mathcal{J}^\mu_\nu] &= i\sigma^\mu_{\nu\alpha}{}^\beta Q_\beta, \\
[Q_{\dot{\alpha}}, \mathcal{J}^\mu_\nu] &= i\sigma^\mu_{\nu\dot{\alpha}}{}^{\dot{\beta}} Q_{\dot{\beta}}, \\
[\mathcal{J}^\mu_\nu, \mathcal{J}^\lambda_\rho] &= i(-\eta^\lambda_\nu \mathcal{J}^\mu_\rho + \eta_{\nu\rho} \mathcal{J}^{\mu\lambda} - \eta^\mu_\rho \mathcal{J}^\lambda_\nu + \eta^{\mu\lambda} \mathcal{J}_{\nu\rho}), \\
[\mathcal{J}^\mu_\nu, P_\lambda] &= i(\eta^\mu_\lambda P_\nu - \eta_{\nu\lambda} P^\mu), \\
[P_\mu, P_\nu] &= 0.
\end{aligned}$$

This is the super Poincaré algebra, as expected. The infinitesimal group element is, from (14), (25)–(26) and (32)–(34),

$$G(1 + \lambda, \Xi) = 1 + \frac{i}{2} \mathcal{J}^\mu_\nu \lambda_\mu{}^\nu - iP_\mu a^\mu - Q_\alpha \xi^\alpha - Q_{\dot{\alpha}} \xi^{\dot{\alpha}}.$$

4 IOSp(3, 1|4) gauge theory

Equipped with the super Lie algebra (22)–(24) we can now construct the gauge theory of IOSp(3, 1|4). The non-super analogue of this is Poincaré gauge theory, and the corresponding formulae of IOSp(3, 1|4) gauge theory will differ only through index-dependent factors of minus one.

We consider the super fibre bundle whose base space is superspace and whose typical fibre is IOSp(3, 1|4). An affine connection on this bundle has a pull-back A which takes the form (see (14))

$$A = A_\Lambda dZ^\Lambda = \left(\frac{i}{2} (-1)^A J^A{}_B A^B{}_{A\Lambda} - iP_A {}^A E_\Lambda \right) dZ^\Lambda, \quad (36)$$

where ${}^A E_\Lambda$ is the superspace vielbein which relates the coordinate frame to the orthosymplectic frame. Under an infinitesimal gauge transformation A transforms to

$$A' = G(1 + \Lambda, \Xi)^{-1} A G(1 + \Lambda, \Xi) + G(1 + \Lambda, \Xi)^{-1} dG(1 + \Lambda, \Xi)$$

$$[(14)] = A - \frac{i}{2} [J^A{}_B, A] {}^B A_\Lambda + i[P_A, A] \Xi^A + \frac{i}{2} J^A{}_B d^B A_\Lambda - iP_A d\Xi^A.$$

We use (36) and (22)–(24) to evaluate the supercommutators and find

$$\begin{aligned}
A' &= A + \frac{i}{2} J^A{}_D A^D{}_{B\Lambda} dZ^\Lambda {}^B A_\Lambda - \frac{i}{2} (-1)^{C(1+A)+B(A+C)} J^C{}_B A^A{}_{C\Lambda} dZ^\Lambda {}^B A_\Lambda \\
&\quad + i(-1)^{BA} P_B {}^A E_\Lambda dZ^\Lambda {}^B A_\Lambda - iP_C A^C{}_{A\Lambda} dZ^\Lambda \Xi^A + \frac{i}{2} J^A{}_B d^B A_\Lambda - iP_A d\Xi^A \\
&=: A + \delta A.
\end{aligned}$$

Comparing this with (36) we see that the component parts of the gauge potential have the gauge transformations

$$\delta {}^A E_\Lambda = \Xi^A{}_{,\Lambda} + (-1)^{B(A+1)} \Xi^B A^A{}_{B\Lambda} - {}^A \Lambda_B {}^B E_\Lambda, \quad (37)$$

$$\delta A^A{}_{B\Lambda} = {}^A \Lambda_{B,\Lambda} + (-1)^{(C+B)(A+C)} {}^C \Lambda_B A^A{}_{C\Lambda} - {}^A \Lambda_C A^C{}_{B\Lambda}, \quad (38)$$

The commutator of two small gauge transformations $\delta(\Lambda_1, \Xi_1)$ and $\delta(\Lambda_2, \Xi_2)$ of A reads, to leading order,

$$[\delta(\Lambda_1, \Xi_1), \delta(\Lambda_2, \Xi_2)] A = \delta(\Lambda', \Xi') A,$$

where

$$\begin{aligned} {}^A\Lambda'_B &= {}^A\Lambda_{1C} {}^C\Lambda_{2B} - {}^A\Lambda_{2C} {}^C\Lambda_{1B}, \\ \Xi'^A &= {}^A\Lambda_{1B} \Xi_2^B - {}^A\Lambda_{2B} \Xi_1^B. \end{aligned}$$

The pull-back of the gauge curvature is

$$\begin{aligned} R &= \frac{1}{2}(-1)^{\Lambda\Pi} R_{\Lambda\Pi} dZ^\Lambda \wedge dZ^\Pi \\ &= \frac{1}{2}(A_{\Pi,\Lambda} - (-1)^{\Lambda\Pi} A_{\Lambda,\Pi} + (-1)^{\Lambda\Pi} A_\Lambda A_\Pi - A_\Pi A_\Lambda) dZ^\Lambda \wedge dZ^\Pi \\ &= \left(\frac{i}{4}(-1)^{B+\Lambda\Pi} J^B{}_A R^A{}_{B\Lambda\Pi} - \frac{i}{2}(-1)^{\Lambda\Pi} P_A R^A{}_{\Lambda\Pi} \right) dZ^\Lambda \wedge dZ^\Pi, \end{aligned}$$

where

$$R^A{}_{\Lambda\Pi} = (-1)^{\Lambda\Pi} {}^A E_{\Pi,\Lambda} - {}^A E_{\Lambda,\Pi} + (-1)^{\Lambda B} A^A{}_{B\Lambda} {}^B E_\Pi - (-1)^{\Pi(B+\Lambda)} A^A{}_{B\Pi} {}^B E_\Lambda, \quad (39)$$

$$\begin{aligned} R^A{}_{B\Lambda\Pi} &= (-1)^{\Lambda\Pi} A^A{}_{B\Pi,\Lambda} - A^A{}_{B\Lambda,\Pi} + (-1)^{\Lambda(C+B)} A^A{}_{C\Lambda} A^C{}_{B\Pi} \\ &\quad - (-1)^{\Pi(C+B+\Lambda)} A^A{}_{C\Pi} A^C{}_{B\Lambda}. \end{aligned} \quad (40)$$

We recognise $R^A{}_{\Lambda\Pi}$ as the torsion supertensor and $R^A{}_{B\Lambda\Pi}$ as the Riemann supertensor:

$$R^A{}_{BC} = (-1)^{(\Lambda+B)\Pi} R^A{}_{\Lambda\Pi} {}^\Lambda E_B {}^\Pi E_C, \quad (41)$$

$$R^A{}_{BCD} = (-1)^{(\Lambda+C)\Pi} R^A{}_{B\Lambda\Pi} {}^\Lambda E_C {}^\Pi E_D. \quad (42)$$

The gauge curvatures satisfy the following Bianchi identities:

$$\begin{aligned} R^A{}_{\Lambda\Pi|\Sigma} + (-1)^{\Lambda(\Pi+\Sigma)} R^A{}_{\Pi\Sigma|\Lambda} + (-1)^{\Sigma(\Lambda+\Pi)} R^A{}_{\Sigma\Lambda|\Pi} &= (-1)^{(B+\Lambda)(\Pi+\Sigma)} R^A{}_{B\Pi\Sigma} {}^B E_\Lambda \\ &\quad + (-1)^{B(\Sigma+\Lambda)+\Sigma(\Lambda+\Pi)} R^A{}_{B\Sigma\Lambda} {}^B E_\Pi + (-1)^{B(\Lambda+\Pi)} R^A{}_{B\Lambda\Pi} {}^B E_\Sigma, \end{aligned} \quad (43)$$

$$R^A{}_{B\Lambda\Pi|\Sigma} + (-1)^{\Lambda(\Pi+\Sigma)} R^A{}_{B\Pi\Sigma|\Lambda} + (-1)^{\Sigma(\Lambda+\Pi)} R^A{}_{B\Sigma\Lambda|\Pi} = 0, \quad (44)$$

where $|$ denotes a covariant derivative with connection $A^A{}_{B\Lambda}$ that acts only on orthosymplectic-frame indices.

We could now go on to construct gauge supersymmetry, the superspace analogue of general relativity. The procedure corresponds exactly to the formulation of general relativity as Poincaré gauge theory. One takes the superspace version of the Hilbert action and finds that it is gauge invariant if and only if the torsion supertensor vanishes, which is in turn the requirement that the $\text{OSp}(3,1|4)$ potential $A^A{}_{B\Lambda}$ satisfy its equation of motion. However, as discussed in the next section, the dynamics of gauge supersymmetry has no relevance for our derivation of supergravity. All of the details of $\text{IOSp}(3,1|4)$ gauge theory necessary to obtain supergravity have been presented so we now turn to the derivation.

5 Supergravity

Section 3 showed how to obtain the super Poincaré algebra from the super Lie algebra of $\text{IOSp}(3,1|4)$. We now use this method to turn the gauge potential and curvature of

IOSp(3, 1|4) gauge theory into super-Poincaré-algebra valued quantities. The result is a blend of gauge theory and the super Poincaré group, both of which are highly significant from a physical standpoint, so we may hope to obtain some supersymmetric physics.

In curved superspace the canonical metric (2) with coordinates (1) cannot, in general, be introduced in an extended region. The corresponding objects in curved superspace are the orthosymplectic frame E_A and the orthosymplectic-frame components of the metric, which we choose in line with (1) and (2):

$$[E_m] = 1, \quad [E_a] = [E_{\dot{a}}] = \frac{1}{2}, \quad (45)$$

$${}^A H_B = \begin{pmatrix} \eta_{mn} & 0 & 0 \\ 0 & k^2 \varepsilon_{ab} & 0 \\ 0 & 0 & -k^2 \varepsilon_{\dot{a}\dot{b}} \end{pmatrix} \Rightarrow {}^A H^B = \begin{pmatrix} \eta^{mn} & 0 & 0 \\ 0 & -\frac{1}{k^2} \varepsilon^{ab} & 0 \\ 0 & 0 & \frac{1}{k^2} \varepsilon^{\dot{a}\dot{b}} \end{pmatrix}. \quad (46)$$

We choose the superspace coordinates Z^Λ to all have the same, standard, units

$$[Z^\Lambda] = -1 \quad \forall \quad \Lambda \quad (47)$$

so that the metric coordinate components

$${}_\Lambda G_\Pi = {}_\Lambda E^A {}_A H_B {}^B E_\Pi$$

are dimensionless. In flat superspace the frame E_A may also be taken as a coordinate frame and the simplest choice of vielbein (now a coordinate transformation) is then, in view of (45)–(47),

$${}^A E_\Lambda = \begin{pmatrix} {}^m \delta_\mu & 0 & 0 \\ 0 & \frac{1}{k} {}^a \delta_\alpha & 0 \\ 0 & 0 & \frac{1}{k} {}^{\dot{a}} \delta_{\dot{\alpha}} \end{pmatrix}. \quad (48)$$

We therefore expect to be able to choose a vielbein that reduces to (48) when superspace is flat.

Following the path we have set out, we must extract the part of the IOSp(3, 1|4) gauge potential corresponding to the infinitesimal transformations (12) and take $k \rightarrow \infty$. Since we wish to discover what happens to the IOSp(3, 1|4) gauge transformations in this super-Poincaré limit, we must perform the same operation on the parameters of infinitesimal gauge transformations. Thus, from (12), the translation potentials ${}^A E_\Lambda$ and parameters Ξ^A are to remain independent, whereas we impose the following constraints on the OSp(3, 1|4) potentials $A^A_{B\Lambda}$ and parameters ${}^A \Lambda_B$:

$$A^A_{B\Lambda} = \begin{pmatrix} A^m_{n\Lambda} & \sigma^m_{b\dot{a}} {}^{\dot{a}} E_\Lambda & \sigma^m_{ab} {}^a E_\Lambda \\ -\frac{1}{k^2} \sigma_n{}^a{}_{\dot{a}} {}^{\dot{a}} E_\Lambda & \frac{1}{2} A^m_{n\Lambda} \sigma_m{}^{na}{}_b & 0 \\ \frac{1}{k^2} \sigma_{na}{}^{\dot{a}} {}^a E_\Lambda & 0 & \frac{1}{2} A^m_{n\Lambda} \sigma_m{}^{n\dot{a}}{}_{\dot{b}} \end{pmatrix}, \quad (49)$$

$${}^A \Lambda_B = \begin{pmatrix} {}^m \Lambda_n & \sigma^m_{b\dot{a}} \Xi^{\dot{a}} & \sigma^m_{ab} \Xi^a \\ -\frac{1}{k^2} \sigma_n{}^a{}_{\dot{a}} \Xi^{\dot{a}} & \frac{1}{2} {}^m \Lambda_n \sigma_m{}^{na}{}_b & 0 \\ \frac{1}{k^2} \sigma_{na}{}^{\dot{a}} \Xi^{\dot{a}} & 0 & \frac{1}{2} {}^m \Lambda_n \sigma_m{}^{n\dot{a}}{}_{\dot{b}} \end{pmatrix}. \quad (50)$$

Note that ${}^A \Lambda_B$ and Ξ^A are the parameters of a gauge transformation and are therefore functions of the superspace coordinates Z^Λ . From the derivation in Section 3 we know

that, with the constraints (49) and the limit $k \rightarrow \infty$, the gauge potential (36) is super-Poincaré-algebra valued. Imposing (49) and (50) on the gauge transformation (37) of ${}^A E_\Lambda$ gives

$$\delta {}^m E_\Lambda = \Xi^m{}_{,\Lambda} + \Xi^n A^m{}_{n\Lambda} - 2\sigma^m{}_{a\dot{a}}(\Xi^a \dot{a} E_\Lambda + \Xi^{\dot{a}} {}^a E_\Lambda) - {}^m \Lambda_n {}^n E_\Lambda, \quad (51)$$

$$\begin{aligned} \delta {}^a E_\Lambda &= \Xi^a{}_{,\Lambda} + \frac{1}{2} \Xi^b A^m{}_{n\Lambda} \sigma_m{}^{na}{}_b - \frac{1}{k^2} \sigma_m{}^a{}_{\dot{a}} \Xi^m \dot{a} E_\Lambda \\ &\quad - \frac{1}{2} {}^m \Lambda_n \sigma_m{}^{na}{}_b {}^b E_\Lambda + \frac{1}{k^2} \sigma_m{}^a{}_{\dot{a}} \Xi^{\dot{a}} {}^m E_\Lambda, \end{aligned} \quad (52)$$

$$\begin{aligned} \delta \dot{a} E_\Lambda &= \Xi^{\dot{a}}{}_{,\Lambda} + \frac{1}{2} \Xi^{\dot{b}} A^m{}_{n\Lambda} \sigma_m{}^{n\dot{a}}{}_{\dot{b}} + \frac{1}{k^2} \sigma_{m\dot{a}}{}^{\dot{a}} \Xi^m {}^a E_\Lambda \\ &\quad - \frac{1}{2} {}^m \Lambda_n \sigma_m{}^{n\dot{a}}{}_{\dot{b}} {}^{\dot{b}} E_\Lambda - \frac{1}{k^2} \sigma_{m\dot{a}}{}^{\dot{a}} \Xi^a {}^m E_\Lambda. \end{aligned} \quad (53)$$

On the other hand, with the constraints (49) the only independent $A^A{}_{B\Lambda}$ are now the 6×8 independent $A^m{}_{n\Lambda}$. The gauge transformation of $A^m{}_{n\Lambda}$ is obtained from (38) with (49) and (50):

$$\begin{aligned} \delta A^m{}_{n\Lambda} &= {}^m \Lambda_{n,\Lambda} + {}^r \Lambda_n A^m{}_{r\Lambda} - \frac{1}{k^2} \sigma_n{}^a{}_{\dot{a}} \Xi^{\dot{a}} \sigma^m{}_{ab} {}^b E_\Lambda + \frac{1}{k^2} \sigma_{na}{}^{\dot{a}} \Xi^a \sigma_{mb\dot{a}} {}^b E_\Lambda \\ &\quad - {}^m \Lambda_r A^r{}_{n\Lambda} - \frac{1}{k^2} \sigma^m{}_{a\dot{a}} \Xi^{\dot{a}} \sigma_n{}^a{}_{\dot{b}} {}^{\dot{b}} E_\Lambda + \frac{1}{k^2} \sigma^m{}_{a\dot{a}} \Xi^a \sigma_{nb}{}^{\dot{a}} {}^{\dot{b}} E_\Lambda. \end{aligned} \quad (54)$$

In light of (48), it is clear from (51)–(53) that we may choose the order Z^α and $Z^{\dot{\alpha}}$ terms of Ξ^A so that a Ξ -transformation of ${}^A E_\Lambda$ leaves its $Z^\alpha, Z^{\dot{\alpha}}$ -independent part in the form

$${}^A E_\Lambda(Z^\alpha = Z^{\dot{\alpha}} = 0) = \begin{pmatrix} e^m{}_\mu & 0 & 0 \\ \frac{1}{2} \phi^a{}_\mu & \frac{1}{k} {}^a \delta_\alpha & 0 \\ \frac{1}{2} \dot{\phi}^{\dot{a}}{}_\mu & 0 & \frac{1}{k} \dot{a} \delta_{\dot{\alpha}} \end{pmatrix}, \quad (55)$$

where $e^m{}_\mu$, $\phi^a{}_\mu$ and $\dot{\phi}^{\dot{a}}{}_\mu$ are functions of Z^μ only ($\phi^a{}_\mu$ and $\dot{\phi}^{\dot{a}}{}_\mu$ a -type). Similarly, (54) shows that we may choose the order Z^α and $Z^{\dot{\alpha}}$ terms of ${}^m \Lambda_n$ so that a Λ -transformation of $A^m{}_{n\Lambda}$ results in

$$A^m{}_{n\alpha}(Z^\alpha = Z^{\dot{\alpha}} = 0) = A^m{}_{n\dot{\alpha}}(Z^\alpha = Z^{\dot{\alpha}} = 0) = 0. \quad (56)$$

The standard choices (55) and (56) amount to a partial gauge fixing.

We identify the bosonic sector of superspace with space-time. Note that this is the largest sector of superspace for which a metric can be retained in the $k \rightarrow \infty$ limit. We see this both from (46) and (55); the latter shows that the matrix ${}^A E_\Lambda$ does not have an inverse when $k \rightarrow \infty$ since its body doesn't [7], but we retain a tetrad ${}^m E_\mu$ in space-time. We must then take account of the experimental fact that physical fields such as the tetrad show no dependence on the a -type coordinates. Accordingly, we must take as the physical fields in space-time

$${}^A E_\Lambda| \quad \text{and} \quad A^m{}_{n\Lambda}|,$$

where $|$ means “ $Z^\alpha = Z^{\dot{\alpha}} = 0$ and $k \rightarrow \infty$ ”.⁵ Then with the partial gauge fixing (55) and (56) we get a physical field content of

$${}^A E_\Lambda | = \begin{pmatrix} e^m{}_\mu & 0 & 0 \\ \frac{1}{2}\phi^a{}_\mu & 0 & 0 \\ \frac{1}{2}\phi^{\dot{a}}{}_\mu & 0 & 0 \end{pmatrix}, \quad (57)$$

$$A^m{}_{n\mu} | =: \Gamma^m{}_{n\mu}(Z^\mu), \quad A^m{}_{n\dot{\alpha}} | = A^m{}_{n\dot{\alpha}} | = 0. \quad (58)$$

Note from (51)–(54) that the transformation of the physical fields is determined solely by

$$\Xi^m | =: \varepsilon^m, \quad \Xi^a | =: -\xi^a, \quad \Xi^{\dot{a}} | =: -\xi^{\dot{a}}, \quad (59)$$

$${}^m \Lambda_n | =: \lambda^m{}_n, \quad (60)$$

so that (59)–(60) are the physically relevant gauge parameters. With (57)–(60), the $k \rightarrow \infty$ limit of the transformations (51)–(54) is

$$\delta e^m{}_\mu = \varepsilon^m{}_{,\mu} + \varepsilon^n \Gamma^m{}_{n\mu} + \sigma^m{}_{a\dot{a}} (\xi^a \phi^{\dot{a}}{}_\mu + \xi^{\dot{a}} \phi^a{}_\mu) - \lambda^m{}_n e^n{}_\mu, \quad (61)$$

$$\delta \phi^a{}_\mu = -2D_\mu \xi^a - \frac{1}{2} \lambda^m{}_n \sigma_m{}^{na} \phi^b{}_\mu, \quad (62)$$

$$\delta \phi^{\dot{a}}{}_\mu = -2D_\mu \xi^{\dot{a}} - \frac{1}{2} \lambda^m{}_n \sigma_m{}^{n\dot{a}} \phi^{\dot{b}}{}_\mu, \quad (63)$$

$$\delta \Gamma^m{}_{n\mu} = \lambda^m{}_{n,\mu} + \lambda^r{}_n \Gamma^m{}_{r\mu} - \lambda^m{}_r \Gamma^r{}_{n\mu}, \quad (64)$$

where D_μ is the familiar space-time covariant derivative operator with connection $\Gamma^m{}_{n\mu}$ that acts only on orthonormal-frame and spinor indices. Eqns. (61)–(64) are the super Poincaré gauge transformations of supergravity [13], with gravitino

$$\psi_\mu{}^a = \frac{1}{\kappa} \phi^a{}_\mu, \quad \kappa = \sqrt{8\pi G}. \quad (65)$$

Turning now to the gauge curvatures, we impose the constraints (49) on the torsion (39) and take $k \rightarrow \infty$, obtaining

$$R^m{}_{\Lambda\Pi} = (-1)^{\Lambda\Pi} {}^m E_{\Pi,\Lambda} - {}^m E_{\Lambda,\Pi} + A^m{}_{n\Lambda} {}^n E_\Pi - (-1)^{\Pi\Lambda} A^m{}_{n\Pi} {}^n E_\Lambda + 2(-1)^\Lambda \sigma^m{}_{a\dot{a}} (\dot{a} E_\Lambda {}^a E_\Pi + {}^a E_\Lambda \dot{a} E_\Pi), \quad (66)$$

$$R^a{}_{\Lambda\Pi} = (-1)^{\Lambda\Pi} {}^a E_{\Pi,\Lambda} - {}^a E_{\Lambda,\Pi} + \frac{1}{2} (-1)^\Lambda A^m{}_{n\Lambda} \sigma_m{}^{na} {}^b E_\Pi - \frac{1}{2} (-1)^{\Pi(1+\Lambda)} A^m{}_{n\Pi} \sigma_m{}^{na} {}^b E_\Lambda, \quad (67)$$

$$R^{\dot{a}}{}_{\Lambda\Pi} = (-1)^{\Lambda\Pi} \dot{a} E_{\Pi,\Lambda} - \dot{a} E_{\Lambda,\Pi} + \frac{1}{2} (-1)^\Lambda A^m{}_{n\Lambda} \sigma_m{}^{n\dot{a}} {}^{\dot{b}} E_\Pi - \frac{1}{2} (-1)^{\Pi(1+\Lambda)} A^m{}_{n\Pi} \sigma_m{}^{n\dot{a}} {}^{\dot{b}} E_\Lambda, \quad (68)$$

⁵The standard procedure of “gauge completion”, used in both the Nath–Arnowitt and Wess–Zumino formulations, can be used to give ${}^A E_\Lambda$ and $A^m{}_{n\Lambda}$ a complete expansion in the a -type coordinates when auxiliary fields are added to the theory. [2, 12, 4]

while the same operation on the Riemann supertensor (40) produces

$$R^m{}_{n\Lambda\Pi} = (-1)^{\Lambda\Pi} A^m{}_{n\Pi,\Lambda} - A^m{}_{n\Lambda,\Pi} + A^m{}_{r\Lambda} A^r{}_{n\Pi} - (-1)^{\Pi\Lambda} A^m{}_{r\Pi} A^r{}_{n\Lambda}, \quad (69)$$

$$\begin{aligned} R^m{}_{a\Lambda\Pi} &= (-1)^{\Lambda\Pi} \sigma^m{}_{a\dot{a}} \dot{a}E_{\Pi,\Lambda} - \sigma^m{}_{a\dot{a}} \dot{a}E_{\Lambda,\Pi} + (-1)^\Lambda A^m{}_{n\Lambda} \sigma^n{}_{a\dot{a}} \dot{a}E_\Pi \\ &\quad + \frac{1}{2} \sigma^m{}_{b\dot{a}} \dot{a}E_\Lambda A^n{}_{r\Pi} \sigma_n{}^{rb}{}_a - (-1)^{\Pi(1+\Lambda)} A^m{}_{n\Pi} \sigma^n{}_{a\dot{a}} \dot{a}E_\Lambda \\ &\quad - \frac{1}{2} (-1)^{\Pi\Lambda} \sigma^m{}_{b\dot{a}} \dot{a}E_\Pi A^n{}_{r\Lambda} \sigma_n{}^{rb}{}_a, \end{aligned} \quad (70)$$

$$\begin{aligned} R^m{}_{\dot{a}\Lambda\Pi} &= (-1)^{\Lambda\Pi} \sigma^m{}_{a\dot{a}} {}^aE_{\Pi,\Lambda} - \sigma^m{}_{a\dot{a}} {}^aE_{\Lambda,\Pi} + (-1)^\Lambda A^m{}_{n\Lambda} \sigma^n{}_{a\dot{a}} {}^aE_\Pi \\ &\quad + \frac{1}{2} \sigma^m{}_{a\dot{b}} {}^aE_\Lambda A^n{}_{r\Pi} \sigma_n{}^{r\dot{b}}{}_{\dot{a}} - (-1)^{\Pi(1+\Lambda)} A^m{}_{n\Pi} \sigma^n{}_{a\dot{a}} {}^aE_\Lambda \\ &\quad - \frac{1}{2} (-1)^{\Pi\Lambda} \sigma^m{}_{a\dot{b}} {}^aE_\Pi A^n{}_{r\Lambda} \sigma_n{}^{r\dot{b}}{}_{\dot{a}}, \end{aligned} \quad (71)$$

$$\begin{aligned} R^a{}_{b\Lambda\Pi} &= \frac{1}{2} (-1)^{\Lambda\Pi} A^m{}_{n\Pi,\Lambda} \sigma_m{}^{na}{}_b - \frac{1}{2} (-1)^{\Lambda\Pi} A^m{}_{n\Lambda,\Pi} \sigma_m{}^{na}{}_b \\ &\quad + \frac{1}{4} A^m{}_{n\Lambda} \sigma_m{}^{na}{}_c A^r{}_{s\Pi} \sigma_r{}^{sc}{}_b - \frac{1}{4} (-1)^{\Lambda\Pi} A^m{}_{n\Pi} \sigma_m{}^{na}{}_c A^r{}_{s\Lambda} \sigma_r{}^{sc}{}_b, \end{aligned} \quad (72)$$

$$\begin{aligned} R^{\dot{a}}{}_{b\Lambda\Pi} &= \frac{1}{2} (-1)^{\Lambda\Pi} A^m{}_{n\Pi,\Lambda} \sigma_m{}^{n\dot{a}}{}_b - \frac{1}{2} (-1)^{\Lambda\Pi} A^m{}_{n\Lambda,\Pi} \sigma_m{}^{n\dot{a}}{}_b \\ &\quad + \frac{1}{4} A^m{}_{n\Lambda} \sigma_m{}^{n\dot{a}}{}_c A^r{}_{s\Pi} \sigma_r{}^{s\dot{c}}{}_b - \frac{1}{4} (-1)^{\Lambda\Pi} A^m{}_{n\Pi} \sigma_m{}^{n\dot{a}}{}_c A^r{}_{s\Lambda} \sigma_r{}^{s\dot{c}}{}_b, \end{aligned} \quad (73)$$

$$R^a{}_{m\Lambda\Pi} = R^{\dot{a}}{}_{m\Lambda\Pi} = 0. \quad (74)$$

We introduce a strange covariant derivative operator $\overleftarrow{\mathcal{D}}_\Lambda$ on superspace that acts only on orthosymplectic-frame indices m, a, \dot{a} with a *Lorentz* connection $A^m{}_{n\Lambda}$. Then $\overleftarrow{\mathcal{D}}_\mu$ is the space-time covariant derivative operator D_μ (acting on the right). We have

$$\sigma^m{}_{a\dot{a}} \overleftarrow{\mathcal{D}}_\Lambda = 0 \quad (75)$$

since

$$\begin{aligned} \sigma^m{}_{a\dot{a}} \overleftarrow{\mathcal{D}}_\Lambda &= \sigma^n{}_{a\dot{a}} A^m{}_{n\Lambda} - \frac{1}{2} \sigma^m{}_{b\dot{a}} A^n{}_{r\Lambda} \sigma_n{}^{rb}{}_a - \frac{1}{2} \sigma^m{}_{a\dot{b}} A^n{}_{r\Lambda} \sigma_n{}^{r\dot{b}}{}_{\dot{a}} \\ &= \sigma^n{}_{a\dot{a}} A^m{}_{n\Lambda} - \frac{1}{2} (\sigma^m{}_{b\dot{a}} \sigma_n{}^{rb}{}_a + \sigma^m{}_{a\dot{b}} \sigma_n{}^{r\dot{b}}{}_{\dot{a}}) A^n{}_{r\Lambda} \\ &= \sigma^n{}_{a\dot{a}} A^m{}_{n\Lambda} - \eta^{m[n} \sigma^r]{}_{a\dot{a}} A_{nr\Lambda} = 0. \end{aligned}$$

By making use of (75) we find a simple relation between (70)–(71) and (67)–(68) so that the new Riemann supertensor (69)–(74) can be expressed as

$$R^m{}_{n\Lambda\Pi} = (-1)^{\Lambda\Pi} A^m{}_{n\Pi,\Lambda} - A^m{}_{n\Lambda,\Pi} + A^m{}_{r\Lambda} A^r{}_{n\Pi} - (-1)^{\Lambda\Pi} A^m{}_{r\Pi} A^r{}_{n\Lambda}, \quad (76)$$

$$R^m{}_{a\Lambda\Pi} = \sigma^m{}_{a\dot{a}} R^{\dot{a}}{}_{\Lambda\Pi}, \quad R^m{}_{\dot{a}\Lambda\Pi} = \sigma^m{}_{a\dot{a}} R^a{}_{\Lambda\Pi}, \quad R^{\dot{a}}{}_{m\Lambda\Pi} = R^{\dot{a}}{}_{m\Lambda\Pi} = 0, \quad (77)$$

$$R^a{}_{b\Lambda\Pi} = \frac{1}{2} R^m{}_{n\Lambda\Pi} \sigma_m{}^{na}{}_b, \quad R^{\dot{a}}{}_{b\Lambda\Pi} = \frac{1}{2} R^m{}_{n\Lambda\Pi} \sigma_m{}^{n\dot{a}}{}_b. \quad (78)$$

Comparing this with the $k \rightarrow \infty$ limit of (49) we see that the various parts of the new gauge curvature are entirely analogous to the corresponding parts of the new gauge potential. This means that, just like the new gauge potential, the new gauge curvature is super-Poincaré-algebra valued.

Using (57) and (58) we obtain from the torsion (66)–(68) and Riemann supertensor (76)–(78)

$$R^m{}_{\mu\nu}| = D_\mu e^m{}_\nu - D_\nu e^m{}_\mu + \frac{1}{2}\sigma^m{}_{a\dot{a}}(\phi^{\dot{a}}{}_\mu \phi^a{}_\nu + \psi^a{}_\mu \psi^{\dot{a}}{}_\nu) \quad (79)$$

$$R^a{}_{\mu\nu}| = D_{[\mu}\phi^a{}_{\nu]}, \quad R^{\dot{a}}{}_{\mu\nu}| = D_{[\mu}\phi^{\dot{a}}{}_{\nu]}, \quad (80)$$

$$R^m{}_{n\mu\nu}| = \mathcal{R}^m{}_{n\mu\nu}, \quad R^m{}_{a\mu\nu}| = \sigma^m{}_{a\dot{a}} R^{\dot{a}}{}_{\mu\nu}|, \quad R^m{}_{\dot{a}\mu\nu}| = \sigma^m{}_{a\dot{a}} R^a{}_{\mu\nu}|, \quad (81)$$

$$R^a{}_{b\mu\nu}| = \frac{1}{2}\mathcal{R}^m{}_{n\mu\nu} \sigma_m{}^{na}{}_{b}, \quad R^{\dot{a}}{}_{b\mu\nu}| = \frac{1}{2}\mathcal{R}^m{}_{n\mu\nu} \sigma_m{}^{n\dot{a}}{}_{b}, \quad (82)$$

where $\mathcal{R}^m{}_{n\mu\nu}$ is the space-time Riemann tensor. The Bianchi identities (43) and (44) are still satisfied by (79)–(82) and (57).⁶

It remains to specify the dynamics of the theory defined by the gauge transformations (61)–(64) and curvatures (79)–(82); we shall do this by writing suitable equations of motion. The requirement for the equations of motion is not that they be covariant under the super-Poincaré gauge transformations (61)–(64), but that they be gauge covariant when $\Gamma^m{}_{n\mu}$ is on-shell. Although we have obtained this theory from IOSp(3, 1|4) gauge theory, it is fruitless to try to obtain suitable field equations by the procedure of imposing our constraints (49) and limit $k \rightarrow \infty$ on the field equations of IOSp(3, 1|4) gauge theory: the latter can be obtained from a ($\Gamma^A{}_{B\mu}$ on-shell) gauge-invariant action by varying *independently* the fields ${}^A E_\Lambda$ and $A^A{}_{B\Lambda}$, but this is inconsistent with the constraints (though not with the limit $k \rightarrow \infty$) so that this procedure would not result in super-Poincaré-covariant equations.

The appropriate equations of motion are

$$R^m{}_{\mu\nu}| = 0, \quad (83)$$

$$R^m{}_{\dot{a}\mu\nu}| e^\mu{}_m = 0, \quad (84)$$

$$\mathcal{R}_{n\nu} e^n{}_\mu e^{\nu m} - R^m{}_{\dot{a}\mu\nu}| \phi^{\dot{a}\nu} - R^m{}_{a\mu\nu}| \phi^{a\nu} = 0, \quad (85)$$

where $\mathcal{R}_{n\nu} = R^m{}_{n\mu\nu}| e^\mu{}_m$ is the space-time Ricci tensor. Using (79)–(81) in these equations, we see that (83) determines the space-time torsion $\mathcal{T}^m{}_{\mu\nu}$,

$$\mathcal{T}^m{}_{\mu\nu} = \frac{1}{2}\sigma^m{}_{a\dot{a}}(\phi^a{}_\mu \phi^{\dot{a}}{}_\nu + \phi^{\dot{a}}{}_\mu \phi^a{}_\nu), \quad (86)$$

(84) is the gravitino field equation (recall (65))

$$\sigma^m{}_{a\dot{a}} D_{[\mu}\phi^a{}_{\nu]} e^\mu{}_m = 0, \quad (87)$$

and (85) is the tetrad field equation

$$\mathcal{R}_\mu{}^m + \sigma^m{}_{a\dot{a}}(\phi^{\dot{a}\nu} D_{[\mu}\phi^a{}_{\nu]} + \phi^{a\nu} D_{[\mu}\phi^{\dot{a}}{}_{\nu]}) = 0. \quad (88)$$

⁶In the Wess–Zumino formulation the constraints on the superspace torsion mean the Bianchi identities are no longer identities; rather, they give the field equations of supergravity [3], and this is the *a posteriori* motivation for the constraints.

Transvecting (88) with e^μ_m and using (87) we obtain

$$\mathcal{R}^m_m = 0,$$

so that we may replace \mathcal{R}_μ^m in (88) by the space-time Einstein tensor $\mathcal{G}_{\mu m}$. Employing also the the following equivalent form of the gravitino field equation

$$D_{[\mu}\phi^a_{\nu]} - \frac{1}{2}i\varepsilon_{\mu\nu}{}^{\lambda\rho} D_{[\lambda}\phi^a_{\rho]} = 0$$

we can rewrite (88) as

$$\mathcal{G}^\mu_m + \frac{1}{2}i\varepsilon^{\mu\nu\lambda\rho} \sigma_{ma\dot{a}}(\phi^{\dot{a}}_\nu D_\lambda\phi^a_\rho - \phi^a_\nu D_\lambda\phi^{\dot{a}}_\rho) = 0, \quad (89)$$

which is the tetrad field equation as it emerges from the supergravity action. As discussed in [15], for example, the action giving rise to (86), (87) and (89) is invariant under (61)–(64) when (86) is satisfied. Therefore (84) and (85) are covariant under (61)–(64) when (83) is satisfied.

The geometrical view of simple $D = 4$ supergravity presented here is that, although it is not a simple theory of superspace geometry, it is related to a simple theory of superspace geometry (gauge supersymmetry) in a physically understandable way. A comparable theory is Newtonian gravity, which also has a rather complicated geometrical formulation, in terms of space-time [14]. Moreover, Newtonian gravity in its space-time form can be derived from general relativity, formulated as Poincaré gauge theory, by a method analogous to that used here to derive supergravity from $\text{IOSp}(3, 1|4)$ gauge theory [16]. In place of (1) and (2) one has Minkowski coordinates with $x^0 = t$, not $x^0 = ct$, so that $\eta_{00} = -c^2$. Instead of (3) and (7) one has infinitesimal Poincaré and Galilean transformations; one obtains the latter from the former by taking the analogous limit $c \rightarrow \infty$, but here no constraints are required. One then imposes a similar limiting procedure on the Lorentz gauge potential of Poincaré gauge theory. In curved space-time one takes a local frame with $\eta_{\widehat{0}\widehat{0}} = -c^2$ and coordinates with $[x^\mu] = -1$ (cf. (46) and (47)); a 3-space metric, but not a space-time metric, is preserved when $c \rightarrow \infty$ [14]. The crucial difference from our derivation of supergravity is the lack of constraints, which allows one to obtain a suitable Newtonian field equation by taking the $c \rightarrow \infty$ limit of the Einstein field equation.

6 Conclusions

We have obtained a new formulation of simple $D = 4$ supergravity in terms of the geometry of superspace. This formulation makes clear a relationship between supergravity and the simplest theory of superspace geometry, namely gauge supersymmetry. When one views gauge supersymmetry as $\text{IOSp}(3, 1|4)$ gauge theory, then supergravity emerges from manipulating quantities in the theory so that $\text{IOSp}(3, 1|4)$ becomes the super Poincaré group. This involves imposing constraints and taking a limit.

In thinking about this relationship between supergravity and gauge supersymmetry, it is interesting to note that a similar relationship exists between Newtonian gravity and general relativity. When one views general relativity as Poincaré gauge

theory, and when one takes $\eta_{00} = -c^2$, then Newtonian gravity in its space-time form emerges from manipulating quantities in the theory so that the Poincaré group becomes the Galilean group. This simply involves taking the limit $c \rightarrow \infty$; no constraints are required.

The only difference between the two relationships

$$\begin{aligned} \text{gauge supersymmetry} &\rightarrow \text{supergravity} \\ \text{general relativity} &\rightarrow \text{Newtonian gravity} \end{aligned}$$

as described here, is the occurrence of constraints in the former. This difference is of crucial significance for the dynamics, however; the dynamics of supergravity has no simple relation to the dynamics of gauge supersymmetry [1]. Nevertheless, one may say that in the geometrical formulation given here, simple $D = 4$ supergravity is more akin to Newtonian gravity than to general relativity.

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