

1 **Mineralisation in Layered Mafic-Ultramafic Intrusions**

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10 See Also: Economic Geology, Layered Intrusions, Igneous Rocks, Igneous Processes, Mining Geology,
11 Mantle, Mantle Plumes, Minerals: Sulphides.

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14 **1. Introduction**

15 Mineral deposits form in mafic-ultramafic layered intrusions (LMI) through a variety of magmatic
16 processes. Chromium, platinum-group elements (PGE) and nickel (Ni) are almost exclusively extracted
17 from mafic-ultramafic rocks (Table 1). LMI also host metals such as gold (Au), copper (Cu), cobalt (Co),
18 vanadium (V), titanium (Ti) and scandium (Sc). Mafic-ultramafic rocks are major sources of dimension
19 stone and aggregate, and increasingly, carbon capture and storage may generate significant demands
20 for Mg-silicate (olivine and associated minerals).

21 Chromium, Ti and V are extracted from the oxide minerals chromite, ilmenite and titanomagnetite.
22 Economic concentrations require several tens of percent of the host minerals, and exploration can
23 therefore utilize remote sensing. Chromite is the only naturally-occurring source of Cr that can be
24 economically extracted (Table 1). Vanadium is extracted from titanomagnetite and titanium from
25 ilmenite (Table 1). In some mines, Fe is recovered as a by-product. Chromite forms exclusively in
26 primitive mafic-ultramafic rocks, while titanomagnetite and ilmenite are more common in the evolved
27 parts of LMI. In contrast to chromium, mafic-ultramafic igneous rocks are not exclusively the hosts of
28 Ti and V resources. Vanadium is also extracted from black shale and crude oil, and titanium is
29 concentrated in beach placers. Nickel, Cu, Co and the PGE associate naturally with sulphide minerals
30 which normally occur in very low abundances in LMI. Sulphide mineralisation is commonly difficult to
31 detect visibly – exploration is therefore dependent upon bulk geochemistry or assay and a sound
32 understanding of igneous processes based on mineral chemistry.

33 Chromium, V and Ni are important alloying metals for steel production. Chromium and Ni provide
34 corrosion resistance and are essential ingredients for stainless steel. Vanadium and Cr add hardness,
35 strength and durability against fatigue, which is critical for the performance of steel tools, springs,
36 bearings and gears. Platinum-group metals are used in jewellery. However, because most PGE are
37 chemically (relatively) inert (except for Os), they are extensively used as catalysts for fuel emission
38 reduction and in hydrogen fuel cells (Table 1).

39 2. Background Geology and Economics

40 2.1 Grades and Processing

41 Chromite (FeCr_2O_4) ore is mined at 20-50% chromite. Metallurgical chromite ore requires a
42 concentrate with at least 30 wt.% Cr (44 wt.% Cr_2O_3), whereas refractory ore needs >60 wt.% (Al_2O_3 +
43 Cr_2O_3) including >20 wt.% Al_2O_3 (Pohl, 2011). Following comminution (crushing and grinding),
44 concentration is achieved by gravity and density methods (shaking tables and spirals) before mixing
45 with a flux and smelting. The major product is ferrochrome (an alloy of Fe and Cr) that is directly used
46 in stainless steel production.

47 Platinum-group metal ore is mined together with Ni, often with co-production of Cr, Cu, Co and Au.
48 With the exception of Cr, these metals are hosted in solid solution in base metal sulphides (typically
49 pentlandite, pyrrhotite, chalcopyrite) or as platinum-group minerals (PGM). The PGM are a diverse
50 group of minerals, of which over 135 are known, including native metals, alloys and various other
51 phases (including sulphides, arsenides etc; O'Driscoll and Gonzalez-Jimenez, 2016). Grades for PGE
52 typically range between 1-10 g/t total PGE + Au, although high-grade deposits have values that might
53 exceed 20 g/t. Nickel, Cu and Co grades are typically 0.5-3.0 wt.%, 0.1-2.0 wt.% and 0.05-0.1 wt.%,
54 respectively. Mineral processing requires comminution and concentration (flotation and gravity)
55 before pyrometallurgical and hydrometallurgical extraction, and refining.

56 Ilmenite ore is mined at around 15 wt.% TiO_2 , which is equivalent to some 30 wt.% ilmenite (Charlier
57 et al. 2006). Vanadium ore typically requires around 0.25 wt.% V_2O_5 , although economic extraction is
58 possible at lower grades where the metal is a by-product. The economic recovery of Ti and V, however,
59 depends on many factors, notably how much Ti is locked into titanomagnetite and the extent to which
60 titanomagnetite and ilmenite are intergrown. Mineral processing requires efficient separation of
61 titanomagnetite from ilmenite, which initially requires magnetic separation (titanomagnetite is
62 ferromagnetic, whereas ilmenite is paramagnetic). Vanadium may be processed into ferrovandium,

63 which is used in the steel industry, or recovered as V_2O_5 through a roast-leach process. Titanium is
64 predominantly processed into white TiO_2 -pigment for paint.

65 An overview of grades, ore minerals, products and uses for the main elements currently extracted
66 from LMI is provided in Table 1.

67

68 **2.2 Magmatic Differentiation**

69 The formation of an ore deposit is the culmination of four key factors: *source*, *pathway*, *agent*, and
70 *deposition* (Fig. 1). In the case of ore deposits hosted within LMI, these factors are largely controlled
71 by magmatic differentiation processes, including partial melting, crystallisation, crustal
72 contamination, and immiscibility of coexisting melts (Fig. 1). Elements that are normally distributed
73 widely across a large volume and at low abundances must become concentrated into a small volume
74 to be amenable and economic for mining. Most of the elements that occur in mineralised deposits in
75 LMI are more abundant in the Earth's mantle (the *source*) than in the crust. Therefore ultramafic-mafic
76 magmas (commonly the *agent*) derived from partial melting of the mantle are relatively enriched in
77 these elements, compared to more felsic magmas from crustal source rocks. These ultramafic-mafic
78 magmas ascend from the mantle via conduits (*pathways*) to form crustal magma chambers. In these
79 chambers, progressive crystallization is accompanied by mineral settling which results in the
80 accumulation of mineral products (cumulates) at their base. Rocks in most LMI therefore generally
81 become successively younger upward through the magmatic stratigraphy, and the roof zones are
82 commonly eroded away. This process of fractional crystallisation means that magmas eventually
83 become saturated in mineral phases that might contain elevated abundances of economically
84 interesting metals (*deposition*). Crystallisation, either *in situ* or with some mechanical sorting,
85 facilitates the accumulation of minerals that may reach sufficient quantities to form a viable ore
86 deposit. Progressive crystallisation may also lead to the exsolution of immiscible liquids (sulphide or

87 Fe-rich silicate) and these have significant effects on the concentration and deposition of target
88 elements. The unmixing of an immiscible liquid fraction from a magma may also be triggered by
89 contamination of the crystallising LMI with crustal rocks.

90 The distribution of an element between coexisting phases and residual magma underpins the
91 magmatic differentiation processes that concentrate metals from source to deposition. This includes
92 the partitioning of an element between liquid magma and cumulus mineral (during the extraction of
93 magma from the source region, or crystallisation from magma), or between coexisting liquids (silicate
94 and sulphide magma). This is described as element partitioning and may be quantitatively expressed
95 by the ratio of the concentration of element (i) in one phase to the concentration of i in the other
96 phase. This ratio is termed the *Nernst partition coefficient* (D) such that:

$$98 \quad D_i^{sulphide/silicate} = \frac{C_i^{sulphide}}{C_i^{silicate}}$$

97 [1]

99 Equation 1 is an example expressing D for element i between sulphide liquid and silicate magma,
100 where C denotes concentration of i . If $D > 1$, then element i is said to be *compatible* and preferentially
101 partitioned into sulphide liquid (in the case of Equation 1). For values of $D < 1$, i is *incompatible* into
102 sulphide liquid and remains preferentially in the silicate magma. Whilst the partition coefficient may
103 be used to numerically model trace element behaviour during partial melting and crystallisation, for
104 the purposes of this chapter, D is used as an indication of an element's affinity for different magmatic
105 phases (Table 1). The value of D effectively indicates the magnitude of this affinity.

106 This chapter will focus on the three main groupings of elements that may be mineralised within LMI:
107 chromium (in chromite), Ti and V (in magnetite and ilmenite), and PGE-Ni (with minor amounts of Co,
108 Cu and Au). Examples of exposures of mineralised LMI locations are also provided in Figure 2. Each of
109 the following sections outlines the main minerals and mineralisation processes thought to operate in
110 the formation of LMI. This chapter will therefore chiefly outline the mainstream models and

111 mechanisms thought to be in operation in LMI mineralisation, most of which are orthomagmatic. We
112 will also discuss some high-temperature metasomatic and fluid-based processes, but acknowledge
113 that some other models are beyond the scope of this short review.

114

115 **3. Chromite**

116 The rock type chromitite typically contains ≥ 60 vol.% chromite crystals, together with minor
117 interstitial plagioclase, pyroxene and/or olivine. In LMI it forms seams and layers that range from < 1
118 mm to several meters thick (Mathez and Kinzler, 2016). LMI-hosted chromitites are one of Earth's
119 main Cr repositories (Mungall and Naldrett, 2008; and references therein), and they are also
120 associated with some of the largest PGE deposits. The formation of chromitite has been the subject
121 of much debate in the petrological community (e.g., Boudreau, 2016; Latypov et al. 2013; 2017). A
122 major problem is one of mass balance and specifically the question of how a trace element such as Cr
123 (with abundances typically of 100-1000 ppm in LMI parental melts) can be concentrated to levels of
124 30-40 wt.% in metre-thick chromitite layers. An explanation is required for the observation that
125 silicate magmas that mainly crystallise silicate minerals are able to precipitate only chromite as a
126 liquidus phase during the formation of chromitite layers. The critical steps in forming LMI-hosted
127 chromitite layers and are essentially controlled by the replenishment of the chamber by new magma.
128 The relevant processes are described in the following sections by referring to both thick massive
129 chromitite layers (tens of cm to metres thick) and thinner chromitite seams (mm to cm thick) in LMI.

130

131 ***3.1 Massive chromitite layers***

132 Several massive chromitite layers occur in the Critical Zone of the ~ 2 Ga Rustenburg Layered Suite
133 (Bushveld Complex) and the Peridotite Zone of the ~ 2.7 Ga Stillwater Complex (Cawthorn, 2015;
134 Boudreau, 2016). The Kemi LMI in Finland also hosts a chromitite layer that reaches ~ 100 m thick in
135 the centre of the intrusion. Massive chromitite layers generally formed in the lower portions of their

136 host LMI, where they appear to be intimately associated with the macro-scale rhythmic layering that
137 developed in response to repeated additions of magma to the chamber. Massive chromitite layers
138 developed systematically at the bases of cyclic units in the Bushveld and Stillwater Complexes (Fig.
139 3a). The layers are extensive and commonly of relatively uniform thickness along strike. As an
140 example, the UG2 chromitite in the Bushveld Complex varies between ~0.4 and 2.5 m in thickness and
141 can be traced continuously for 300 km across the Eastern and Western Limbs of the Bushveld Complex
142 (Schouwstra and Kinloch, 2000). A thin (10-20 cm) 'leader' seam is systematically developed
143 immediately above the UG2 chromitite, and thin (< 20 cm) chromitite seams also formed in the
144 footwall rocks. These minor seams (above and below) locally merge and bifurcate with the main
145 chromitite. In the UG2 layer, the chromitite has 75-90 vol.% chromite with interstitial plagioclase and
146 pyroxene. Generally, the UG2 footwall comprises a pegmatoidal feldspathic pyroxenite, with locally
147 developed anorthosite, and the hanging wall is made up of pyroxenite. One of the remarkable features
148 of many massive chromitites of the Critical Zone (particularly well-expressed in the UG2 chromitite) is
149 that they locally cut down through several metres of their footwall cumulates giving rise to structures
150 referred to as 'potholes'. Potholes associated with the UG2 chromitite are circular or elliptical
151 depressions (10 cm to 100 m wide and < 1-40 m deep) with gentle-to-steeply inclined sides that
152 truncate the underlying layered rocks (Latypov et al. 2017a).

153

154 The formation of massive chromitite layers such as those in the Bushveld and Stillwater Complexes
155 has been suggested to be triggered either by magma mixing (following the classic model of Irvine,
156 1977; Fig. 4a), by crustal contamination (Fig. 4a), and by pressure and/or fO_2 fluctuations. In the case
157 of both the magma mixing and contamination mechanisms, the hybrid composition of the resultant
158 magma becomes saturated in chromite only (Fig. 4a), leading to crystallisation of a monomineralic
159 chromite layer as chromitite. Mondal and Mathez (2007), however, argued that the enormous amount
160 of chromite in the UG2 layer must have involved the emplacement of magma that carried entrained
161 chromite phenocrysts, with subsequent gravity-driven accumulation onto the magma chamber floor.

162 While this idea addresses the mass balance problem, more recent work by Junge et al. (2014) showed
163 cryptic mineral compositional variations within the UG2 chromitite, that suggest it was constructed
164 incrementally, rather than in a single 'event'.

165

166 In the Stillwater Complex, the association of chromitite with cyclic harzburgite-orthopyroxenite units
167 has similarly been taken as evidence of chromite crystallisation in response to mixing of resident
168 magma with repeated influxes of olivine-saturated (high-Mg) parental magmas (i.e., the mechanism
169 outlined in Fig. 4a). Horan et al. (2001) used Os isotopes to argue for chromitite formation by mixing
170 of variable proportions of a relatively primitive magma with one that had already seen crustal
171 contamination. In a variant on this theme, Spandler et al. (2005) analysed chromite-hosted inclusions
172 in the chromitites and proposed that assimilation of silica-rich country rock by the incoming magma
173 forced the contaminated magma into crystallising chromite only, thereby forming the chromitite
174 layers (Fig. 4a). Other suggestions include the crystallisation of chromite in response to pressure
175 fluctuations in the magma chamber (Lipin, 1993) and hydration of magma to stabilise chromite as the
176 only crystallising phase. Boudreau (2016; and references therein) discussed how hydration melting
177 could work to produce the successive dunite/chromitite layers of the Peridotite Zone as restitic
178 products of this process. This works by shifting the cotectic in the olivine-chromite-SiO₂ ternary phase
179 diagram (Fig. 4b), expanding the chromite stability field such that the overall composition of the hybrid
180 magma lies in the chromite-only field, and resulting in crystallisation of chromite alone with
181 subsequent formation of the monomineralic layers by crystal settling or in situ crystallisation as for
182 the mixing/assimilation model(s) (Fig. 4b).

183

184 **3.2 Chromitite seams**

185 Chromitite seams are typically a few millimetres to several centimetres thick and therefore
186 uneconomic for chromium. Instead, the chromitite in the Merensky Reef (Bushveld Complex) is mined
187 as one of the largest PGE deposits on Earth, highlighting the important association between chromitite

188 and PGE mineralisation in LMI. Similar to the massive chromitite layers, many seams formed directly
189 at lithological contacts that preserve evidence of magma replenishment. The Merensky Reef
190 chromitites formed between 20-400 m above the UG2 chromitite, close to the top of the transition
191 between the Critical Zone and the Main Zone. Like the UG2, they are continuous around the Western
192 and Eastern Limbs of the Bushveld Complex. Although large variations occur in thickness and
193 composition, the Merensky Reef is generally < 30 cm thick and comprises a pegmatoidal feldspathic
194 pyroxenite, bound above and below by thin (1-2 cm) chromitite seams (Fig. 3b). It is generally
195 underlain by anorthosite and overlain by pyroxenite. Like the UG2 chromitite, the Merensky package
196 locally cuts down into its footwall rocks to form potholes (Latypov et al. 2015).

197

198 Other examples of chromitite seams are found in the ~60 Ma Rum Layered Suite (NW Scotland;
199 Emeleus et al. 1996). The Rum Layered Suite comprises the Eastern and Western Layered Intrusions
200 and the Central Intrusion (ELI, WLI and CI, respectively). The ELI comprises 16 coupled peridotite-
201 troctolite ± gabbro ± anorthosite macro-rhythmic units, each of which is considered to be the result
202 of differentiation of batch fractionation within the Rum magma chamber (Brown, 1956; Emeleus et al.
203 1996). Thin (0.5 mm – 2 cm) chromitite seams formed in two principal environments. Chromitite
204 seams drape the contacts of several of these units – for example, between the anorthositic tops
205 (underneath) and peridotitic bases (above) of Units 6/7, 7/8 and 11/12 (O’Driscoll et al. 2010). These
206 boundary chromitite seams are continuous over lateral distances of 100s of metres, generally
207 maintaining a uniform thickness of ~2 mm around topographic culminations, depressions and even
208 overhangs in the footwall anorthosite (Fig. 3c). Locally, pockets of chromitite protrude down into the
209 underlying anorthosite; these were termed ‘cone-structures’ by O’Driscoll et al. (2010) and may
210 represent small-scale analogues of the Bushveld potholes. The second mode of occurrence on Rum is
211 as chromitite seams wholly contained within ELI and WLI peridotite bodies. These seams are more
212 numerous than the boundary chromitites – hundreds have been documented from the Unit 10
213 peridotite alone, and there are likely to be thousands present in the ~420 m thick WLI (Hepworth et

214 al. 2017; 2018). These seams are also typically associated with a change in olivine texture on either
215 side of them (Fig. 3d).

216

217 Many of the processes invoked for massive chromitites (e.g., magma mixing, crustal contamination)
218 are also candidates for the formation of chromitite seams. Because the amount of Cr hosted in seams
219 is so much less than massive layers, the mass balance problem is less acute. The complex field relations
220 of the Merensky Reef chromitites, in particular around potholes (e.g., Latypov et al. 2015; 2017a),
221 demonstrate that chromitites drape all parts of the outer edges of potholes, including vertical
222 sidewalls and overhangs. These relations are very difficult to explain by crystal settling (or any sort of
223 mechanical accumulation of chromite crystals) alone. Instead, they suggest that the chromitite
224 crystallised in situ. Latypov et al. (2017b) emphasised magmatic erosion of the footwall to explain
225 potholes and associated features, such as ‘undercutting’ MR, where sill-like bodies of chromitite from
226 pothole edges protrude laterally into the footwall rocks.

227

228 Alternatively, it has been argued that chromitite seams may represent reaction products rather than
229 products of direct crystallisation. For example, Nicholson and Mathez (1991) proposed that the
230 Merensky Reef chromitites formed at metasomatic fronts, driven by dissolution of the crystal mush
231 by percolating H₂O-rich melts. The presence of H₂O during these reactions causes the crystallisation
232 temperatures of all of the coexisting silicate phases (relative to chromite) to be suppressed, allowing
233 the latter to crystallise alone. This suggestion has similarities to the hydration melting model of
234 Boudreau (2016) for the Stillwater massive chromitites. Mathez and Kinzler (2017) drew attention to
235 the similarities between the Merensky and Rum chromitite seams. The Rum unit boundary seams
236 were initially interpreted in the context of the batch fractionation model for the ELI macro-rhythmic
237 units; so that a new replenishment of magma would first crystallise chromite which, after gravitational
238 settling to form a chromitite seam, would be followed by successive crystallisation and settling of
239 olivine and olivine + plagioclase to form the overlying peridotite and troctolite portions of each unit

240 (cf. Brown, 1956). However, the observation that the Rum chromitites faithfully follow footwall
241 topography, including overhangs, without significant changes in seam thickness, led to a reevaluation
242 of the traditional crystal settling model for chromitite formation and a proposal that new batches of
243 superheated magma might thermochemically erode their footwall on entry to the magma chamber
244 (O’Driscoll et al. 2010). In the latter scenario, chromitites represent an in situ reaction product
245 between the underlying feldspathic cumulate and the incoming primitive magma, where assimilated
246 troctolitic cumulate modified the composition of the magma to the extent that it crystallised chromite
247 alone. Subsequent studies have interpreted the intra-peridotite chromitite seams in a similar way,
248 such that reactive infiltration within the crystal mush occurred as a consequence of the incoming
249 magma being emplaced into the crystal mush rather than onto the magma chamber floor (Hepworth
250 et al. 2017; 2018).

251

252 **4. Magnetite-Ilmenite Ti-V**

253 Layered mafic-ultramafic intrusions become enriched in titanium (Ti) and vanadium (V) in their more
254 fractionated portions and economic Fe-Ti-V resources may consequently form. The dominant minerals
255 are ilmenite (as part of the ilmenite-haematite solid solution series) and titanomagnetite (as part of
256 the ulvöspinel-magnetite solid solution series) – Figure 5. Titanium is a major component in both
257 ilmenite (FeTiO_3) and the ulvöspinel component of titanomagnetite ($\text{Fe}_2\text{TiO}_4 - \text{Fe}_3\text{O}_4$). In contrast,
258 vanadium partitions into titanomagnetite and to a lesser extent ilmenite where it substitutes for Fe^{3+}
259 (Fig. 5). Notable resources include the Bushveld Complex in South Africa, the Panzhihua Intrusion in
260 China, Windimurra Igneous Complex in Australia, and the Mustavaara Fe-Ti-V oxide deposit in Finland
261 (Figure 2 and see Table 1). Resources within LMI are typically stratiform layers that accumulated at
262 the bases of the magma chambers. Additional resources are found in norite dykes and intrusions
263 associated with anorthosite provinces – notably at Allard Lake in Canada and Tellnes in Norway.

264 **4.1 Source regions**

265 Similar to other metals discussed in this chapter, titanium and vanadium enter mafic-ultramafic
266 magma during partial melting of the mantle. While mid-ocean ridge basalt is low in Ti (1.53 wt.% TiO₂)
267 and V (280 ppm) (Gale et al. 2013), many mantle hot spots produce lavas with much higher
268 concentrations (for example, up to 5 wt.% TiO₂ and 500 ppm V in the Scoresby Sund basalts of
269 Greenland (Larsen et al. 1989) and up to 4.3 wt.% TiO₂ and 405 ppm V (Peate 1997) in the Paraná
270 basalts of Brazil). These concentrations reflect variations in the mantle source, notably the abundance
271 of the dominant Ti and V hosts (phlogopite, rutile and perovskite; Foley et al. 2000; Corgne and Wood
272 2002). The elevated concentrations are indicative of contributions from mantle material that either
273 hasn't been depleted by melt extraction, or become enriched through subduction, melt or volatile
274 addition (from subcontinental lithospheric mantle). Mafic-ultramafic magmas derived from hot spots
275 (and thus LMI) are, therefore, much more fertile for Ti and V than their mid-ocean ridge equivalents.

276 ***4.2 Titanomagnetite and ilmenite crystallisation***

277 As mafic-ultramafic magmas start to crystallize, Ti and V become concentrated into the residual melt
278 fraction. Ultramafic cumulates do not contain much Ti or V, and the crystallization of chromite ceases
279 before ilmenite and titanomagnetite start to form. There is typically a significant stratigraphic
280 separation (known as the spinel gap) in many LMI between the deepest strata that carry chromite (at
281 the base of the magmatic stratigraphy) and the upper layers that carry ilmenite and titanomagnetite.

282 Ilmenite and titanomagnetite crystallisation is controlled by the availability of Ti, Fe, Si and O in the
283 magma. The minerals crystallize as separate, homogeneous grains of (haematite-bearing) ilmenite and
284 ulvöspinel-rich titanomagnetite. The partitioning of Fe and Ti is governed by the equilibrium



286 which is sensitive to temperature and the oxidation state of iron (Fe²⁺ and Fe³⁺, which reflects the
287 prevailing oxygen fugacity) (Fig. 5). This equilibrium is commonly used as a geothermometer and
288 oxygen barometer. In mafic-ultramafic magmatic systems, the element partitioning is part of the

289 wider equilibrium that also encompasses iron-bearing mafic silicates and quartz (Lindsley et al. 1990;
290 Andersen et al. 1993; Sauerzapf et al. 2008). During primary crystallization, ilmenite carries significant
291 haematite in solution and titanomagnetite is ulvöspinel-rich. Therefore, the primary magmatic
292 assemblage may cause significant Ti to be partitioned into titanomagnetite (Fig. 5).

293 Vanadium is highly compatible in the titanomagnetite crystal lattice. Consequently, once
294 titanomagnetite starts to crystallize, the residual magma becomes rapidly depleted in V (Fig. 6a). As
295 dense minerals accumulate on the floor of a magma chamber, the lowermost, first-formed magnetite-
296 rich layer will therefore contain the highest V content (Fig. 6b).

297 **4.3 Late- or post-magmatic evolution**

298 During cooling, titanomagnetite is prone to oxidation leading to a transformation known as "oxy-
299 exsolution" (Lindsley, 1991; Fig. 5). This process converts ulvöspinel to ilmenite and magnetite,
300 following the reaction:



302 Ilmenite cannot remain in solid solution and segregates to form lamellae, subgrains or trellis-
303 microstructures within the titanomagnetite host. The composition of the exsolved ilmenite and
304 remaining titanomagnetite reflects the temperature and oxygen fugacity at the time of exsolution (Fig.
305 5). The lower the temperature, the purer the ilmenite and magnetite. Conversely, the size of these
306 exsolved sub-grains is controlled by the diffusion of Fe and Ti, and diffusion rates are generally highest
307 at high temperature. Titanomagnetite-ilmenite intergrowths are ubiquitous within mafic igneous
308 rocks (Fig. 5). The scale and purity of the mineral pairs strongly influence the potential for mineral
309 processing and therefore the economic prospects of metal recovery. An ideal ore will have undergone
310 prolonged oxy-exsolution to low temperature at a slow enough rate for the exsolved minerals to
311 separate into large grains.

312 **4.4 Formation of economic resources**

313 While the mineral chemistry of ilmenite and titanomagnetite is well understood, the process of ore
314 formation in LMI is still enigmatic. The traditional view is that the minerals crystallize from their host
315 magmas and settle (because of their high specific gravity) to the base of the magma chamber.
316 However, Ti and V have similar mass balance problems as Cr, mentioned above. Mafic magmas reach
317 saturation with ilmenite and titanomagnetite at TiO_2 contents of less than 5 wt.%, so there remains a
318 problem as to how rocks with >60 wt.% of these minerals (such as the Main Magnetite Layer in the
319 Bushveld complex, Cawthorn and Ashwal 2009) can form. Unlike chromite, phase relations do not
320 support crystallization of ilmenite and/or titanomagnetite without simultaneous crystallization of
321 pyroxene and plagioclase. It is therefore not feasible that such concentrations can be achieved by
322 crystal accumulation alone since this would require an extremely efficient sorting mechanism. This
323 problem is linked to the question as to whether residual magmas become iron- or silica-enriched
324 during progressive fractionation (e.g., Hunter and Sparks 1987; Brooks et al. 1991; Tegner 1997).

325 The issues outlined above have led to the realization that many mafic magmas separate into
326 conjugate, or immiscible, Si-rich and Fe-rich magmas as they fractionate (McBirney 1975; Charlier et
327 al. 2011; Holness et al. 2011; Namur et al. 2012). The Fe-rich magmas may contain up to 40 wt.% FeO
328 and 3 wt.% TiO_2 , as well as up to 18 wt.% P_2O_5 (Hou et al. 2018). These magmas are dense and likely
329 to become preferentially incorporated into the cumulates at the bottom of LMI – by contrast, the Si-
330 rich magmas are relatively buoyant and may eventually percolate into the roof zone (Nielsen et al.
331 2015). Erosion of layered intrusions preferentially removes the Si-rich upper parts, leading to over-
332 representation of their deeper Fe-rich portions. The immiscibility of Fe-rich magma offers an
333 additional mechanism of fractionation and provides a feasible explanation for the concentration of
334 ilmenite and titanomagnetite in LMI. Interestingly, it also offers a linkage between the evolution of
335 LMI and a group of enigmatic magnetite-apatite ores known as Nelsonites or Kiruna-type iron ores
336 (e.g., Nyström and Henríquez 1994; Tornos et al. 2016).

337

338 **5. PGE-Ni-(Co-Cu-Au) mineralisation**

339 Like other magmatic ore deposits mentioned in this chapter, mineralisation of precious metal
340 magmatic sulphide deposits can be considered in terms of a source, a pathway, and mechanisms for
341 collection of the metals of interest (Fig. 1). Magmatic sulphide deposits are accumulations of
342 immiscible sulphide melt which segregated from a silicate magma and within which elevated
343 concentrations of base and precious metals occur. Magmatic sulphide deposits are intrinsically
344 polymetallic but can be categorised into (i) precious metal PGE-Ni-(Co-Cu-Au) and (ii) base metal Ni-
345 Cu-(Co-PGE) deposits (where those elements listed in brackets are by-products). Precious metal
346 magmatic sulphide deposits typically form stratiform ore bodies (or 'reefs') in LMI and generally
347 comprise a low volume (< 1 modal %) of sulphide minerals. In this section, we will outline the factors
348 that control the formation of PGE and nickel (Ni) magmatic sulphides, with their concentrations of less
349 abundant cobalt (Co), copper (Cu) and gold (Au).

350

351 **5.1 Source regions**

352 The processes governing mineralisation of the PGE (Os, Ir, Ru, Rh, Pt, Pd), Ni, and to some extent Co,
353 Cu and Au, are interlinked due to their affinity for sulphide melts at mantle and crustal conditions –
354 hence their geochemical behaviour is broadly classified as chalcophile. The PGE have traditionally
355 been further subdivided into a compatible Ir-group (IPGE) and an incompatible Pt-group PGE (PPGE),
356 based on their behaviour during partial melting of the mantle (Table 1). The IPGE (Os, Ir and Ru, which
357 behave in a similar fashion to Ni and Co) are sometimes hosted as nano- or microscale platinum-group
358 minerals (Pitcher et al., 2009) within silicate and oxide minerals, especially olivine and spinel
359 (chromite)). In contrast, the PPGE (Rh, Pt and Pd, with similar behaviour to Au and Cu) are
360 incompatible with silicate and oxide minerals and largely hosted by sulphides in the mantle (Lorand
361 and Luguet, 2016).

362 As partial melting starts, the most easily fusible minerals in the mantle source region will melt first –
363 typically clinopyroxene, garnet and sulphides – and continued melting incorporates increasing
364 amounts of spinel and olivine (Pearson, 2005 and Figure 7). Progressively higher degrees of partial
365 melting (expressed as melt fraction, F) incorporate more sulphides, until the sulphide budget of the
366 source is exhausted. Given the strong affinity of chalcophile elements, and in particular the PPGE, for
367 sulphide (expressed as high partition coefficients, D_i), any residual sulphide in the mantle source will
368 suppress the uptake of PGE into the melt being formed. Thus, silicate magmas which are S-saturated
369 are not fertile for PGE. In contrast, if melting continues to the point of sulphide exhaustion, the full
370 PGE budget of the mantle source will be imparted to the S-undersaturated silicate magma, maximising
371 magma fertility for the PGE (Fig. 7). Beyond this point, continued melting merely dilutes the
372 concentration of these elements in the silicate magma. For example, assuming a peridotite mantle
373 source with a starting concentration of 200 ppm S, all sulphides would be melted and dissolved in a
374 silicate magma by the time $F=13.5\%$ of equilibrium batch partial melting (Naldrett, 2011). In this
375 model, fractionation of the PGE in the mantle occurs largely because of the different melting points
376 of the minerals hosting these elements (i.e., sulphide vs oxide – Fig. 7). Ore deposit formation models
377 are based upon this paradigm and can be used to describe the fertility of basaltic magmatic systems
378 for PGE-Ni-(Co-Cu-Au) mineralisation.

379 Whilst this paradigm is broadly applicable, empirical studies of mantle sulphides have shown that
380 other factors influence the behaviour of the PGE, Ni, Co, Cu and Au during partial melting. For most
381 pressures and temperatures at which partial melting takes place, the mantle sulphide budget occurs
382 as two coexisting sulphide phases (e.g., Bockrath et al., 2004; Lorand and Luguét, 2016) – a crystalline
383 Fe-rich monosulphide solid solution (MSS) relatively enriched in IPGE, and an immiscible molten Ni-
384 Cu sulphide enriched in PPGE (Fig. 8). During partial melting, the immiscible sulphide gradually
385 becomes incorporated into silicate magma, while the MSS tends to be retained in the mantle source.
386 Thus, silicate magmas derived from mantle partial melting are generally enriched in PPGE and
387 depleted in IPGE, relative to the initial bulk mantle composition (Bockrath et al., 2004).

388 Empirical studies have also shown that the sulphides in different mantle lithologies (e.g., peridotite
389 vs. eclogite; McDonald et al., 2017) and reservoirs (i.e., enriched or depleted; Hamlyn and Keays, 1986)
390 may have very different metal budgets. Furthermore, multiple generations of sulphides may be
391 present within a single mantle reservoir, with each generation preserving a distinct precious metal
392 inventory (Hughes et al., 2017). Some mantle sulphides are interstitial to the major silicate and oxide
393 minerals within that reservoir, but other sulphides may occur as inclusions within these minerals.
394 Thus, mantle sulphides may not always be available for entrainment into silicate magmas during
395 partial melting and, if entrained, will impart varying precious metal characteristics to that silicate
396 magma. Yet classic models (e.g., Fig. 7) used to predict the fertility of mantle-derived magmas assume
397 that interstitial sulphides of uniform composition are always accessible for melting. Unravelling these
398 discrepancies leads to a better understanding of why certain parts of the Earth's crust appear pre-
399 disposed to precious metal mineralisation.

400

401 ***5.2 Mechanisms of mineralisation***

402 Fertile, S-unsaturated silicate magma ascends from a partially melted mantle source reservoir through
403 the lithosphere. The decrease in pressure associated with this ascent increases the ability of the
404 silicate magma to dissolve S such that the Sulphur Content at Sulphide Saturation (SCSS) increases (Li
405 and Ripley, 2005). This inverse relationship between pressure and SCSS means that without further
406 modification the magma would be unlikely to reach S-saturation leading to mineralisation. The
407 following subsections outline mechanisms by which a silicate magma may reach S-saturation in the
408 crust, in order to form an immiscible sulphide liquid that can accumulate to form an ore deposit.

409 An important consideration in the mineralisation of PGE-Ni-(Co-Cu-Au) deposits is the volume of
410 sulphide in the magmatic system. This is described as the mass ratio of the silicate magma divided by
411 that of the sulphide liquid, termed the R-factor (Campbell and Naldrett, 1979):

413
$$R = \frac{m_{silicate}}{m_{sulphide}}$$

412

414 Given the very high partition coefficients of the PGE from silicate magma into sulphide (several orders
415 of magnitude greater than D_{Ni} or D_{Cu}), the optimal concentrations of PGE in sulphide (the tenor) will
416 be achieved in smaller volumes of sulphide. Therefore, deposits with a high R-factor (e.g., $> 10^4$) have
417 low volumes of high tenor sulphide (Fig. 9). However, the efficiency with which an immiscible sulphide
418 liquid circulates in a silicate magma, and therefore the extent to which chalcophile metals diffuse into
419 the sulphide liquid, is also a function of the R-factor (Mungall, 2002) – Figure 9.

420 The formation of a potentially economic ore deposit requires the accumulation of this small volume
421 of sulphide liquid into a narrow stratiform horizon (or reef), again usually at the base of the magma
422 chamber. In some cases, sulphide abundance may be so low that mineralisation is not easily visible.
423 Thus, an understanding of the mechanisms of mineralisation in LMI facilitates exploration for new
424 PGE-Ni-(Co-Cu-Au) reefs.

425 **Magma mixing:** Replenishment of a silicate magma chamber with a new pulse of magma can lead to
426 S-oversaturation. This appears to be particularly effective in magmatic systems with a significant
427 ultramafic component, and for magmas close to S-saturation prior to mixing. During magma mixing,
428 several factors can contribute to a magmatic system achieving S-saturation. These are (i) pressure
429 changes associated with the introduction of a new magma batch (increasing pressure leads to a
430 decrease in SCSS), (ii) temperature changes (increasing temperature causes an increase in SCSS) and
431 (iii) compositional changes such that increasing FeO content results in an increase in SCSS (Wykes et
432 al., 2014). These factors are summarised in Figure 10, which shows how magma mixing may lead to S-
433 saturation in compositional space.

434 As with the formation of chromitite layers, open-system replenishments are thought to underpin the
435 formation of many classic PGE-Ni-(Co-Cu-Au) reefs. Most notable amongst these are the Merensky

436 Reef and UG-2 of the Bushveld Complex in South Africa. Other examples include the Main Sulphide
437 Zone of the Great Dyke in Zimbabwe, the Munni Munni Complex in Australia, Penikat and Portimo (and
438 other Fennoscandian intrusions) and other smaller bodies such as the Rum Layered Suite.

439 **Fractional crystallisation:** During fractional crystallisation of a basaltic silicate magma, the compatible
440 chalcophile elements (Ni, Co, IPGE) gradually become incorporated (by partitioning or as discrete
441 PGM) into minerals such as olivine and chromite. This fractionation may be monitored, for example,
442 by the abundance of Ni in olivine – early-formed olivine (i.e., with high forsterite content) will contain
443 higher abundances of Ni than low-Fo olivine formed much later in the fractionating assemblage. In
444 contrast, incompatible chalcophile elements (PPGE, Cu, Au) and S will not be incorporated into silicate
445 and oxide mineral phases during fractional crystallisation and consequently the concentration of these
446 elements increases in the residual silicate magma. The increase in S concentration, coupled with
447 cooling, can culminate in the silicate magma reaching S-saturation (Fig. 10), with the accompanying
448 segregation of a fraction of immiscible sulphide liquid into which all remaining chalcophile elements
449 partition (Table 1). Thus, prolonged fractional crystallisation (typically > 40% crystallisation; Naldrett,
450 2011) can achieve S-saturation with a high R-factor.

451 Protracted fractionation has implications for the relative proportions of chalcophile elements in a
452 potentially mineralised sulphide-bearing layer, as well as for the location of the resultant reef in the
453 LMI. For example, mineralisation may be particularly enriched in PPGE (especially Pd), Cu and Au, but
454 have extremely low abundances of IPGE and Ni in comparison to other reefs formed via different
455 mechanisms (such as magma mixing). These Pd-Au-Cu enriched reefs may also occur higher in the
456 layered sequence and be associated with gabbroic cumulates, as opposed to pyroxenitic or noritic
457 cumulates lower down in the magmatic succession. Notable examples of reefs formed by extended
458 fractional crystallisation include the Platinova Reef of the Skaergaard Intrusion in Greenland, Sonju
459 Lake and the Layered Series at Duluth in the USA. These magmatic systems may not have incorporated
460 any new magma injections and the Skaergaard Intrusion is a notable case where protracted magmatic

461 differentiation resulted in the formation of the Platinoval Reef relatively high up in the layered
462 sequence (Nielsen et al., 2015).

463 **Contamination:** Contamination of a basaltic magma by addition of silica (by assimilation of silica-rich
464 sedimentary sequences or silica-rich melts from wall rock anatexis) or carbon-rich crustal lithologies
465 (such as mudrocks, graphite, coal, etc) may trigger a redox-controlled S-saturation event (Fig. 10).
466 Most effective, however, is the contamination of basaltic magmas with sulphide-rich crustal rocks
467 (such as shales, evaporates and coal) mainly because this increases the concentration of S in the
468 magma, inducing S-oversaturation (Ripley and Li, 2013). This process generally leads to the production
469 of large volumes of sulphide liquid and thus a magmatic system with low R-factor. Sulphides produced
470 in this way tend to have much lower concentrations of PGE than sulphides in high R-factor magmatic
471 systems. As a result, the mechanism of crustal contamination could be considered, in certain cases, to
472 be counterproductive to the formation of stratiform PGE-Ni-(Co-Cu-Au) reefs in LMI.

473 In other cases, however, evidence from stable isotopes ($\delta^{34}\text{S}$ or $\Delta^{33}\text{S}$) indicates that contamination may
474 in fact have played an important role, particularly during the formation of several well-known
475 examples of PGE-Ni-(Co-Cu-Au) mineralisation such as the Platreef in the Northern Limb of the
476 Bushveld Complex (Kinnaird and McDonald, 2018) and the Ahmavaara deposit of the Suhanko
477 Intrusion in Finland (Makkonen et al., 2017). The impact of contamination on the formation of PGM
478 mineralisation associated with high tenor sulphides is still debated, particularly in the case of the
479 Platreef (Kinnaird and McDonald, 2018). Studies indicate that the high tenor Platreef sulphides may
480 have formed deeper in the magmatic plumbing system and prior to local contamination of the Platreef
481 parental magma. The high tenor sulphide liquid was then transported via the magma conduits to the
482 current level of the Platreef and was ultimately responsible for the mineralisation seen today. If this
483 is correct then the widespread local contamination of the Platreef footwall by S-rich crustal rocks is
484 not likely to have been an important ore-forming process in this deposit.

485 **Fluids:** Hydrothermal or deuteritic high-temperature late-magmatic hydrous fluids may play an
486 important role in the formation of PGE reefs. Models for fluid-based mineralisation in LMI generally
487 suggest that the chalcophile elements and S are scavenged by upward-migrating fluids percolating
488 through the cooling cumulate pile. These fluids are exsolved from cooling intercumulus melt, may be
489 aqueous, saline and carbonic in composition, are likely halogen-rich, and are capable of dissolving
490 appreciable amounts of S and precious metals (e.g., Hanley et al., 2005). Notable examples of PGE-Ni-
491 (Co-Cu-Au) mineralisation that have been attributed to metasomatic processes are the J-M Reef of
492 the Stillwater Complex in Montana, USA (Boudreau, 1999) and the Merensky Reef in the Bushveld
493 Complex (e.g., Ballhaus and Stumpfl, 1986; Kanitpanyacharoen and Boudreau, 2013). Recent studies
494 have also highlighted the critical role of fluids in mineralisation in some portions of the Main Zone of
495 the Bushveld Complex which were previously thought unmineralised, thus representing an exciting
496 new exploration frontier (McFall et al., 2019).

497

498 **6. The Future of LMI Resources**

499 The transformation of society away from carbon-based fossil fuel will fundamentally change the
500 resource industry. There will be a much greater demand for metals and minerals that are used in clean
501 energy production, energy storage, and decarbonisation. LMI host important resources of Sc, PGE, Co,
502 V and Mg-silicate minerals that will be essential for this transformation. Although not discussed in this
503 article, the Mesoproterozoic Ilímaussaq layered intrusion of southern Greenland is highly enriched in
504 the critical rare earth element (REE) metals and will undoubtedly play an important role in the
505 availability of these metals to society in future decades.

506 **6.1 Future uses and demands on existing commodities**

507 **Chromium** has a well developed market for primary production and recycling. Stainless steel is not
508 expected to undergo substantial market transformations in the circular economy, although there may
509 be more emphasis on sustainable manufacturing and recycling.

510 **Platinum-Group Elements (PGE)** (particularly the PPGE) are extensively used in catalytic fuel emission
511 converters (2-4 g used per catalytic converter) and as industrial catalysts. As society is expected to
512 transition away from the use of fossil fuel, one might expect a long term decline in their demand.
513 However, Pt and Pd are also important components in hydrogen fuel cells, where they are employed
514 to catalyse the oxidation of hydrogen (currently requiring over 28 g of Pt per fuel cell). Fuel cell
515 technology is attractive for mobile energy supply where batteries are not a viable option. While the
516 automotive industry may develop into a large market, fuel cells are particularly attractive for the
517 aviation industry, because of the much higher energy density of hydrogen (around 130 MJ/kg) than
518 traditional aviation fuel (45 MJ/kg) and Li-ion batteries (0.5 MJ/kg).

519 **Cobalt and nickel** are important components of cathodes in lithium-ion batteries, and the metals are
520 therefore critical for the development of energy storage systems. Lithium-ion batteries dominate the
521 market for electric vehicles and domestic storage, and the demand is projected to surge with the
522 transformation of the global vehicle fleet to electricity.

523 **Copper** is essential for electricity infrastructure. National and international power grids will need
524 substantial expansion to deal with the demands from electrification of domestic vehicles as well as
525 the wider transport sector. Further demands can be expected to reduce emissions from the industry
526 sector.

527 **Vanadium** is used for energy storage in Vanadium Redox Batteries. An increased reliance on
528 renewable energy sources (wind, solar) to replace coal- and oil-fuelled power stations requires storage
529 solutions to secure continuous electricity supply. Their large capacity and ability to survive prolonged
530 periods under complete discharge makes these batteries ideal for grid storage.

531

532 **6.2 Future commodities**

533 **Scandium (Sc)** partitions into clinopyroxene and hornblende and is locally concentrated into
534 economically interesting resources in lateritic soils above zoned ultramafic complexes, such as
535 Owendale and Nyngan in Australia. Given the mineralogical similarities, LMI (particularly those with
536 alkaline affinities) may also host future scandium resources. Scandium alloyed with aluminium (Al) is
537 a light and strong high-performance material that can be used to reduce the weight (and therefore
538 the energy consumption) of airplanes and vehicles. Scandium is also used in solid oxide fuel cells.

539 **Mg silicates** (olivine, talc and serpentine) are essential for carbon capture technologies that offer
540 permanent greenhouse gas removal. The reaction of atmospheric CO₂ with Mg silicate is energy
541 positive, so although the slow reaction kinetics are problematic, the process is both energy efficient
542 and energy economic. Other geological alternatives for carbon capture, such as underground storage
543 in oil wells, cannot rival carbonation of magnesium silicate minerals in terms of capacity or
544 permanency. Biological storage in the Earth's ecosystems is less efficient and far less permanent and
545 would require extensive reforms of global land use. Considering that global carbon emissions exceed
546 10 billion tonnes per year (Boden et al, 2017), reductions in emission targets will invariably require
547 large-scale development of carbon capture.

548

549 **Table captions**

550 **Table 1** – Summary of partitioning behaviour of metals discussed in this chapter, the grades and the
551 concentrations of metals in ore deposits relative to mantle abundances, key minerals in ore deposits,
552 uses and products of the relevant metals, and a selection of notable ore deposits worldwide. Note
553 that D_i^* is the bulk partition coefficient between mantle minerals and silicate melt during partial
554 melting of the mantle, under conditions typical for the generation of basaltic magmas. D_i is the
555 partition coefficient for the typical phase which concentrates these elements (e.g., sulphide melt for
556 chalcophile elements like Cu and the PGE; oxides for Cr, V and Ti). The PPGE are the Pd-group
557 platinum-group elements (Rh, Pt, Pd) and IPGE are the Ir-group platinum-group elements (Os, Ir, Ru).
558 PGM in this table denotes the platinum-group minerals. Where elements are listed in brackets, this
559 indicates that these occur as co- or by-products of the main ores listed in this chapter. Information for
560 elements in brackets therefore does not show the main ore locations. Note that 1 ppm = 1 g/t.

561

562 **Figure captions**

563 **Figure 1** – Factors and processes governing mineralisation and the formation of an ore deposit.

564 **Figure 2** – (a) Chromitite: UG-1 chromitite seams (the dark bands) at Dwars River in the Bushveld
565 Complex. These chromitites bifurcate, may have irregular anastomosing contacts with underlying
566 anorthosites, and sometimes contain xenoliths of anorthosite. (b) PGE reef: Merensky Reef
567 photographed in an underground mine in the Western Bushveld. From bottom to top: anorthosite
568 footwall underlying the basal chromitite of the Merensky Reef package, the Merensky pegmatoid and
569 overlying feldspathic pyroxenite. In this image there are three chromitite stringers, including the basal
570 chromitite. (c) Magnetitite: Dark layers anomalously rich in ilmenite and titanomagnetite in the
571 Skaergaard intrusion, Greenland. Layering in the normal gabbro (observed toward the right) is much

572 poorer in these minerals. The layers are approximately planar and would have formed parallel to the
573 magma chamber floor, the succession was rotated 20° after crystallisation. The undulating
574 appearance is due to topography. The younging direction is to the right. Hammer for scale. (d)
575 Magnetite: Two metre thick dark titanomagnetite-rich layer in the Bushveld Complex, South Africa.
576 Similar to Skaergaard, the layer would have formed along a planar surface parallel to the magma
577 chamber floor. The younging direction is up. Photo courtesy of Paul Nex, University of the
578 Witwatersrand.

579 **Figure 3** – (a) Field photograph of the massive G chromitite of the Stillwater Complex, USA. Chromite
580 is black and olivine is brown. Note the interlayering defined by the relative abundances of these two
581 minerals. The entire chromitite package, as well as others in the Stillwater intrusion, is characterised
582 by complex variations in chromite and olivine modal abundances and textural relationships. The
583 person's finger on the left gives scale. (b) QEMSCAN® image showing the lower chromitite of the
584 Merensky Reef package of the Bushveld Complex. The brownish-coloured crystals are orthopyroxene,
585 blue is plagioclase, and black is chromite. Other (brightly-coloured) phases are accessory hydrous
586 minerals (biotite, amphibole etc). (c) Field photograph of a chromitite seam (black) at the contact
587 between peridotite (brown, above) and anorthosite (white, below) from the Unit 11-12 boundary of
588 the ELS in the Rum Layered Suite. This is an oblique view showing the undulose character of the
589 chromitite seam in cross section. The hand lens at the top of the photograph gives scale. (d)
590 QEMSCAN® image of an intra-peridotite chromitite seam from Unit 10 (ELS, Rum Layered Suite). Way
591 up is to the top left of the image. Olivine crystals are yellow, plagioclase is light blue and chromite is
592 black. Other (secondary) phases such as serpentine and amphibole are represented by the dark blue,
593 pink and purple colours. Note the change in texture (olivine grainsize reduction) across the chromitite
594 seam.

595 **Figure 4** – Ternary phase diagrams of the chromite-olivine-SiO₂ system illustrating possible models for
596 chromitite formation in LMI. In (a), theoretical magma mixing and contamination are shown. A

597 possible 'normal' differentiation path (in orange) for a magma with starting composition is labelled A.
598 Crystallisation of olivine will result in the liquid reaching the olivine-chromite cotectic, and eventually
599 orthopyroxene saturation at the end of the orange path (after Irvine, 1977). Replenishment of the
600 magma chamber and mixing of the primitive (B) and differentiated liquids (C, green dot) will move the
601 hybrid composition into the chromite field and lead to crystallisation of chromite only (red dot).
602 Contamination of the differentiating liquid by a SiO₂-rich assimilant (e.g., Spandler et al. 2005) also has
603 the effect of moving the hybrid into the chromite-only field (blue dot/path). In (b), the effects of
604 magma hydration are shown and, in particular, the expansion of the chromite phase field with addition
605 of H₂O. The relatively dry liquid represented by the purple dot on the red dashed line is saturated in
606 chromite and orthopyroxene. With H₂O addition, the phase boundaries shift (blue lines) and the liquid
607 is saturated in chromite alone. See text for further discussion. Redrawn after Boudreau (2016) and
608 references therein.

609 **Figure 5** – Phase diagrams demonstrating exsolution of magnetite, ilmenite and spinel-group minerals
610 with decreasing temperature, and the process of oxy-exsolution. (a) Crystallisation and cooling of Fe-
611 Ti oxide minerals controlling distribution of V and Ti in ore (from Lindsley, 1991). (b) Stages in oxy-
612 exsolution: (i) Oxidation of FeO to Fe₂O₃, (ii) Exsolution of equilibrium pair of titanomagnetite and
613 ilmenite and (iii) Magnetite with ilmenite domains, lamellae-, or trellis-like textures.

614 **Figure 6** – (a) Concentration of V in a fractionating magma crystallising magnetite. (b) Variations of
615 magnetite compositions in the Bushveld Upper Zone, based on Klemm et al. (1985).

616 **Figure 7** – Partial melting of the mantle. (a) Cartoon demonstrating the partial melting process. Fusible
617 mineral phases, including sulphides, melt first. In a partial melting model there will come a point that
618 interstitial sulphides will become exhausted in the mantle source. (b) Sketch model of the
619 concentration of key elements in the silicate magma formed through partial melting of the mantle. As
620 sulphides host most of the chalcophile elements, they control the budget of these elements in the
621 melts that are formed. Continued melting beyond the exhaustion of these sulphides effectively dilutes

622 the concentration of chalcophile elements in the melt, thus a theoretical peak fertility is defined. This
623 model is complicated by sulphides that may be included within other minerals and variation in the
624 trace element composition of mantle sulphides.

625 **Figure 8** – The physical state of mantle sulphides at a range of pressures and temperatures. The shaded
626 area delineates pressure and temperature conditions at which most mantle partial melting takes
627 place. Diagram adapted from Bockrath et al. (2004).

628 **Figure 9** – R-factor models for chalcophile elements Ni, Cu and Pt (as an example for the PGE): (a) The
629 concentration of elements according to various R-factor values for a silicate magmatic system with an
630 immiscible sulphide liquid, adapted from Campbell & Naldrett (1979) and assuming a starting
631 concentration of 250 ppm Ni, 70 ppm Cu and 10.5 ppb Pt. (b) Cartoon models visualising low and high
632 R-factor processes.

633 **Figure 10** – Processes that cause silicate-sulphide liquid immiscibility: Magma mixing, contamination
634 and incompatible concentration during fractional crystallisation change the mass fraction of Fe
635 (described as FeO) and S in the resultant magma, causing that magma to change from S-
636 undersaturated to S-saturated conditions. Diagram after Poulson & Ohmoto (1990).

637

638 **Suggested Further Reading**

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640

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Table 1

Elements	D_i^*	D_i	Ore grade	Concentration factor	Key minerals	Product	Uses	Notable mineralised LMI locations
Ni	1 - 10	500	0.5 - 3 wt.%	5 - 10 x mantle	Pentlandite	Metal	Steel, metals, alloys	Bushveld Complex (South Africa), Great Dyke (Zimbabwe), Monchegorsk (Russia), Stillwater (USA)
(Co)	1 - 10	45 - 80	0.05 - 0.1 wt.%		Pentlandite and other Ni,Fe- or Ni-sulphides			
(Cu)	< 0.1	1500	0.1 - 2 wt.%		Chalcopyrite			
(Au)	< 0.1	1000 - 10000	> 1 - 10 ppm	1000 x mantle	Electrum or alloys with PGE / metalloids, sulphides	Metal	Jewellery, electronics, finance, medicine	Bushveld Complex (South Africa), Skaergaard (Greenland)
PPGE	< 0.1	3000 - 500000	> 1 - 5 ppm total PGE	1000 x mantle	PGM and sulphides	Platinum-group metals	Catalytic converters, jewellery, electronics, fuel cells	Bushveld Complex (South Africa), Great Dyke (Zimbabwe), Monchegorsk (Russia), Stillwater (USA), Lac des Isles (Canada), Skaergaard (Greenland)
IPGE	1 - 10	2000 - 450000			PGM, chromite, sulphides			
Cr	< 0.1		20 - 50 % chromite: > 44 wt.% Cr ₂ O ₃ for metallurgical uses, > 60 wt.% Al ₂ O ₃ + Cr ₂ O ₃ and > 20 wt.% Al ₂ O ₃ for refractory uses.	20 - 60 x mantle	Chromite	Ferrochrome, chromium metal	Stainless steel, superalloys	Bushveld Complex (South Africa), Great Dyke (Zimbabwe), Kemi (Finland), Isle of Rum (Scotland)
V	< 0.1	10 - 100	~ 0.25 wt.% V ₂ O ₅ in magnetite with > 50 % magnetite (low Ti)	25 - 160 x mantle	Magnetite	Titanium metal, TiO ₂	White paint, metal (aerospace/automobile), chemicals, medicine, jewellery	Bushveld Complex (South Africa), Mustavaara (Finland), Panzhihua (China), Lac Doré Complex (Canada), Windimurra (Australia)
Ti	< 0.1		15 wt.% TiO ₂ (equivalent to ~ 30 % ilmenite)	100 - 200 x mantle	Ilmenite	Ferrovandium	Steel, chemicals, bonding agents, ceramics, magnets	

D_i^* is bulk partition coefficient between mantle minerals and silicate melt during partial melting of the mantle and under condition typical for the generation of basaltic magmas.

D_i is partition coefficient into the typical phase which concentrates these elements (e.g., sulphide melt for chalcophile elements like Cu and the PGE; oxides for Cr, V and Ti).

PPGE is Pd-group platinum-group elements (Rh, Pt, Pd). IPGE is Ir-group platinum-group elements (Os, Ir, Ru). Here, PGM stands for platinum-group minerals.

Where elements are listed in brackets, indicates that these occur as co- or by-products to the main ores listed in this chapter. Information for elements in brackets therefore does not show the main ore locations.

Note that 1 ppm = 1 g/t

Figure 1

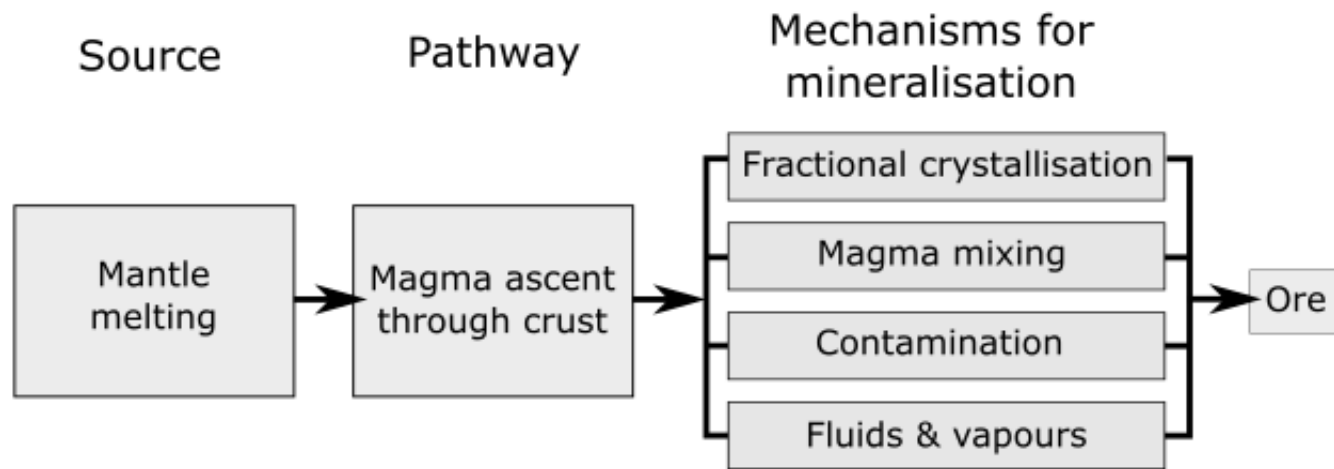


Figure 2

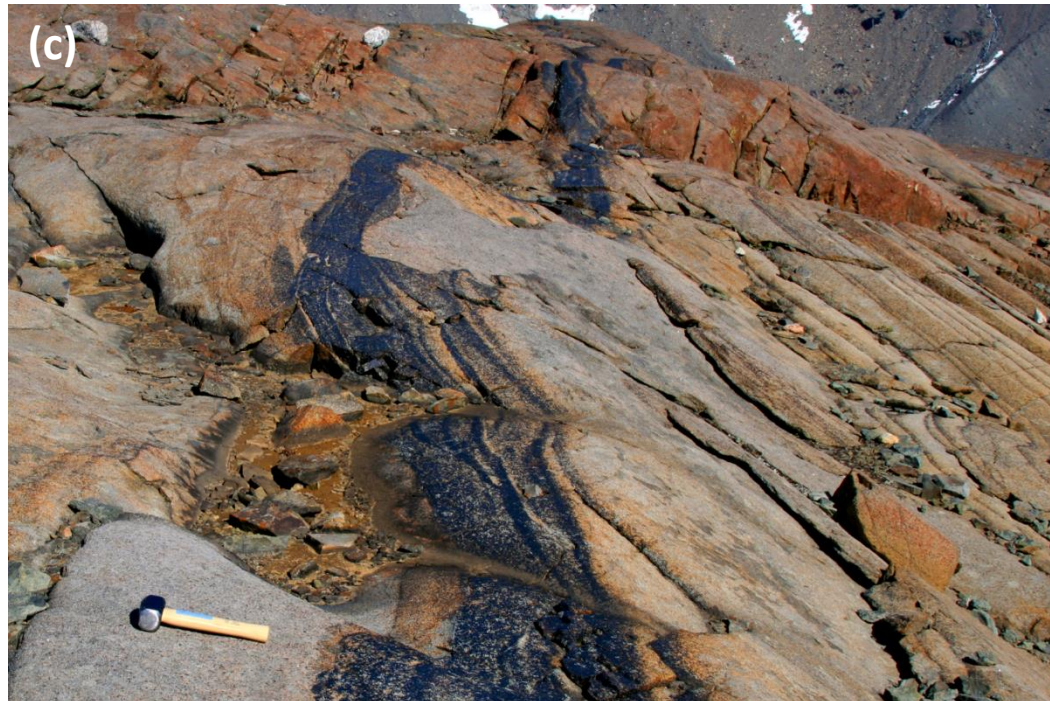
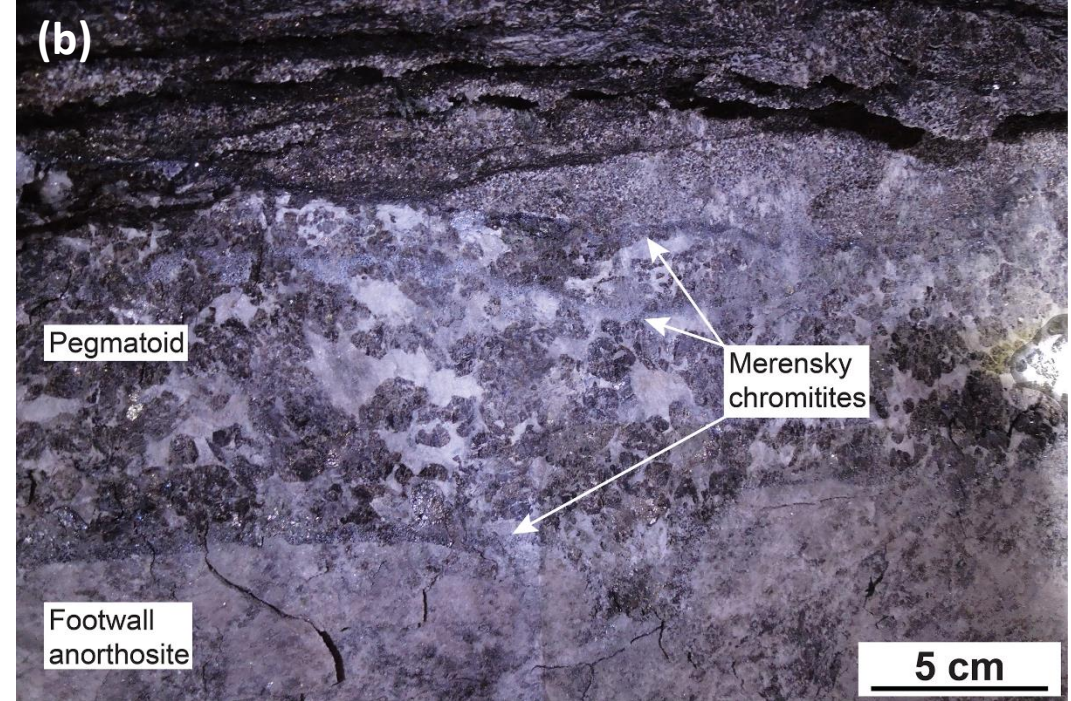


Figure 3

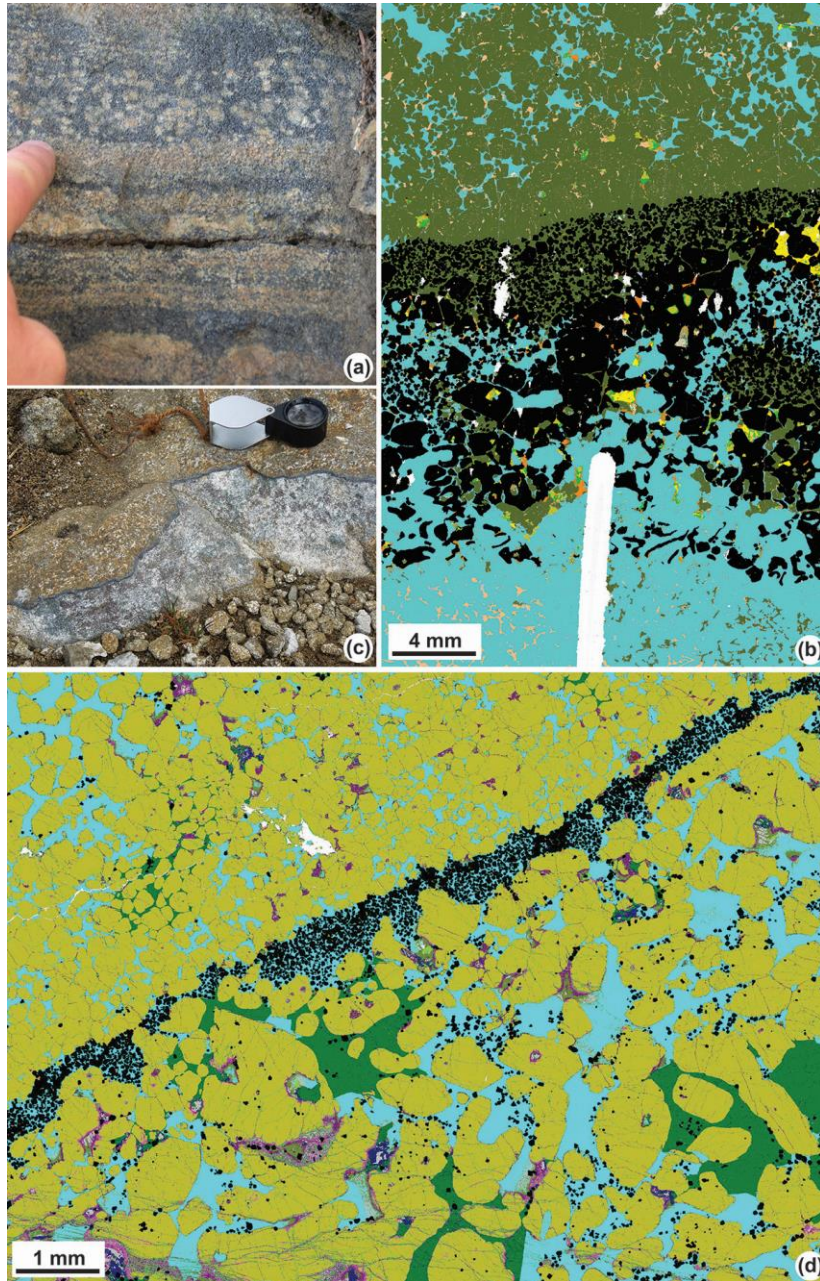


Figure 4

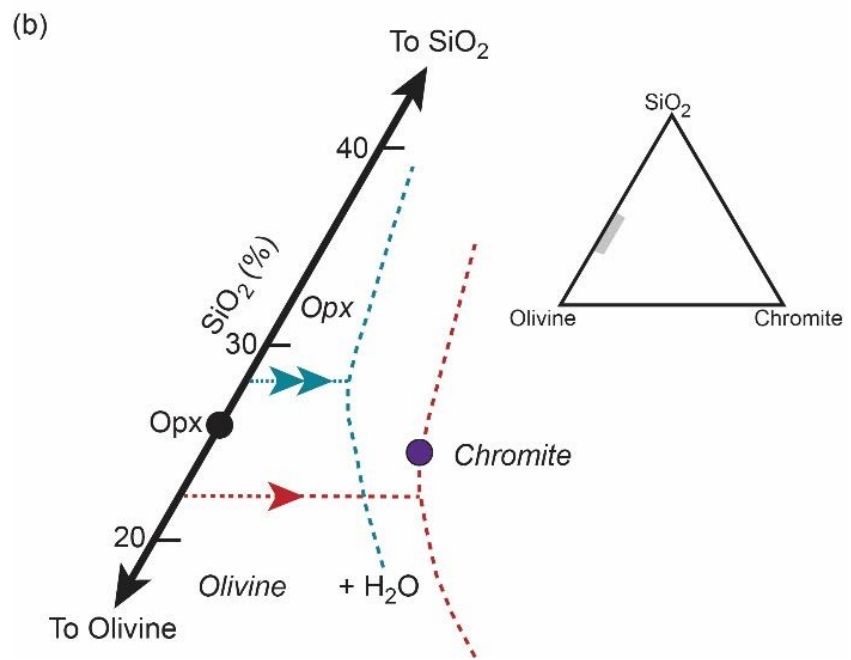
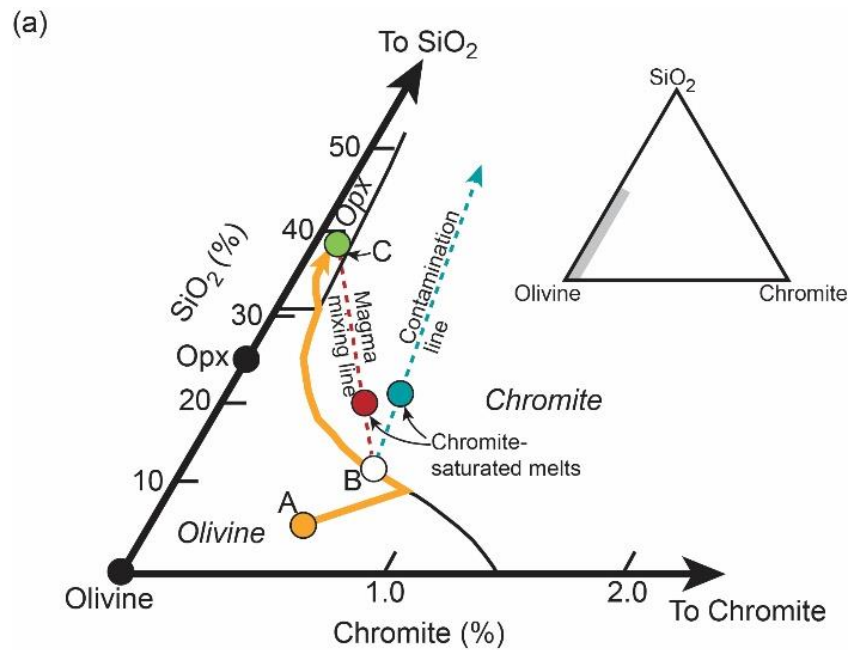


Figure 5

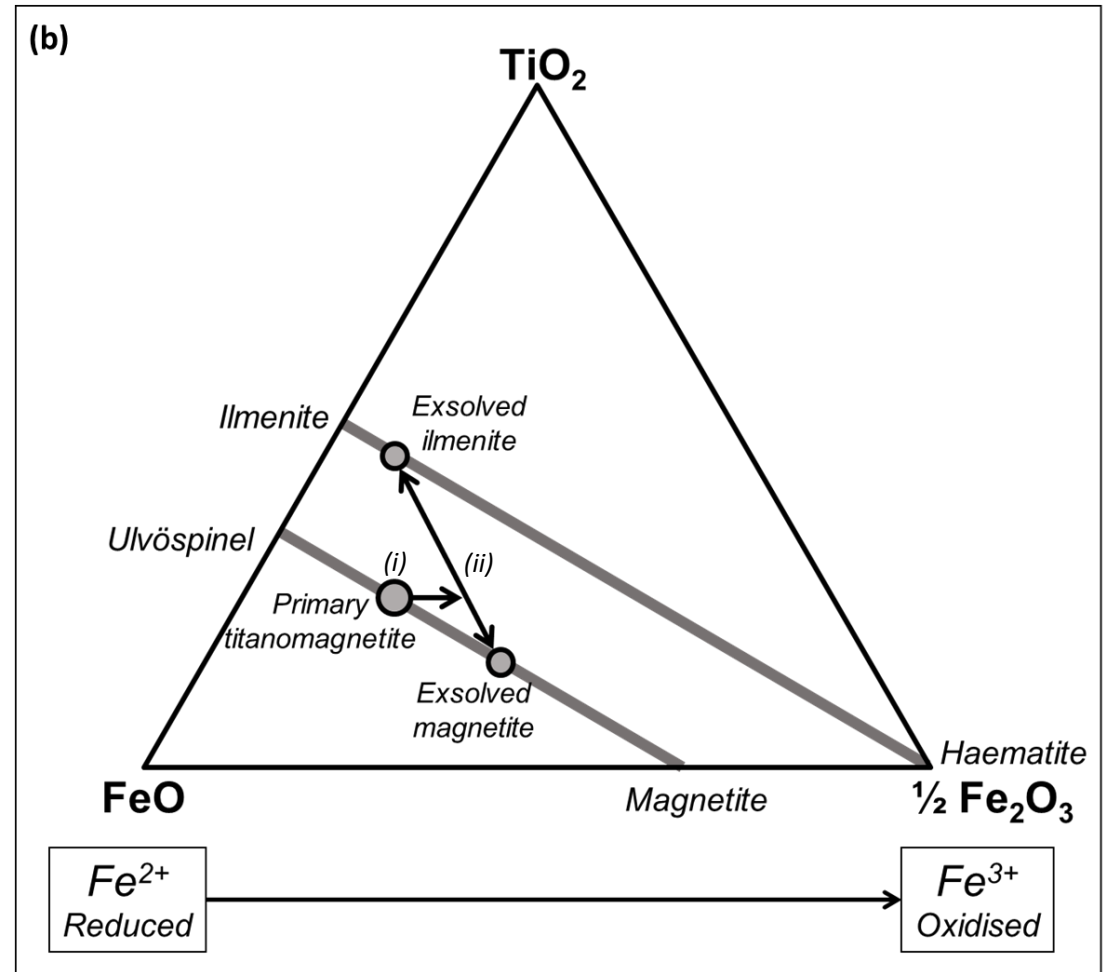
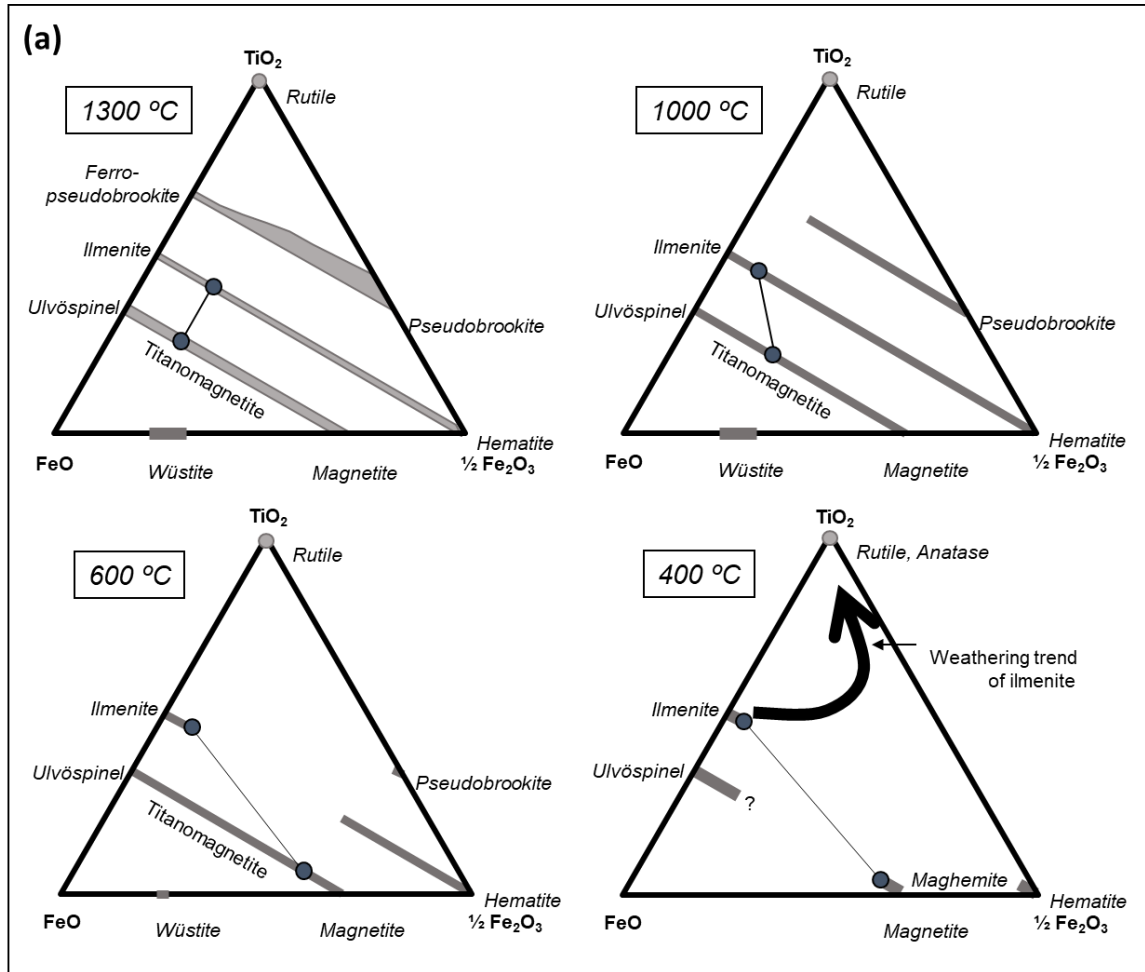


Figure 6

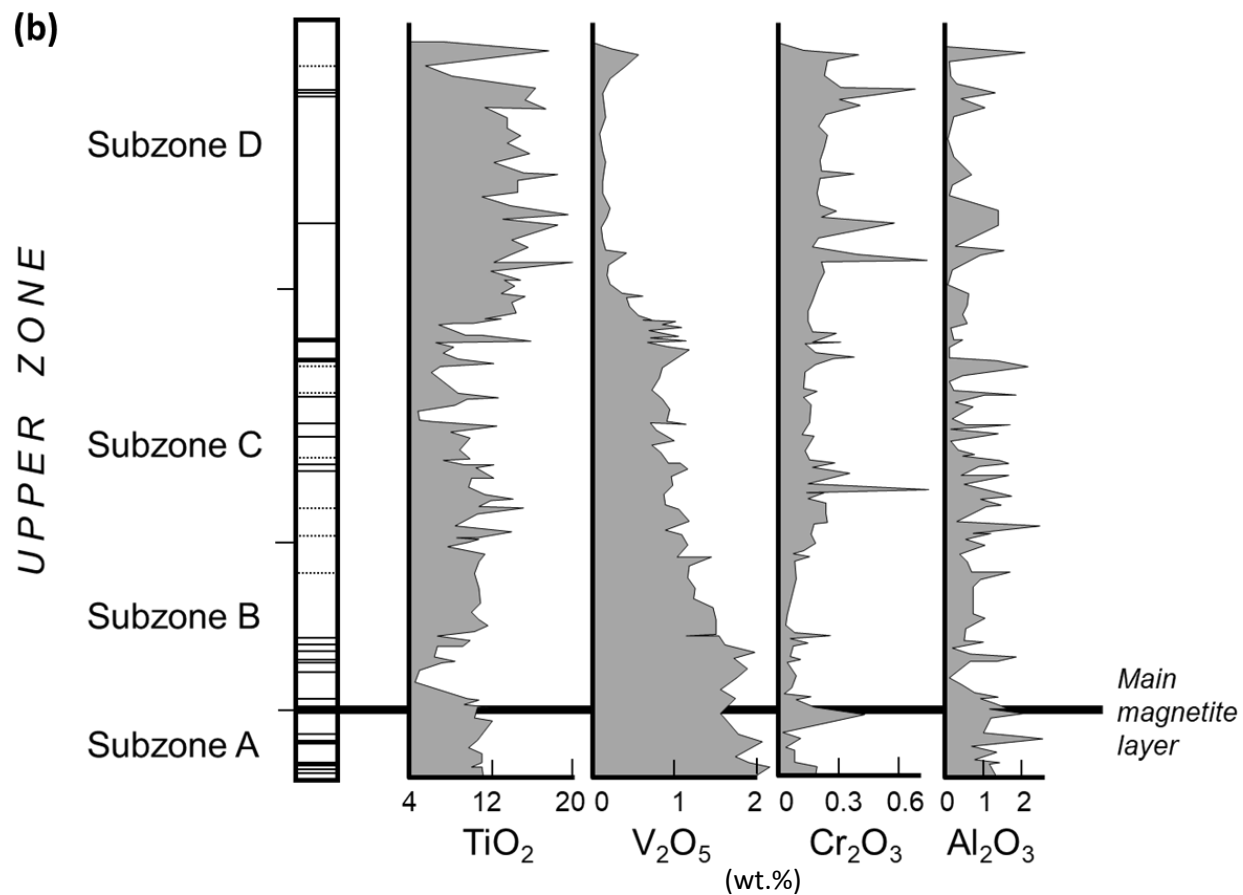
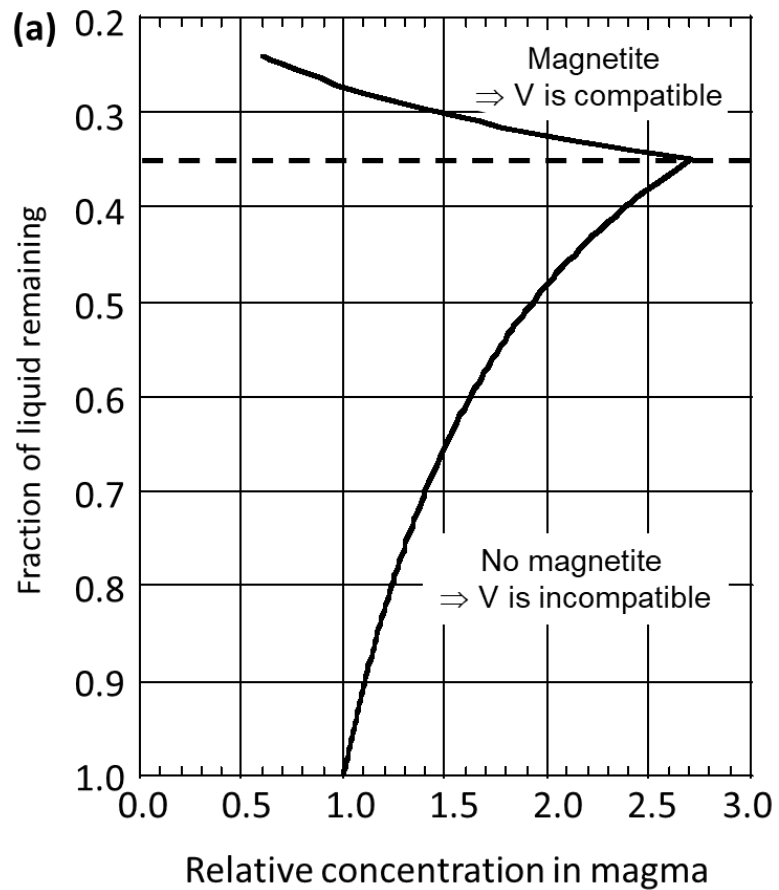


Figure 7

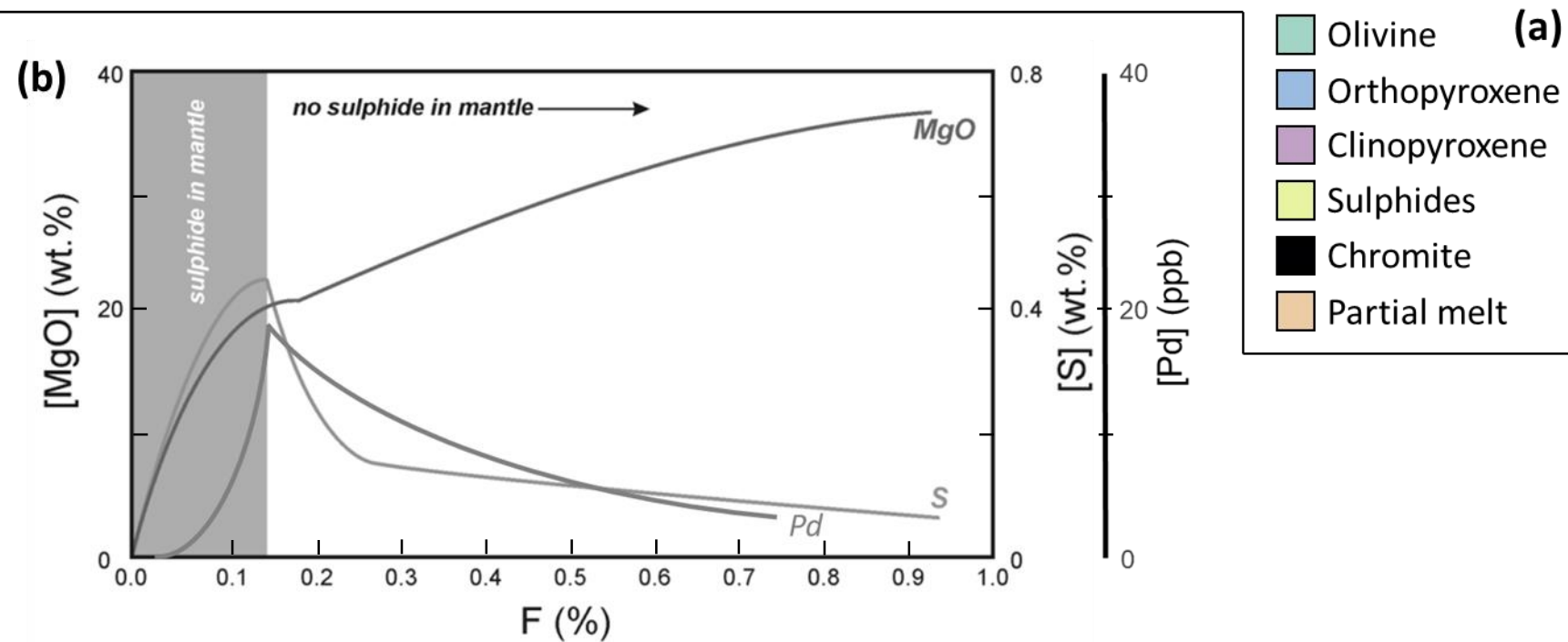
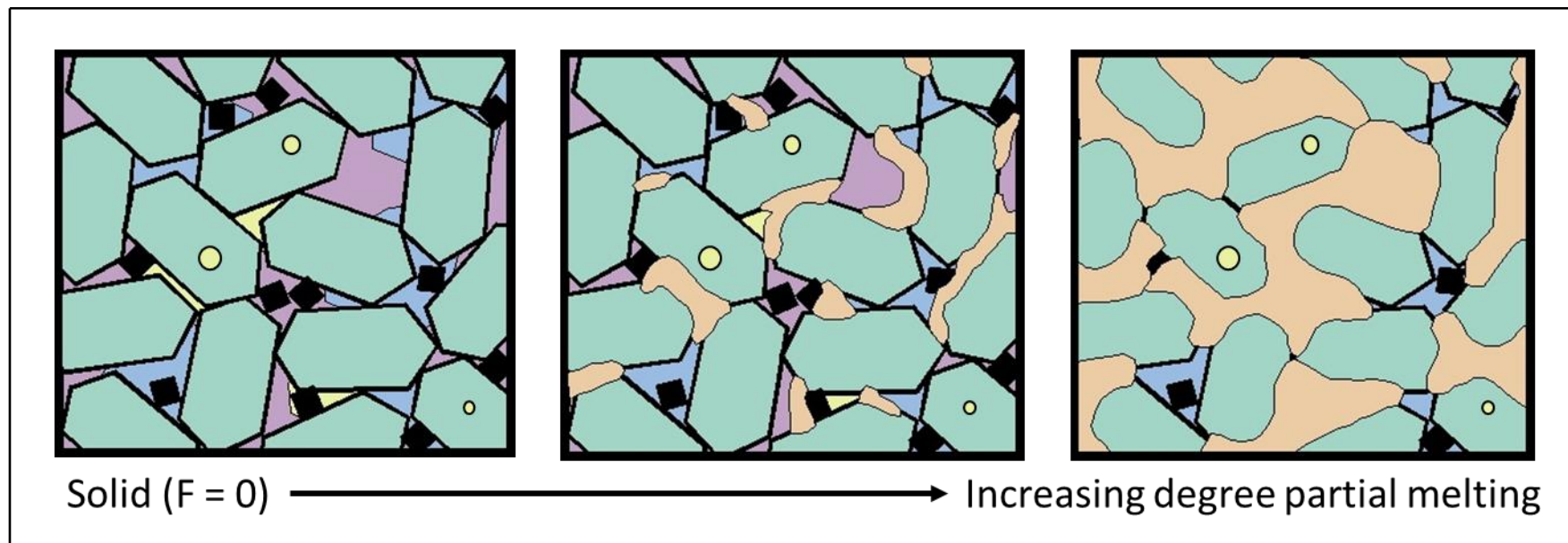


Figure 8

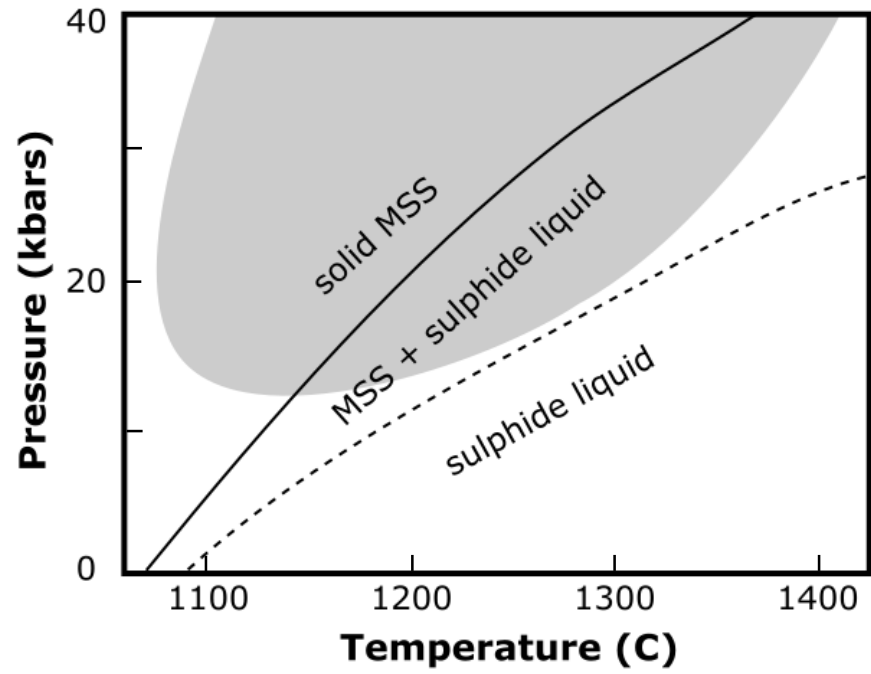
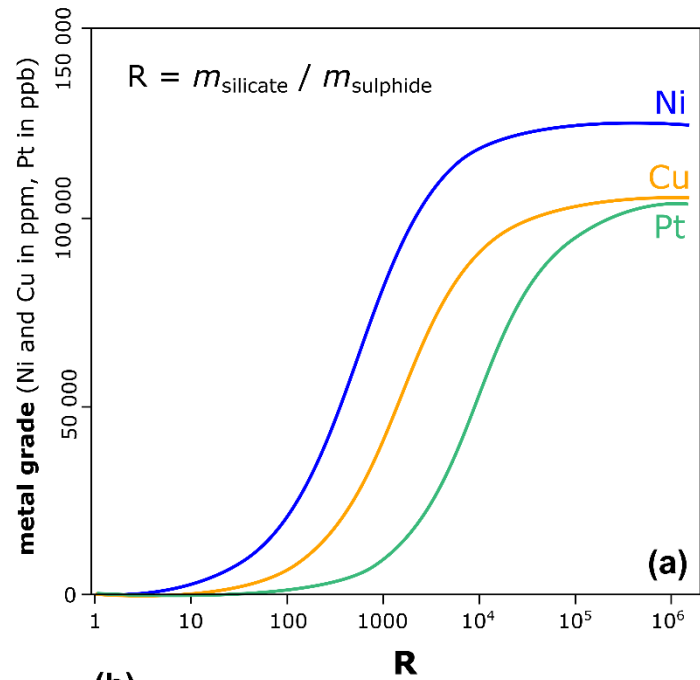
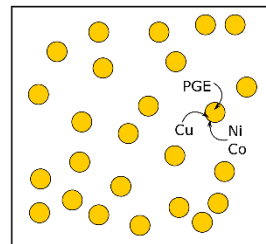


Figure 9

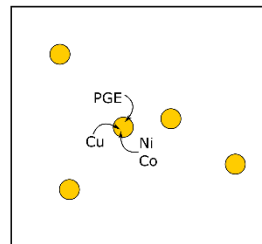


(b)



Low R

minimal uptake
of chalcophiles
(proportional to D)



High R

lots of uptake
of chalcophiles
(limited by circulation)

Figure 10

