Critical layer and radiative instabilities in shallow water shear flows

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In this study a linear stability analysis of shallow water flows is undertaken for a representative Froude number $F = 3.5$. The focus is on monotonic base flow profiles U without an inflexion point, in order to study critical layer instability (CLI) and its interaction with radiative instability (RI). First the dispersion relation is presented for the piecewise linear profile studied numerically by Satomura (1981) and using WKBJ analysis an interpretation given of mode branches, resonances and radiative instability. In particular surface gravity waves can resonate with a limit mode (or Rayleigh wave), localised near the discontinuity in shear in the flow; in this piecewise profile there is no critical layer.

The piecewise linear profile is then continuously modified in a family of nonlinear profiles, to show the effect of the vorticity gradient $Q' = -U''$ on the nature of the modes. Some modes remain as modes and others turn into quasi-modes, linked to Landau damping of disturbances to the flow, depending on the sign of the vorticity gradient at the critical point. Thus an interpretation of CLI for continuous profiles is given, as the remnant of the resonance with the limit mode. Numerical results and WKBJ analysis of CLI and RI for more general smooth profiles are provided. A link is made between growth rate formulae obtained by considering wave momentum and those found via the WKBJ approximation. Finally the competition between the stabilising effect of vorticity gradients in a critical layer and the destabilising effect of radiation (RI) is studied.

1. Introduction

Fluid mechanical phenomena in astrophysics and geosciences are a motivation for the study of shear instability in shallow, stratified, rotating fluid layers. Many types of waves and instabilities can occur, and the investigation of very idealised models is relevant to teasing out mechanisms and interactions, especially in linear regimes. In this paper we consider fluid flow governed by the shallow water (or Saint Venant) equations in a Cartesian geometry (without imposed rotation). Here a natural class of problems involves understanding the instabilities of a shear flow of constant depth bounded on one side by an impermeable boundary (here $y = 0$) and unbounded on the other side. A range of profiles $U(y)$ can be considered as models of possible flows adjacent to the boundary, for example modelling a boundary current in an ocean. A piecewise linear profile was studied numerically by Satomura (1981). In this case with an open domain, he obtained a family of surface gravity modes (localised near the boundary) and a single branch of modes linked to the discontinuity in the slope U' of the profile. He observed resonances

between these modes and the presence of radiative instability, in which an unstable mode incorporates waves that propagate to infinity, with outwards group velocity. For any piecewise linear shallow water flow an exact dispersion relation may be written in terms of Kummer functions, and this is done for the Satomura (1981) profile by Knessl & Keller (1995); these authors did not develop this further, and we note that these dispersion relations are sufficiently unwieldy that asymptotic approximations are needed to extract useful information. Glatzel (1985) discusses a related stability equation for a compressible shear flow, with Kelvin–Helmholtz instability resulting from resonances between modes localised at the two discontinuities of the profile.

When the profile $U(y)$ is no longer piecewise linear the possibilities for instabilities become richer, while exact solutions can no longer be written down. A key study is Balmforth (1999) for linear and nonlinear shear flow profiles confined to a channel; this author identifies three kinds of instabilities. First there is the classical inflexional instability (Rayleigh, 1880; Fjørtoft, 1950), present in the limit $F \to 0$ of incompressible fluid flow, when the surface gravity wave speed is infinite. Secondly, for $F > 0$ the wave speed becomes finite and such waves may then be destabilised if there is a critical layer where the wave speed and the flow speed are the same. In this case the behaviour of a mode is closely linked to the potential vorticity gradient in the layer: if its sign is the same as that of the wave momentum M the mode will be destabilised, otherwise it will be damped. The latter process is an example of Landau damping, first explained in the fluid context by Briggs *et al.* (1970), as we discuss further below. Finally, there is an unstable resonance between pairs of surface gravity modes, localised on opposite sides of the channel. For the flows considered in the present paper, this third instability is not present as the flow is bounded only on one side, but a fourth type of instability is allowed, a radiative instability with waves propagating to infinity, away from the wall. Finally for the discontinuous profiles there is the resonance instability obtained by Satomura.

Radiative instability was first found in compressible shear flows, in vortices by Broadbent & Moore (1979) and in many works by Lindzen (e.g., Lindzen & Tung, 1978; Lindzen & Barker, 1985) who first showed that the instability is based on over-reflection. Here waves are trapped and totally internally reflected; and at each reflection a wave packet draws energy from the underlying shear flow, while at the same time radiating a wave packet to infinity. Thereafter radiative modes were found in various open flows where both shear and stratification are present, for example in rotating flows, boundary and shear layers, and jets, in the presence of linear stratification, shallow water dynamics, and compressibility. Radiative instability is observed experimentally by Riedinger *et al.* (2011) in the case of the potential flow around a rotating cylinder in a stratified fluid. For a Froude number and a Reynolds number at the marginal stability limit, two networks of internal waves are generated, one corresponding to a helicoidal wave going down the cylinder and the other up. There are several ways of viewing the instability mechanism: it may be seen as an over-reflection process (Takehiro & Hayashi, 1992), as a consequence of the conservation of wave-activity or pseudomomentum (Schecter & Montgomery, 2004), or in terms of a reversal of the wave group velocity (Le Dizès & Billant, 2009).

A WKBJ analysis may be employed in these linear problems to obtain growth rates and to understand the over-reflection process and the role of the critical layer and turning points. Stratified vortices are considered in papers by Le Dizès and co-workers, for example Le Dizès & Billant (2009). For these flows the dominant term in the growth rate arises from radiation and a secondary damping term results from the critical layer. The competition between radiation and critical layer damping has also been studied in Schecter & Montgomery (2004) and Park & Billant (2012, 2013). In a recent work on

compressible jets (Parras & Le Dizès, 2010), a WKBJ analysis for radiative modes shows that the term in the growth rate resulting from the over-reflection at the critical level can also be destabilising.

Critical layers are better known for their stabilising effect, as for example explained in Briggs *et al.* (1970). Modes are found to be damped by a fluid–wave interaction similar to the Landau damping of plasma oscillations. In the fluid context this amounts to the generation of vorticity fluctuations in the presence of a background vorticity gradient, and the feedback on the mode as they are sheared out in the local flow. The mechanism is intimately linked to conservation of potential vorticity. Mathematically the resulting linear perturbation is not described by a normal mode and to obtain the decay rate the linear eigenvalue problem has to be integrated on a complex contour which cannot be deformed to the real axis. The corresponding eigenvalue is called a Landau pole, and is associated with a quasi-mode, which can be considered as formed from the continuous spectrum, in other words a combination of singular modes that naturally arises in the initial value problem (Briggs *et al.*, 1970). Quasi-modes always have a damping effect on perturbations, an effect studied for vortices, and nonlinear effects can lead to the formation of structures such cat's eyes or tripoles (e.g., Rossi *et al.*, 1997; Balmforth, 1999; Bassom & Gilbert, 1999; Schecter *et al.*, 2000; Balmforth *et al.*, 2001; Turner *et al.*, 2008), also observed in experiments (Van Heijst, 1991). Growth of unstable modes whose structure includes a critical layer has been observed in experiments of a columnar vortex in a stratified fluid (Riedinger *et al.*, 2010*b*). Little work has been published on the potential destablising effect of a critical layer, in particular for model geophysical flows, although it has been identified as the effect of a gradient in the background potential vorticity by Kubokawa (1985), Papaloizou & Pringle (1987) and Perkins & Renardy (1997). Otherwise the term 'critical layer instability', although it is not well established, has been used for baroclinic flows (Bretherton, 1966) and two layer flows (Iga, 1999).

The goal of the present paper is to study instabilities of shallow water shear flows, numerically and analytically, with particular interest in resonances, critical layer instability and radiative instability. The paper is organised as follows. In section 2 the governing equations are given, together with the general WKBJ formulation. The latter leads to a classification of types of modes, depending on the presence of critical points and turning points. Section 3 concerns the piecewise linear profile of Satomura (1981) (see also Knessl & Keller, 1995) which is the basis of all our subsequent analysis. We discuss the various modes and resonances, supported by WKBJ and related analysis. In particular we link a branch of 'limit' modes to the discontinuity in the piecewise profile, and give asymptotic formulae for these modes. In section 4 we consider a family of flows which includes the piecewise linear profile of Satomura (1981) but allows a quadratic profile, with non-vanishing vorticity gradient. This highlights the role of the critical layer and, as the profile is distorted from linear, numerical results indicate the effect on the branches and resonances. Importantly critical layer damping or destabilisation can now take place. Some previously neutral branches of modes remain as modes, while others turn into quasi-modes. To give an analytical basis to our results, the piecewise defined profiles are inconvenient, and in section 5 we discuss several smooth profiles. Some are bounded as $y \to \infty$ and some unbounded: in the latter case all modes become radiative at infinity. These profiles have neutral modes that may be stabilised or destabilised by potential vorticity gradients in the critical layer, and may be subject to radiative instability. Asymptotic formulae for these effects are derived using WKBJ theory and matching to local solutions near critical points, and confirmed by means of numerical calculations. Finally section 6 offers concluding discussion.

2. Governing equations

Our study concerns fluid motion governed by the shallow water equations, which we write in a standard dimensionless form,

$$\partial_t \boldsymbol{u} + \boldsymbol{u}\cdot\nabla\boldsymbol{u} + F^{-2}\nabla h = 0, \tag{2.1}$$

$$\partial_t h + \nabla\cdot(h\boldsymbol{u}) = 0, \tag{2.2}$$

where F is the Froude number based on the surface gravity wave velocity (with $F^2 = U^2/gH$ in dimensional quantities) (e.g. Balmforth, 1999). The shallow water dynamics gives material conservation of potential vorticity q,

$$\partial_t q + \boldsymbol{u}\cdot\nabla q = 0, \qquad q = h^{-1}\hat{\boldsymbol{z}}\cdot\nabla\times\boldsymbol{u}. \tag{2.3}$$

We consider an unperturbed, or basic, state of a steady shear flow with uniform depth,

$$\boldsymbol{u} = \boldsymbol{U} \equiv U(y)\hat{\boldsymbol{x}}, \quad h = H \equiv 1, \quad q = Q \equiv -U', \tag{2.4}$$

in a half plane given by $y \geq 0$ with an impermeable boundary $y = 0$. The velocity profiles we consider have $U(0) = 1$ and do not have inflection points, which rules out the inflectional instabilities discussed in Balmforth (1999) and Mak *et al.* (2013). Such instabilities are linked to the classical Rayleigh theory, since in the limit $F \to 0$ the shallow water equations reduce to those for incompressible two-dimensional fluid flow. We consider flows with different limits as $y \to \infty$, with $U \to 0$ for the flows in sections 2–5.2 and unbounded $U(y)$ in section 5.3.

We take the equations for small amplitude perturbations, make the substitutions

$$\boldsymbol{u} \to (U,0) + (u(y),v(y))e^{ik(x-ct)}, \quad h \to H + h(y)e^{ik(x-ct)}, \quad q \to Q + q(y)e^{ik(x-ct)}, \tag{2.5}$$

in (2.1,2.2), and linearise to obtain

$$ik(U-c)u + U'v + F^{-2}ikh = 0, \tag{2.6}$$

$$ik(U-c)v + F^{-2}\partial_y h = 0, \tag{2.7}$$

$$ik(U-c)h + iku + \partial_y v = 0 \tag{2.8}$$

(Balmforth, 1999). These govern normal modes of the perturbation fields u, v and h with wave number $k \geq 0$ in the x-direction and (possibly complex) wave speed $c = c_{\rm r} + ic_{\rm i}$. We use a prime to denote a y-derivative of the basic state (only). The perturbation potential vorticity q is given by

$$q = ikv - \partial_y u - Qh \tag{2.9}$$

and satisfies

$$ik(U-c)q + Q'v = 0. \tag{2.10}$$

The term $Q'v$ gives generation of perturbation potential vorticity in a background gradient $Q' = -U''$. For a purely linear shear flow, $Q' = -U'' = 0$, and this generation term is absent, making q identically zero.

A useful quantity is the wave momentum which in the full time dependent problem is defined by $M\exp(2kc_{\rm i}t)$, with

$$M = \tfrac{1}{2}\int_0^\infty (uh^* + hu^*)\,dy. \tag{2.11}$$

This is given in Balmforth (1999), together with the wave energy E which we do not need here. M is quadratic in the disturbance fields and its time evolution under (2.6–2.8) is

linked to the transport of perturbation potential vorticity through

$$2kc_{\rm i}M = -\tfrac{1}{2}\int_0^\infty (vq^* + qv^*)\,dy, \tag{2.12}$$

for modes that are evanescent as $y \to \infty$. For a purely linear profile, $U'' = 0$, q is zero and so M is conserved; in this case a growing or decaying mode, $c_{\rm i} \neq 0$, must have zero total momentum M. This is relevant to unstable resonances in bounded shear flows (Hayashi & Young, 1987; Takehiro & Hayashi, 1992).

2.1. *WKBJ formulation*

The key to understanding numerical results, as well as giving approximations to growth rates and frequencies, is a WKBJ analysis in the limit of large wave number $k \gg 1$. Equations (2.6–2.8) are a second order system of ODEs, which may be written in various forms. We first eliminate in favour of h to give

$$\partial_y^2 h - 2U'(U-c)^{-1}\partial_y h - k^2\Delta_0 h = 0. \tag{2.13}$$

Here Δ_0 is a function of y defined by

$$\Delta_0(y) = 1 - F^2(U-c)^2. \tag{2.14}$$

Although our numerical work is based on (2.13), for analysis in the limit of large k we eliminate the term in $\partial_y h$ by setting

$$g(y) = (U-c)^{-1}h, \tag{2.15}$$

to give

$$\partial_y^2 g = k^2\Delta\, g, \tag{2.16}$$

where the function $\Delta(y)$ is

$$\Delta(y) = 1 - F^2(U-c)^2 - k^{-2}U''(U-c)^{-1} + 2k^{-2}U'^2(U-c)^{-2}. \tag{2.17}$$

WKBJ approximations to (2.16) then involve $\Delta(y)$ in (2.17) in the standard form,

$$g(y) = |\Delta(y)|^{-1/4}\exp\left(\pm k\int^y \sqrt{\Delta(s)}\,ds\right). \tag{2.18}$$

With reference to (2.14) we may write the quantity $\Delta(y)$ as

$$\Delta = \Delta_0 + k^{-2}\Delta_1, \qquad \Delta_1 = -U''(U-c)^{-1} + 2U'^2(U-c)^{-2}. \tag{2.19}$$

For large k, the first term $\Delta_0(y)$ in (2.19) is nominally the largest. This gives rise to oscillatory solutions in regions where Δ_0 is negative and evanescent behaviour where Δ_0 is positive. Two adjacent regions are separated by a turning point $y_{\rm t}$ with

$$\Delta_0(y_{\rm t}) = 0, \tag{2.20}$$

and when a mode has two of these we label them as $y_{\rm t1}$ and $y_{\rm t2}$ with $y_{\rm t1} < y_{\rm t2}$. The other significant feature of (2.19) is the possible presence of a critical point $y_{\rm c}$ where the wave speed is equal to the flow velocity and the term involving Δ_1 can increase. Such a point is defined by

$$U(y_{\rm c}) = c. \tag{2.21}$$

Except close to a critical point $y_{\rm c}$, the terms in Δ_1 in (2.19) may be neglected compared

with Δ_0. However within a distance of order k^{-1} of the critical point, Δ_1 increases to become comparable with Δ_0 and a new expansion must be sought.

For some solution branches, modes are purely oscillatory with real wave speed c, in which case any relevant turning points y_t and critical point y_c will lie on the real y-axis. In other situations, where there is instability or damping, c becomes complex and these points lie in the complex plane. However despite this, our asymptotic calculations are based on $c = c_\mathrm{r} + ic_\mathrm{i}$ being approximately real ($c_\mathrm{r} = O(1)$, $c_\mathrm{i} \ll 1$). For this reason it is helpful to work on the basis that c is real together with real points y_t and y_c, all correct at leading order, and then calculate the small correction c_i perturbatively. In what follows we often speak as if y_t and y_c are real, even though these points may be 'pushed' above or below the real axis by small values of c_i. This approximation is valid, as the magnitude of c_i is found to be exponentially small in terms of k when WKBJ solutions are linked across an evanescent region of finite width (independent of k).

Given a profile $U(y)$, the problem then is to solve (2.13) or (2.16) subject to the boundary condition at the origin of no normal flow, $v = 0$, amounting to

$$h'(0) = 0 \quad \text{or} \quad g'(0)/g(0) = -U'(0)(U(0) - c)^{-1}, \tag{2.22}$$

and to the condition as $y \to \infty$ that

$$h(y) \to 0 \quad \text{or} \quad g(y) \to 0 \qquad (y \to \infty), \tag{2.23}$$

for evanescent modes, decaying at infinity. For radiative modes, that is modes which are oscillatory as $y \to \infty$, we need instead the radiation condition that waves propagate outwards, namely

$$h(y)(U - c)^{-1} = g(y) \sim |\Delta(y)|^{-1/4} \exp\left(\pm ik \int^y \sqrt{-\Delta(s)}\, ds\right), \tag{2.24}$$

with the upper/lower sign for positive/negative c_r. For each applicable set of boundary conditions there results an eigenvalue problem giving branches of modes $c(k)$ for a given value of the Froude number F.

As well as discrete normal modes, there are two branches of continuous spectrum, namely,

$$S_\mathrm{crit} = \{c : c = U(y), y \in \mathbb{R}\}, \tag{2.25}$$

$$S_\mathrm{rad} = \{c : \Delta_0(\infty) < 0\} = \{c : c > U(\infty) + F^{-1} \ \text{ or } \ c < U(\infty) - F^{-1}\}. \tag{2.26}$$

The first branch is linked to the presence of critical points y_c on the real axis where the differential equation is singular and the second is a range of values of c for which there are outward-going waves at great distances (cf. Riedinger *et al.*, 2010*a*). These branches correspond to an integration contour $y \in \mathbb{R}$ for (2.13) or (2.16), but they can be deformed in the complex plane by distorting this contour.

2.2. *Classes of modes*

For a given profile $U(y)$ there are several distinct classes of WKBJ solutions depending on the presence and location of turning and critical points. To illustrate this and establish notation, we take the smooth profile $U(y) = 1 - \tanh y$ (shown as U_2 in figure 2(b) below) with $F = 3.5$ as an example, and consider real values of $c = c_\mathrm{r}$, plotted on the vertical axis in figure 1. For each value of c we show the location of the turning points $y_{\mathrm{t}1}$, $y_{\mathrm{t}2}$ given by (2.20) (solid curves) on the horizontal axis, together with the critical point y_c from (2.21) (thin curve). In the shaded regions $\Delta_0 < 0$ and the WKBJ solution is oscillatory; otherwise the solutions are exponential.

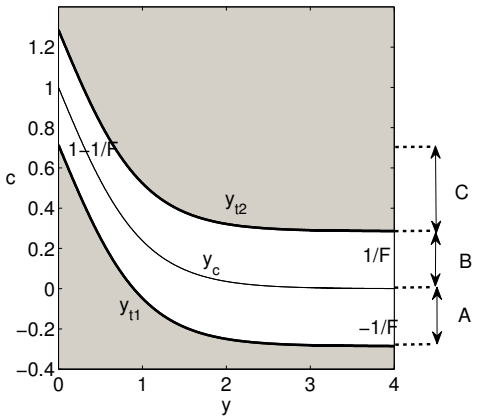


FIGURE 1. Location of turning and critical points, and sign of Δ_0, in the (y, c_r)–plane with $c = c_\mathrm{r}$ real, according to the profile $U_2 = 1 - \tanh y$ with $F = 3.5$. The shaded zones correspond to $\Delta_0 < 0$ and oscillatory modes. The turning points $y_{\mathrm{t}1}$, $y_{\mathrm{t}2}$ are marked by thick curves and the critical point y_c by a thin curve.

Given a value of c we may read horizontally to find turning and critical points, and so identify the following classes of modes:

- class A, $c_\mathrm{r} \in (-F^{-1}, 0)$: modes that are oscillatory for $[0, y_{\mathrm{t}1}]$, evanescent for larger values of y and with no critical point,
- class B, $c_\mathrm{r} \in (0, F^{-1})$: as for class (A) but with a critical point y_c in the open evanescent region,
- class C, $c_\mathrm{r} \in (F^{-1}, 1 - F^{-1})$: modes that are oscillatory in $[0, y_{\mathrm{t}1}]$ and for $y > y_{\mathrm{t}2}$ with an evanescent region in between; these are radiative modes.

From the figure, radiative modes in C exist when $F^{-1} < 1 - F^{-1}$, that is $F > 2$, and modes in class (B) exist when $1 - F^{-1} > 0$, that is $F > 1$. For the present study we fix a representative value of the Froude number $F = 3.5$ for all our simulations, as this gives all three classes of modes. The sketch of modes in figure 1 is similar to the one obtained by Parras & Le Dizès (2010) for a study of instability in a compressible round jet: class A modes correspond to counterflow waves, class B to subsonic coflow waves, and class C to supersonic coflow waves. Note that in the figure the characteristics of the modes are plotted according to the real part of the velocity c_r, but also constitute the leading order approximation when $c_\mathrm{i} \ll 1$ and modes are weakly damped or destabilised

3. Piecewise linear profile

The starting point for our study is the piecewise linear profile of Satomura (1981),

$$U(y) = \begin{cases} 1 - y & (0 \le y \le 1), \\ 0 & (y > 1). \end{cases} \tag{3.1}$$

This is the profile with $\mu = 0$ shown in figure 2(a). The piecewise linear property means that there is zero potential vorticity gradient $Q' = -U''$, except for a delta-function concentration at $y = y_\mathrm{d} = 1$. This allowed Knessl & Keller (1995) to write down an exact but awkward dispersion relation in terms of Kummer functions. We will not proceed this

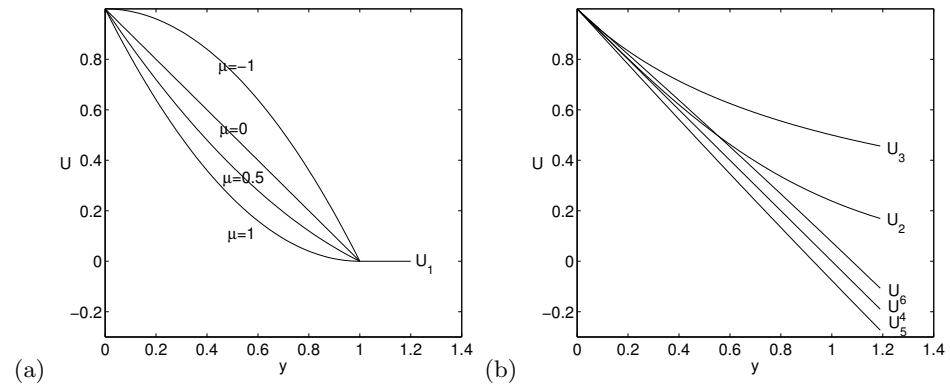


FIGURE 2. Basic flow profiles. Shown are (a) flow profiles U_1 that are zero for $y \geq 1$ and given by (4.1). The case $\mu = 0$ gives the piecewise linear profile U in (3.1). (b) Everywhere smooth profiles U_2 and U_3 in (5.1), U_4 in (5.19), U_5 in (5.20), and U_6 in (5.21).

way, but instead apply WKBJ approximations from the outset below; this allows easier generalisation to profiles that are not piecewise linear in later sections.

As $U = 0$ for $y > 1$ we may write g in (2.16) as a decaying exponential satisfying (2.23),

$$g \propto \exp\bigl(-ky\sqrt{1 - F^2c^2}\bigr), \tag{3.2}$$

or a wave with the appropriate sign for the radiation condition given in (2.24),

$$g \propto \exp\bigl(\pm iky\sqrt{F^2c^2 - 1}\bigr). \tag{3.3}$$

The solution is required to have h and h' continuous across $y = 1$; the latter makes the normal flow component v continuous there but there is generally a discontinuity in tangential flow u. We are then left with the problem on the reduced range $0 \leq y \leq 1$, to solve the differential equation for g in the case of constant shear,

$$\partial_y^2 g = k^2 \Delta\, g, \qquad \Delta = \Delta_0 + k^{-2}\Delta_1, \tag{3.4}$$

$$\Delta_0 = 1 - F^2(1 - y - c)^2, \qquad \Delta_1 = 2(1 - y - c)^{-2}, \tag{3.5}$$

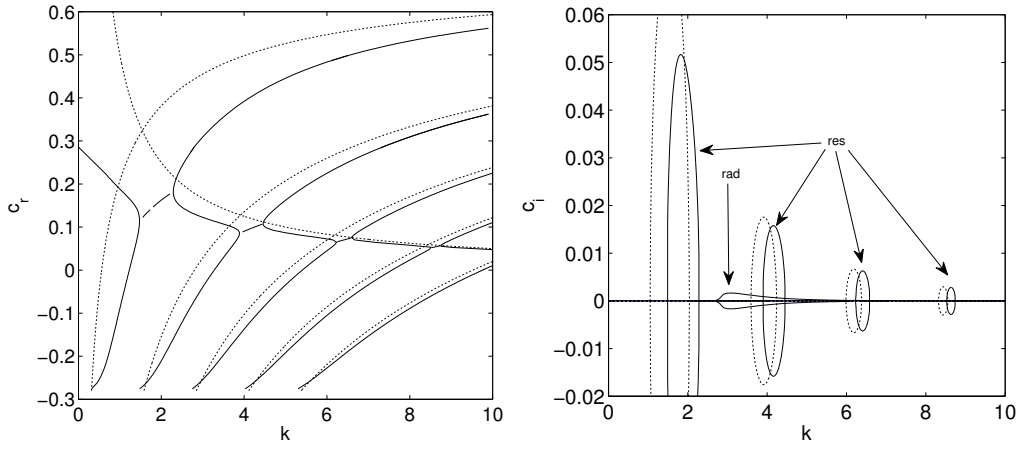
subject to

$$g'(0)/g(0) = (1 - c)^{-1}, \quad g'(1)/g(1) = -c^{-1} - k\sqrt{1 - F^2c^2} \text{ or } -c^{-1} \pm ik\sqrt{F^2c^2 - 1}. \tag{3.6}$$

3.1. *Numerical results*

The eigenvalue problem was solved using a shooting code for (2.13) and the eigenmode branches, that is c_r and c_i (solid curves), are shown in figure 3 as functions of k. When a mode is neutral, $c_\mathrm{i} = 0$, any critical point y_c lies on the real axis and makes the differential equation singular there. Although the singularity is easily treated analytically for this piecewise linear profile, the shooting has to be done on a path in the complex plane, for example a parabolic arc from zero to one above or below y_c.

This figure reproduces results of Satomura (1981) and shows a variety of eigenmode branches. We turn first to frequencies plotted in figure 3(a). Ignoring for a moment resonances near to branch crossings, we have two different types of branches. There is a single branch (somewhat broken up by resonances) with $dc_\mathrm{r}/dk < 0$, tending to the



(a) (b)

FIGURE 3. (a) Freqency c_r and (b) growth rate c_i for $F = 3.5$ and the piecewise linear profile (3.1). In each case the solid lines give the numerical solutions, and the dotted lines the WKBJ approximations. Radiative ('rad') and resonant ('res') instabilities are present.

horizontal axis as $k \to \infty$: we refer to these modes as 'limit modes'. These modes are localised on the discontinuity at $y = 1$ and continue to exist in the incompressible limit $F \to 0$; these can be interpreted as Rayleigh waves or edge modes on a piecewise linear profile (e.g., Sutherland, 2010). There is also a sequence of branches with increasing frequency as $k \to \infty$, $dc_\mathrm{r}/dk > 0$, and we refer to these as surface gravity wave modes, since branches show increasing numbers of oscillations in the fluid domain adjacent to the boundary $y = 0$. We observe all three classes of modes discussed at the end of section 2, leaving aside resonances at branch crossings.

First, for $c_\mathrm{r} \in (-F^{-1}, 0)$ the branches of surface gravity modes are all neutral and there is no critical point: these modes belong to class A. Secondly, for $c_\mathrm{r} \in (0, F^{-1})$ and away from branch crossings, the branches of surface gravity modes are again neutral, and so is the limit branch: these fall in class B. In this case there is a critical point y_c in the domain (with $0 \leq y_\mathrm{c} < 1$) but this does not lead to damping or destabilisation of the mode. This is a consequence of the piecewise linear profile: the background potential vorticity gradient Q' is zero at the critical point, and there is none of the feedback from vorticity transport in a critical layer. Linked to this, a Frobenius development of the equations (3.4, 3.5) near to the critical point y_c gives a simple pole in $g(y)$ and a non-singular height field $h(y)$ (Satomura, 1981). This means that numerical eigenvalues can be obtained equally from shooting above or below y_c in the complex plane.

For $c_\mathrm{r} \in (F^{-1}, 1 - F^{-1})$ the surface gravity modes are radiative and fall in class C. They have a positive or negative growth rate which pushes the critical point y_c off the real axis. In figure 3 (b) they correspond to long, low elongated 'bubbles' with growth rates of $O(10^{-3})$ or less; however only the first can be seen, clinging to the axis for $k > 2.7$, labelled 'rad'. For the growing radiative mode, the boundary condition (3.6) is used with the sign chosen to correspond to an outgoing wave as $y \to \infty$. However for the decaying mode it is an incoming wave with the opposite sign of the square root taken in (3.6). Thus it is only the upper portion of the bubble which corresponds to the physically important case of radiative instability.

The structure of three neutral modes, namely limit mode, third and fourth surface gravity

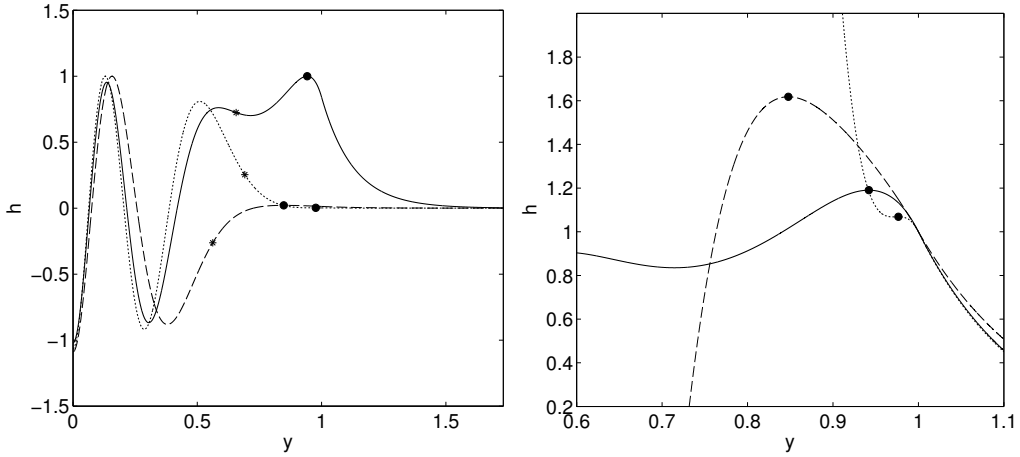


FIGURE 4. Structure of the three neutral modes for the piecewise linear profile (3.1) with $k = 8$, (a) scaled by their maximum amplitude, and (b) view zoomed in near the critical points with solutions scaled by their values at $y = 1$. The (real) $h(y)$ field is plotted against y for the limit mode with $c = 0.0581$ (solid), the third discrete mode for $c = 0.152$ (dash), and fourth discrete mode for $c = 0.0235$ (dot). The positions of critical points y_c ($\bullet$) are shown in (a,b) and turning points $y_t \equiv y_{t1}$ ($*$) in (a) only.

modes, is shown in figure 4 for $k = 8$. The curves were obtained by integrating along the real axis on both sides of the critical point, using the eigenvalue c first obtained by means of shooting along a complex path. Note that the surface gravity modes (dash and dot) are oscillatory up to their turning points y_{t1} ($*$) and then die away rapidly to very low amplitude at their critical points y_c ($\bullet$). On the other hand the limit mode (solid) has a peak at its critical point y_c; although it decays only a little for $y < y_c$, for large k this branch of solutions is represented at leading order by evanescent decay as y is reduced below y_c. We will pick up this structure in the WKBJ analysis below.

Where the branches for c_r cross in figure 3(a), the corresponding modes become resonant, giving the tall, narrow bubbles of instability labelled 'res' in figure 3(b). There are two solutions with positive or negative growth rate and each solution corresponds to a mixture of the two individual non-resonant modes. For these resonant bubbles, the two solutions with opposite growth rates can again be obtained by shooting above or below the critical point y_c, which is now pushed off the real axis.

3.2. *WKBJ theory for surface gravity modes*

We now turn to analysis of the features seen in figure 3, and in the remainder of this section are interested in modes in class B, with wave speed $c_r \in (0, F^{-1})$. Our aim is to describe limit and surface gravity modes using WKBJ analysis for large k (e.g., Bender & Orszag, 1978) and to give a description of the resonances that occur when branches cross. At the outset, in sections 3.2, 3.3 we seek the two separate families of surface gravity modes, localised near the boundary, and limit modes, localised near the discontinuity, as separate branches by requiring exponential decay in the evanescent region that separates them. We then investigate resonant interactions by connecting solutions through the evanescent region, giving effects exponentially small in k as $k \to \infty$, in section 3.4.

Supposing first that $c = c_r$ is real, we have a single real turning point y_t and a real critical point y_c, with $0 < y_t < y_c < 1$. We divide space into three regions, defined loosely as

region I, $0 \le y < y_{\rm t}$, region II $y_{\rm t} < y < y_{\rm c}$ and III $y \simeq y_{\rm c}$. We start in region I with the standard oscillatory form of the WKBJ solution in (2.18). For large k the solution satisfying the boundary condition at $y = 0$ in (3.6), which at leading order in k simply amounts to $g'(0) = O(1) \ll k$, is

$$g_{\rm I} = A(-\Delta)^{-1/4} \cos\Big(k \int_0^y \sqrt{-\Delta}\, dy\Big). \tag{3.7}$$

This is valid in region I, in which $\Delta \simeq \Delta_0 < 0$, and may be rewritten as

$$g_{\rm I} = A(-\Delta)^{-1/4} \sin\Big(k \int_y^{y_{\rm t}} \sqrt{-\Delta}\, dy - \Phi + \pi/4\Big), \tag{3.8}$$

with a phase defined by

$$\Phi = k \int_0^{y_{\rm t}} \sqrt{-\Delta}\, dy - \pi/4. \tag{3.9}$$

We move to region II where $\Delta_0 > 0$ and suppose that we are sufficiently far from the critical point that the WKBJ approximation remains valid. We may write this solution in the form

$$g_{\rm II} = \Delta^{-1/4}\Big[C \exp\Big(k \int_{y_{\rm t}}^y \sqrt{\Delta}\, dy\Big) + D \exp\Big(-k \int_{y_{\rm t}}^y \sqrt{\Delta}\, dy\Big)\Big]. \tag{3.10}$$

To find branches of solutions we now ignore the effect of the critical point and just require evanescent solutions in region II. In fact the presence of a critical point can have a weak destabilising effect, but only near to mode crossings as we will discuss below. A standard argument (e.g., Bender & Orszag, 1978) involves matching up the oscillatory solution (3.7) to the evanescent solution with $C = 0$ (via an Airy function approximation) and gives the leading order dispersion relation for these modes as $\Phi = n\pi$ or

$$k \int_0^{y_{\rm t}} \sqrt{-\Delta}\, dy = n\pi + \pi/4. \tag{3.11}$$

Away from any critical point we may approximate Δ by Δ_0 up to corrections of order k^{-2}. Evaluating the integral then gives the dispersion relation explicitly as

$$\tfrac{1}{2} k F^{-1} \Big[F(1-c)\sqrt{F^2(1-c)^2 - 1} - \cosh^{-1}(F(1-c))\Big] = n\pi + \pi/4. \tag{3.12}$$

This approximation (dot) gives the surface gravity branches depicted in figure 3(a) with excellent agreement for large k. The two surface gravity modes shown in figure 4 (dot and dash curves) have the correct qualitative structure, in particular the exponential decay beyond the turning points ($*$).

3.3. *Asymptotic theory for limit modes*

The branch of limit modes which approaches the horizontal axis of figure 3(a) has its origin elsewhere. As seen in figure 4 such a mode (solid curve) has a peak close to $y = 1$ and from figure 3, c is real (except at the mode crossings), positive and tends to zero for large k. Experimentation suggests that $c = c_{\rm r} = O(k^{-1})$ is the appropriate scaling and we take this as a working assumption. In this case a critical point, with $U(y_{\rm c}) = c$, given by $y_{\rm c} = 1 - c$ approaches $y = 1$ as k tends to infinity and we seek a eigenmode localised there. We call this region III, which is defined formally by $y - y_{\rm c} \ll k^{-1/2}$ and includes the point $y = 1$. In region III we can no longer ignore Δ_1 in (3.4, 3.5) but it is

now legitimate to approximate Δ_0, so that

$$\Delta \simeq 1 + k^{-2}\Delta_1 = 1 + 2k^{-2}(y - y_c)^{-2}, \tag{3.13}$$

and then the resulting differential equation

$$\partial_y^2 g = [k^2 + 2(y - y_c)^{-2}]g \tag{3.14}$$

transforms to a Whittaker equation for $w = 2k(y - y_c)$,

$$\partial_w^2 g + (-\tfrac{1}{4} - 2w^{-2})g = 0. \tag{3.15}$$

This equation is discussed in section 13.14 of Olver (2010) and in its standard form ((5.7) below) has two parameters which here take the values $\kappa = 0$, $\mu = 3/2$. The solution may be expressed as

$$g_{\mathrm{III}} = EW_{0,3/2}(w) + GW_{0,3/2}(-w), \tag{3.16}$$

or explicitly

$$g_{\mathrm{III}} = E[1 + k^{-1}(y - y_c)^{-1}]e^{-k(y-y_c)} + G[1 - k^{-1}(y - y_c)^{-1}]e^{k(y-y_c)}\,. \tag{3.17}$$

It may be checked that this solution matches to the exponential WKBJ solution (3.10) in region II. Note that the solution g_{III} has a simple pole at $y = y_c$ whereas the corresponding height field h (2.15) is regular there. We will later see that the situation becomes more complicated for profiles where $Q' = -U''$ is non-zero at the critical point and the Whittaker functions gain branch cuts.

We can now impose the boundary condition (3.6) at $y = 1$ on the solution (3.17), and we may neglect F^2c^2 on the right-hand side as $c = O(k^{-1})$, leaving us to require only $g'(1)/g(1) = -c^{-1} - k$. Some rearrangement gives

$$E/G = (1 - 2kc)e^{2kc}. \tag{3.18}$$

Finally for a solution localised about the critical point we require exponential decay as y decreases below y_c, in other words $E = 0$, leaving the branch described by the approximation

$$c = (2k)^{-1}, \tag{3.19}$$

valid for large k, independent of Froude number. This approximation is shown in figure 3(a) (dot) and shows good agreement with increasing k. These limit modes have a peak localised close to $y = 1$ for large k, as illustrated in figure 4 (solid), and so can be thought of as driven by the jump in potential vorticity Q at $y = 1$, much like a normal mode on a Rankine vortex in the analogous problem in plane polar geometry (Ford, 1994).

3.4. *WKBJ theory for resonances*

We have identified a set of branches of surface gravity modes, trapped between $y = 0$ and $y = y_t$ and a branch of limit modes, localised at a critical point y_c close to the boundary $y = 1$ of the shear flow. We observe resonances between these modes where the branches cross in figure 3, and now sketch the analysis of these following Knessl & Keller (1995). This involves keeping track of both exponential components of the evanescent WKBJ solution (3.10) in $y_t < y < y_c$, even though we neglect terms of order k^{-1} in each one; we need to keep the two independent solutions at leading order (no matter how weak one is) to capture an effect that is exponentially small as $k \to \infty$ (Shepard, 1983).

Matching the WKBJ solutions (3.8) and (3.10) across the turning point y_t gives

$$C/D = -2\tan\Phi \tag{3.20}$$

and matching the solution (3.17) for $y < y_c$ to (3.10) for $y > y_t$ yields (omitting details),

$$G = C\exp(\tfrac{1}{4}\pi k F^{-1}), \quad E = D\exp(-\tfrac{1}{4}\pi k F^{-1}). \tag{3.21}$$

Using also (3.18), the resulting dispersion relation is found to be

$$(kc - \tfrac{1}{2})\tan\Phi = \tfrac{1}{4}\exp(-\tfrac{1}{2}\pi k F^{-1} - 2kc). \tag{3.22}$$

The right-hand side is exponentially small for large k and if neglected, we regain the surface gravity and limit branches we already have, namely (3.11) and (3.19). We work near to a branch crossing, labelled by n, where simultaneously both conditions are satisfied

$$\Phi_n = n\pi, \quad k_n c_n = \tfrac{1}{2}, \tag{3.23}$$

and set $c = c_n + \delta c$, $k = k_n + \delta k$. The dispersion relation (3.22) takes the form

$$(\delta c + \alpha\,\delta k)(\delta c + \beta\,\delta k) + \gamma = 0 \tag{3.24}$$

to leading order in δc, δk, where the coefficients are

$$\alpha = c_n k_n^{-1} = \tfrac{1}{2}k_n^{-2}, \quad \beta = -k_n^{-2}(F^2-1)^{-1/2}(n\pi + \pi/4), \tag{3.25}$$

$$\gamma = \tfrac{1}{4}k_n^{-2}(F^2-1)^{-1/2}\exp(-\tfrac{1}{2}\pi k_n F^{-1} - 1). \tag{3.26}$$

The approximate dispersion relation (3.24) is solved as

$$2\delta c = -(\alpha+\beta)\delta k \pm \sqrt{(\alpha-\beta)^2\delta k^2 - 4\gamma}\,. \tag{3.27}$$

Now $\gamma > 0$ as $F > 1$ for class B modes (see figure 1) and so for small δk we obtain a pair of complex roots and instability. We have an interval about k_n given by $|\delta k| < 2\sqrt{\gamma}|\alpha-\beta|^{-1}$ in which we have complex roots. The maximum growth rate will occur at $\delta k = 0$ and is given by $\delta c = \pm i\sqrt{\gamma}$. The theoretical calculation is plotted on figure 3(b) (dot), with the resonance bubbles centred on the crossings (k_n, c_n) of the approximate branches in (3.23). We observe improving agreement as k increases.

The resonant interaction presented in this part is very similar to the resonant instability of a purely linear shear flow, with $Q' = -U'' = 0$, between two walls. In this case two neutral surface gravity modes, with the same frequencies and localised on opposite walls interact (Balmforth, 1999). The resulting instability has been explained as a coupling between two surface gravity waves with opposite signs of wave action, energy or momentum (Hayashi & Young, 1987), or as an over-reflection process, or as two waves propagating energy in opposite directions (as they have opposite group velocities) with the region of the critical point acting as an energy source (Le Dizès & Billant, 2009).

To link this case of a linear profile with the *piecewise* linear profile of this section and Satomura (1981), note first that the existence of the limit mode is a result of the discontinuity of Q at $y = 1$. Its existence as a neutral mode is also linked to the finite limit $U(\infty)$. If $U(y)$ is unbounded as $y \to \infty$ then all modes radiate at infinity, and none of the modes would be neutral. Secondly arguments based on coupling of modes with opposite signs of wave momentum cannot be applied here. The discontinuity in Q also gives a contribution to the rate of change of M in (2.12) through a delta function source of q. Thus M is no longer conserved and wave momentum can now be extracted from the background shear flow. Whereas for a strictly linear profile, with $Q = 0$ everywhere, any unstable mode must have zero M and so can only arise from the resonance between modes with positive and negative M, for the piecewise linear profile this is no longer the case. Both surface gravity modes and limit mode have negative M, as discussed below.

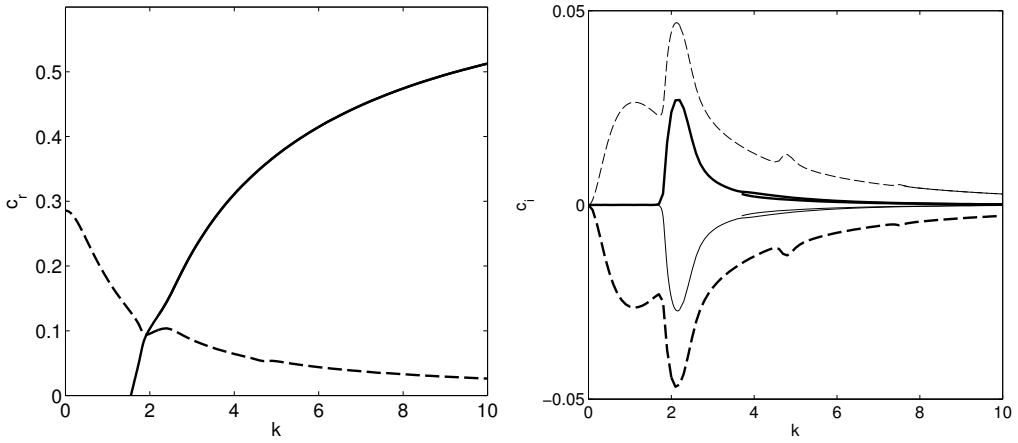


FIGURE 5. First surface gravity mode SGM (thick solid) and limit quasi-mode LQM (thick dash) for the profile U_1 in (4.1) with $\mu = 0.5$. Depicted are (a) c_r, and (b) c_i. The thin curves in (b) result from integrating below the critical point y_c in the complex y-plane.

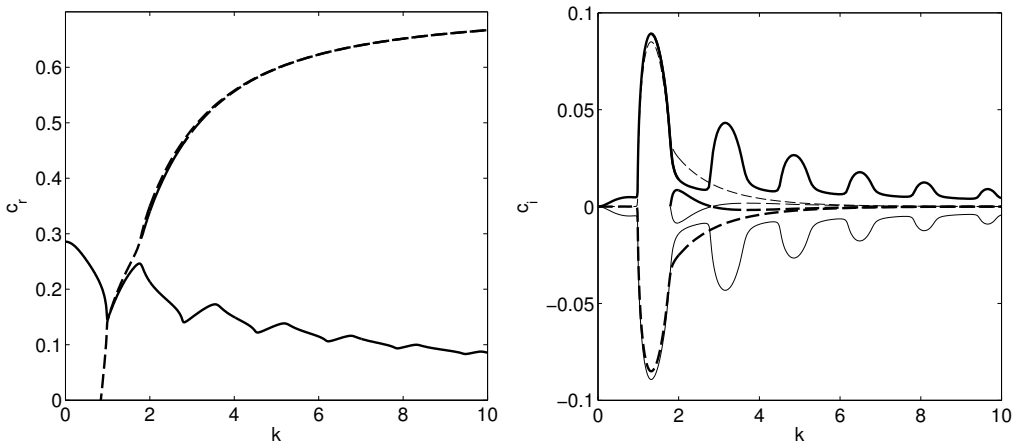


FIGURE 6. First surface gravity quasi-mode SGQM (thick dash) and limit mode LM (thick solid) for the profile U_1 in (4.1) with $\mu = -1$. Depicted are (a) c_r, and (b) c_i. The thin curves in (b) result from integrating below the critical point y_c in the complex y-plane. The radiative mode for $c_\mathrm{r} \in (0.28, 0.47)$ becomes a quasi-mode for $c_\mathrm{r} \in (0.47, 0.71)$

4. Nonlinear profiles

In our analysis of the limit modes in the piecewise linear profiles in the previous section, the gradient of potential vorticity at the critical point is zero, and this is a key simplification. In this section we allow nonlinear shear in the region $0 < y < 1$, using members of the family of profiles,

$$U_1(y) = \begin{cases} (1-\mu y)(1-y) & (0 \le y \le 1), \\ 0 & (y > 1), \end{cases} \tag{4.1}$$

depicted in figure 2(a). The parameter μ gives the potential vorticity gradient and the curvature of the profile with $U'' = -Q' = 2\mu$, and $\mu = 0$ is the previous piecewise linear case. We consider how frequencies c_r and growth rates c_i are changed as $|\mu|$ is increased from zero. Note that the situation has some similarities with that discussed by Balmforth

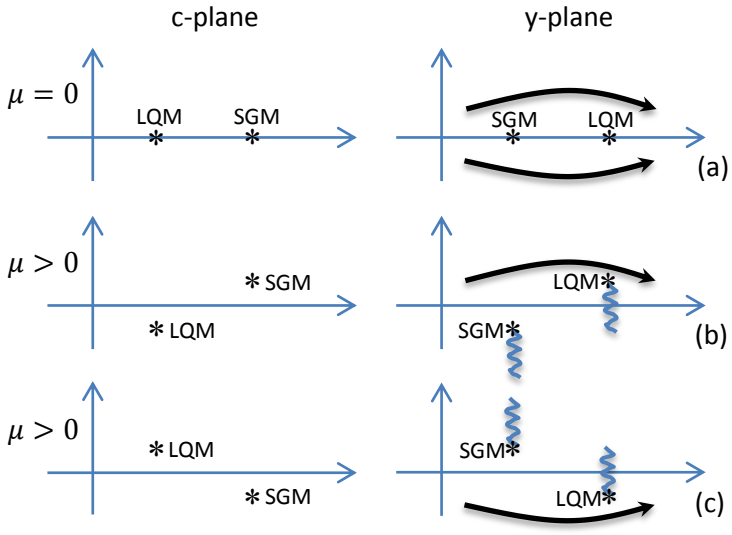


FIGURE 7. Schematic picture of eigenvalues in the complex c-plane (left panels) and critical points y_c in the complex y-plane (right panels) for surface gravity (SG) and limit (L) modes (M) and quasi-modes (QM). In (a) $\mu = 0$, in (b) $\mu > 0$ with y-contour taken above y_c, and in (c) $\mu > 0$, with contour taken below y_c. Branch cuts are also indicated.

(1999) for a finite channel, but in that case the symmetry of the channel increases the number of modes. We focus on just the first surface gravity mode and the limit mode; further branches of surface gravity modes show similar behaviour.

Referring to the classification of section 2.2, modes in class A do not have a critical point and so remain neutral. However this slight change in the profile results in significant effects on modes in classes B and C. First the critical point is slightly moved in the complex plane. For radiative modes already $c_i > 0$ and so this is just a modification to the growth rate. However for the previously neutral modes there is the appearance of a small imaginary part c_i, and this is the origin of critical layer instability.

Secondly, the eigenvalue problem now has different solutions depending on whether the path of integration is taken above or below a critical point of the governing differential equation in the complex plane. When $\mu \neq 0$, the U'' term of Δ_1 in equation (2.19) is non-zero, and the height field h at the critical point is singular, generally gaining a branch cut. For this reason the two branches of solutions for $\mu = 0$ in figure 3(b) become four branches, shown in figure 5 for $\mu > 0$ and figure 6 for $\mu < 0$. Solid lines give the modes and dashed lines the quasi-modes. Thick lines correspond to integrating above the critical point and thin lines below, giving complex conjugate values of c and so stable/unstable pairs. The resonant bubbles from figure 3 leave a clear imprint in each case, but to explain the overall structure it is more helpful to focus on a case in between such bubbles, where the original $\mu = 0$ modes are neutral in figure 3.

The effect of introducing a background potential vorticity gradient is then shown schematically for class B modes in figure 7 which indicates the locations of eigenvalues for c and corresponding values of y_c in the complex c-plane (left panels) and complex y-plane (right panels). In 7(a) the two neutral modes (LM, SGM) are depicted for $\mu = 0$ (away from resonance and with, say, $k > 2$ for definiteness). Now for $\mu > 0$ an integration path is taken well (b) above or (c) below the complex y-axis and the two neutral modes in

(a) move off the real axis; the sense in which they are tipped off is opposite in c- and y-planes since $U'(y) \leq 0$ for this and other flows we consider. Because the solutions gain a branch point at y_c and a branch cut in the y-plane, the neutral surface gravity mode in (a) becomes a pair of normal modes (SGM) in (b,c), but the neutral limit mode becomes a pair of limit quasi-modes (LQM). For $\mu < 0$, the situation is reversed and we gain a pair of surface gravity quasi-modes (SGQM) and a limit mode (LM) (not depicted). The two modes and the two quasi-modes have the same value for c_r but have opposite growth rates: see figure 5 for positive μ, and figure 6 for negative μ. Note that in general a solution is called a quasi-mode when the critical point y_c lies between the real axis and the complex y-path used to obtain it; an attempt to distort the contour to the real axis would encounter the branch cut and so give a discontinuous solution for the physical fields (Briggs *et al.*, 1970). This discontinuity would signal the need for a critical layer, in which the vorticity field is sheared to small scales, for a complete description of the time evolution. For a normal mode there is no barrier (i.e. critical point) between distorted contour and the real axis.

We make some further remarks. First the situation is complicated because of the presence of the two branches of continuous spectrum noted in (2.25, 2.26). When a neutral mode for $\mu = 0$ turns into two quasi-modes, these are really on separate Riemann surfaces of the dispersion relation, only apparent by either going through a branch cut of continuous spectrum, or by distorting the y-contour so as to shift this branch cut and so reveal the quasi-mode (Briggs *et al.*, 1970). Related to this is the issue of causality. In our discusson so far, both branches of quasi-modes are given on an equal footing, as in Balmforth (1999), to show how solutions of the eigenvalue problem are modified because of the nonlinearity of the profile and how modes are connected to quasi-modes as the parameter μ is taken to pass through zero. However considering an initial value problem using a Laplace transform setting (Briggs *et al.*, 1970; Schecter & Montgomery, 2004), requires a inversion contour sitting above any singularities in the c-plane. Then, to reveal a quasi-mode on another Riemann sheet, it is necessary to *raise* the y-contour to as to *lower* the continuous spectrum in the c-plane (in our flows with $U' \leq 0$), as depicted in figure 7. Note that as a Landau pole can be considered as a component of the continuous spectrum, it dominates the full evolution during a transient period before being overtaken by other components that decay only algebraically. This period increases the closer the pole is to the c_r-axis, i.e. the less the quasi-mode is damped; however information about Landau poles is relevant for the evolution of disturbances and processes such as cat's eye formation and mixing, even when the quasi-mode is not weakly damped (Schecter *et al.*, 2000; Turner *et al.*, 2008).

In any case, only the damped quasi-mode is relevant to the intial value problem and we can discard quasi-modes obtained with a path taken below the y-axis (for example the thin, dashed curves in figures 5(b) and 6(b). While quasi-mode damping can be considered as a result of fine structure being sheared to finer and finer scales in a critical layer, the possibility of quasi-mode amplification would correspond to indefinitely fine structure being unsheared in the critical layer, at odds with the nature of a (smooth) initial condition. Similar considerations apply to the robustness of modes in the presence of weak viscosity (Lin, 1945; Balmforth, 1999), and we have confirmed this by time-stepping (2.1) (with an additional viscous term). From here onwards we only consider the physical solutions obtained with the path taken above critical points in the y-plane (thick curves). With this, for class B modes, $c_r \in (0, F^{-1})$, the mode and the quasi-mode have independent dispersion relations: profiles with $\mu > 0$ have amplified surface gravity modes, and damped limit quasi-modes; for $\mu < 0$ it is the opposite way round.

Introducing nonlinearity $\mu \neq 0$ has a similar effect on radiative modes as on a surface gravity mode: $\mu > 0$ leads to amplification and $\mu < 0$ leads to damping. Note that, as for class B modes, there are again two solutions for a given c_r in the interval of class C modes. In fact the solution for $c_\mathrm{i} > 0$ is obtained with the boundary condition of an outgoing wave and the other with $c_\mathrm{i} < 0$ with an incoming wave. For small μ the growing outgoing solution is a mode and the damped incoming solution a quasi-mode. If μ is large, positive or negative, the sign of c_i can change and hence the mode change to quasi-mode or conversely. This can be seen in figure 5 and in figure 6 for $c_\mathrm{r} > 0.47$. Note that in figure 5 the growing mode associated with an incoming wave boundary condition is not physically relevant.

We can summarize the effect of weak nonlinearity of the profile according to the range of wave speed c_r. For class B, neutral modes become a mode/quasi-mode pair with two very different structures and frequencies. For class C, radiative modes become a mode/quasi-mode pair with very close frequencies corresponding to a growing radiative mode with an outgoing wave and a damped quasi-mode with an incoming wave. An approximate formula for critical layer instability of a smooth flow may be obtained by integrating (2.12) along the real y-axis (Kubokawa, 1985; Balmforth, 1999) linking the growth rate c_i to the wave momentum M and the characteristics of the flow at the critical point,

$$|c_\mathrm{i}| = \frac{\pi Q_c' |v_c|^2}{2k^2 M |U_\mathrm{c}'|} \,. \tag{4.2}$$

These modes split into growing/decaying pairs or disappear entirely depending on whether Q_c'/M is positive or negative. When the modes disappear, for $Q_c'/M < 0$, they become quasi-modes and the left hand side of (4.2) is to be replaced by $-|c_\mathrm{i}|$, from integrating on the other side of the critical point in the y-plane (Balmforth, 1999). In fact for the profiles we study, the limit and surface gravity mode both have both negative momentum M and the same sign for $Q_\mathrm{c}' = -2\mu$. However they appear and disappear for opposite values of μ, in apparent contradiction with (4.2). The reason is that this formula does not apply to the profiles in (4.1) because U' is discontinuous at $y = 1$. Incorporating a term $Q'(y) = Q_\mathrm{d}\delta(y - y_\mathrm{d})$ in the vorticity gradient (see appendix A) gives instead

$$|c_\mathrm{i}| = \frac{\pi Q_c' |v_c|^2}{2k^2 |U_c'|} \left(M - \frac{Q_\mathrm{d} |v_\mathrm{d}|^2}{2k^2 c_\mathrm{r}^2} \right)^{-1} , \tag{4.3}$$

and predictions from this formula are in line with our results. So, for example the $\mu = 0$ surface gravity mode is weak near the discontinuity (i.e., in the evanescent region), has $|v_\mathrm{d}|$ small, and so (4.3) becomes the same as (4.2), this mode disappearing for $\mu < 0$. For the limit mode, localised near the discontinuity, $|v_\mathrm{d}|$ is larger giving a key sign change between the right-hand sides of (4.2) and (4.3).

5. Smooth profiles

We have determined the structure and growth rates of disturbances to the piecewise linear profile (3.1) of Satomura (1981), and how introducing curvature to the flow profile leads to critical layer instability and the creation of quasi-modes. However (unless $\mu = 1$) these profiles have a discontinuity in derivative $U'(y)$ at $y = 1$ which has two implications. First the limit mode discussed above and given explicitly by (3.17) for $\mu = 0$ is localised near to the discontinuity of $U'(y)$. Smoothing the profile will have a big impact on this mode and it may disappear entirely. Secondly, it is is difficult to present an analysis because of having to impose the boundary condition (3.6) at $y = 1$: this point sits in the

region where the solution is described by Whittaker functions and imposing a boundary condition here does not lead to useful or explicit formulae. For these reasons we consider profiles that have continuous $U'(y)$ in this section. The focus is on profiles with finite limiting velocity $U(\infty)$ as $y \to \infty$ in section 5.1 and unbounded profiles, for which all modes are radiative, in section 5.3.

5.1. *Profiles with bounded velocity*

We consider the piecewise profile $U_1(y)$ in (4.1) with $\mu = 1$, and the smooth profiles,

$$U_2(y) = 1 - \tanh y, \qquad U_3(y) = (1+y)^{-1}. \tag{5.1}$$

U_1, U_2 and U_3 tend to zero increasingly slowly for large y; see figure 2. As in Riedinger *et al.* (2010*a*) results were obtained using the spectral code on a complex path defined by the variable $\arg y = \pi/10$. Results were checked afterwards using a shooting code. The phase velocity c_r and growth rate c_i are given for the unstable mode and the limit quasi-mode in figure 8 for the three profiles. Using the spectral code on this complex path has the effect of lowering both branches of continuous spectrum (2.25,2.26) in the c-plane, and revealing quasi-mode eigenvalues (Riedinger *et al.*, 2010*a*). The integration path is always taken above critical points, so imposing causality or effects of weak viscosity.

Looking first to the curves for the piecewise profile U_1 in figure 8 (which can also be compared with figure 5 for $\mu = 0.5$) the gravity wave mode (solid, thickest) becomes unstable as soon as c_r becomes positive, when the critical layer is present for $k > 1.99$. The maximum $c_\mathrm{i} = 0.0157$ is obtained for $k \simeq 2.8$ a little after the crossing of c_r branches in 8(a) at $k \simeq 2.475$. The c_i branch for the limit quasi-mode in 8(b) (dash, thickest) shows a similar but inverted bump, i.e. a trough, for these values of k, corresponding to increased damping. Returning to the surface gravity mode, there is now no obvious distinction between this and the resonance which we had in figure 3(b) and this allows us to draw a new interpretation of the critical layer instability in these smoother profiles with $U'' \neq 0$ as the remnant or 'ghost' of the interaction between the surface gravity mode and the limit mode in the piecewise linear profile. Informally, in the latter case all of the U'' is concentrated in a delta function at $y = 1$, and the unstable resonance with the resulting limit mode is analogous to the critical layer instability when the non-zero U'' is distributed, over the critical layer.

The U_2 and U_3 profiles are entirely smooth, but tend to zero exponentially and algebraically, respectively. For U_2 the curves for c_i (thick) in figure 8(a) are similar to those for U_1, but the growth rate for the limit quasi-mode 8(b) (thick, dash) becomes increasingly negative as k is increased, and so it is strongly damped. The surface gravity mode (thick, solid) is amplified, but more weakly now. For U_3, these effects are more pronounced: the surface gravity mode is amplified (thin, solid) but it cannot be seen on the figure as the maximum of c_i is 1.61×10^{-4}. We have not been able to obtain the quasi-mode for U_3, indicated by the absence of thin dashed curves in the figure; indeed the critical point that must be circumnavigated is expected to be far above the real axis. In summary, the smoother the profile, i.e. the smaller the values of U'', the lower is the critical layer instability growth rate for the surface gravity modes and the stronger is the damping of the limit quasi-mode, to the point where it cannot easily be detected numerically.

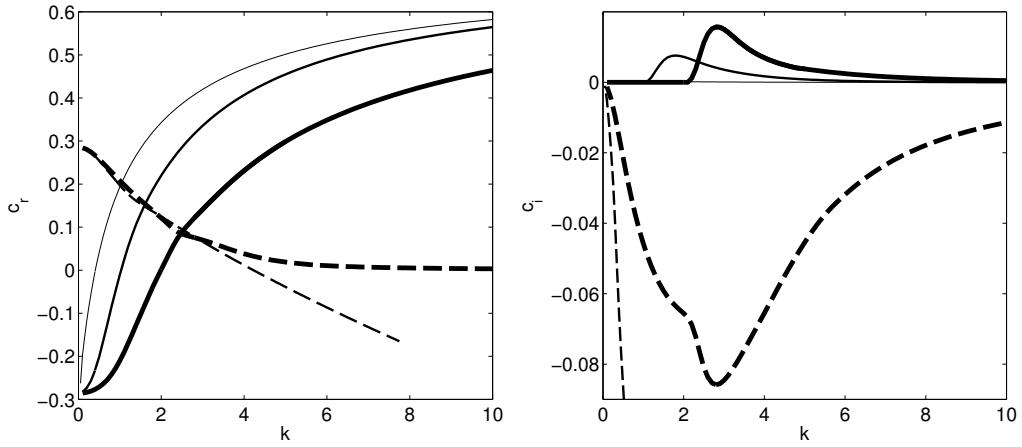


FIGURE 8. Critical layer instability: (a) c_r and (b) c_i for first surface gravity mode (SGM) (solid) and limit quasi-mode (LQM) (dash) for $F = 3.5$ and profiles U_1, $\mu = 1$ in (4.1) (thickest), U_2 (thick) and U_3 (thin) in (5.1).

5.2. *WKBJ theory for critical layer instability*

We now develop a theoretical framework for the above numerical results in the case of smooth profiles. We work on $0 \le y \le \infty$ and return to the full set of equations (2.16–2.19) for any smooth profile $U(y)$, with boundary conditions (2.23) and (2.24). We assume we have a single turning point y_t and a critical point y_c with $0 < y_\mathrm{t} < y_\mathrm{c} < \infty$. The difficulty for a nonlinear profile is the more complicated form of Δ_1 in (2.19) which generally gives a branch cut in the solution at a critical point y_c if U'' is non-zero there, i.e. $Q'_\mathrm{c} = -U''_\mathrm{c} \neq 0$.

In line with our discussion in section 3.2, we specify region I as $0 \le y < y_\mathrm{t}$, region II as $y_\mathrm{t} < y < y_\mathrm{c}$, region III as a neighbourhood of y_c and region IV as $y > y_\mathrm{c}$. In regions I and II we have the WKBJ solutions (3.7) and (3.10) exactly as before and these are connected through the turning point via (3.20). In region IV we have an evanescent wave

$$g_\mathrm{IV} \propto \Delta^{-1/4} \exp\Big(-k \int^y \sqrt{\Delta}\, dy\Big). \tag{5.2}$$

The key problem is to link up the solutions in regions IV and II across the critical point. As before we consider region III, defined formally by $|y - y_\mathrm{c}| \ll k^{-1/2}$, and approximate

$$\Delta \simeq 1 + k^{-2}\Delta_1. \tag{5.3}$$

Using the smoothness of the profile and a Taylor series expansion, the term Δ_1 has the following singular component as y_c as approached,

$$\Delta_1 = 2\tilde{y}^{-2} + (U''_\mathrm{c}/U'_\mathrm{c})\tilde{y}^{-1} + \cdots, \tag{5.4}$$

where we have set $\tilde{y} = y - y_\mathrm{c}$ for convenience. Thus in region III equation (2.16) for g becomes at leading order

$$\partial_y^2 g = [k^2 + (U''_\mathrm{c}/U'_\mathrm{c})\tilde{y}^{-1} + 2\tilde{y}^{-2}]g, \tag{5.5}$$

or, with

$$w = 2k\tilde{y}, \quad \kappa = -U''_\mathrm{c}/2kU'_\mathrm{c}, \quad \mu = 3/2, \tag{5.6}$$

we obtain the standard form of the Whittaker equation (Olver, 2010)

$$\partial_w^2 g + [-\tfrac{1}{4} + \kappa w^{-1} + (\tfrac{1}{4} - \mu^2) w^{-2}] g = 0. \tag{5.7}$$

The general solution may be written

$$g = E W_{\kappa,3/2}(2k\tilde{y}) + G W_{-\kappa,\mu}(e^{\pm i\pi} 2k\tilde{y}), \tag{5.8}$$

using (13.14(v)) of Olver (2010). Either sign can be taken, giving different but related branches of the Whittaker function.

For $|z| \to \infty$, Whittaker functions have the dominant asymptotic behaviour

$$W_{\kappa,\mu}(z) \sim e^{-z/2} z^{\kappa} \qquad (-3\pi/2 < \arg z < 3\pi/2) \tag{5.9}$$

(Olver, 2010, (13.14.21)) and so in (5.8) we set $G = 0$ to match to evanescent decay in region IV (5.2). The resulting solution in (5.8) also matches to an exponential solution in region II (3.10) with $C = 0$ and this gives the leading order structure of a neutrally stable mode. However to pick up the effect of the critical layer on the stability of the mode we need also to keep track of the subdominant component of the Whittaker function. Although this is an exponentially small contribution, by stabilising or destabilising the mode, this pushes c below the real axis, $c_{\rm i} < 0$, or above, $c_{\rm i} > 0$. Similarly y_c is pushed below the real axis if $U_{\rm c}' c_{\rm i} < 0$ or above if $U_{\rm c}' c_{\rm i} > 0$. As we integrate above the critical point for reasons of causality and $U' \le 0$, we obtain a normal mode if $c_{\rm i} > 0$ and a quasi-mode if $c_{\rm i} < 0$ in what follows (cf. figure 7).

We relegate the details to appendix B and give the connection formula, from region IV to region II above the critical point in region III, written in terms of $\tilde{y}$

$$e^{-k\tilde{y}}(2k\tilde{y})^{\kappa} \quad \longrightarrow \quad e^{-k\tilde{y}}(-2k\tilde{y})^{\kappa} e^{i\pi\kappa} + i\pi\kappa\, e^{k\tilde{y}}(-2k\tilde{y})^{-\kappa} e^{i\pi\kappa}. \tag{5.10}$$

This indicates that (5.8) for $\tilde{y} > 0$ (and $G = 0$) connects to

$$g = E \exp[-k\tilde{y} + \kappa \log(-2k\tilde{y}) + i\pi\kappa] + i\pi\kappa E \exp[k\tilde{y} - \kappa \log(-2k\tilde{y}) + i\pi\kappa], \tag{5.11}$$

which includes both the dominant and subdominant components, valid for $\tilde{y} < 0$.

The approximation (5.11) needs to be matched to (3.10) and we do this first quickly and then more carefully. We let $\tilde{y} < 0$ be in an overlap region $k^{-1} \ll \tilde{y} \ll k^{-1/2}$ where both the WKBJ form (3.10) and the Whittaker function form (5.11) are valid. In (3.10) we approximate Δ by Δ_0 and evaluate

$$I_1 \equiv k \int_{y_{\rm t}}^{y} \sqrt{\Delta}\, dy \simeq k \int_{y_{\rm t}}^{y} \sqrt{\Delta_0}\, dy = V_{\rm tc} - k \int_{y}^{y_{\rm c}} \sqrt{\Delta_0}\, dy \simeq V_{\rm tc} + k\tilde{y}, \tag{5.12}$$

where

$$V_{\rm tc} = k \int_{y_{\rm t}}^{y_{\rm c}} \sqrt{\Delta_0}\, dy. \tag{5.13}$$

Comparing the leading order k terms (3.10) and (5.12) with (5.11) yields

$$C e^{V_{\rm tc}} = i\pi\kappa E, \quad D e^{-V_{\rm tc}} = E, \tag{5.14}$$

which with (3.20) gives

$$\tan\Phi = -\tfrac{1}{2} i\pi\kappa e^{-2V_{\rm tc}}. \tag{5.15}$$

This corresponds to a shift in the complex wave speed c, with

$$\tan(\Phi + \delta\Phi) \simeq \delta\Phi = \frac{\partial\Phi}{\partial c}\,\delta c, \qquad \frac{\partial\Phi}{\partial c} \simeq -F^2 k \int_0^{y_{\rm t}} (U - c)(-\Delta_0)^{-1/2}\, dy. \tag{5.16}$$

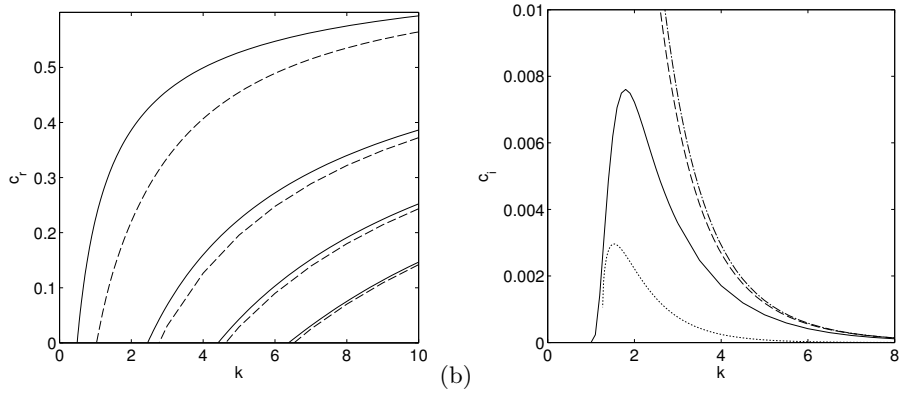


FIGURE 9. Comparison between numerical results (solid) and WKBJ theory for the profile $U_3 = 1 - \tanh(y)$. Depicted are (a) frequency c_r with formula (3.11) (dash) for the first four surface gravity wave branches, and (b) the growth rate c_i for the first branch and the critical layer term using (5.18) (dash), and the radiative term using (5.22) (dot), and combined effects (5.23) (dash–dot).

The result is a non-zero leading order value of c_i, exponentially small for large k, but that stabilises or destabilises a mode,

$$c_\mathrm{i} = -\tfrac{1}{2}\pi\kappa e^{-2V_{tc}}(\partial\Phi/\partial c)^{-1}. \tag{5.17}$$

This formula derived by WKBJ theory is equivalent to the equation (4.2) derived by considering the evolution of wave momentum, as discussed in appendix A.

In matching the solutions in (3.10) and (5.11) we ignored the prefactors and concentrated on the exponential pieces to obtain (5.14). Further analysis in appendix C confirms that this is correct at leading order, but also gives a better connection formula including a term in $k^{-1}\log k$, which is the above with V_tc replaced by

$$V'_\mathrm{tc} = V_\mathrm{tc} + \kappa\int_{y_\mathrm{t}}^{y_\mathrm{c}} \tilde{y}^{-1}(1-\Delta_0^{-1/2})\,dy + \kappa\log[2k(y_\mathrm{c}-y_\mathrm{t})]. \tag{5.18}$$

Figure 9 shows a comparison between numerical and analytical results for U_3, showing gravity wave mode branches (but with no limit mode present — see figure 8). Notably, in 9(b) there is good agreement between the growth rate obtained from (5.18) (dash) and the numerical result (solid), especially given that k is not very large. We have also checked this agreement, and likewise for figure 10, on plots of $\log|c_\mathrm{i}|$ (not shown). As the frequency c_r increases above F^{-1} the mode becomes radiative at large y, and this gives a corrected growth rate (dot-dash). As this is a small correction here and both curves are consistent with the numerical results, we defer discussion of the combined effects of critical layer and radiation until the next section.

5.3. *Profiles with unbounded velocity*

Finally we consider three profiles in which the velocity is unbounded at infinity, a linear profile and two with opposite signs of U'',

$$U_4 = 1 - y, \tag{5.19}$$

$$U_5 = 1 - y - 0.1\tanh y \qquad (U'' > 0, \kappa > 0), \tag{5.20}$$

$$U_6 = 1 - y + 0.1\tanh y \qquad (U'' < 0, \kappa < 0). \tag{5.21}$$

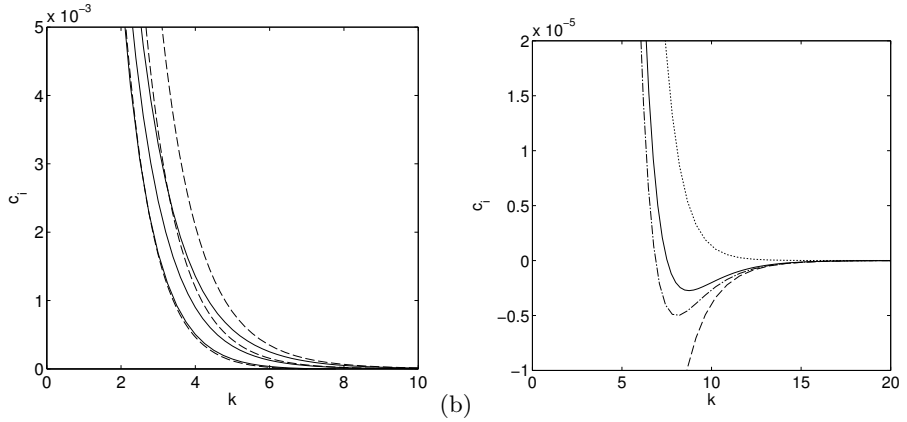


FIGURE 10. Effect of potential vorticity gradient on radiative modes: (a) growth rate c_i for U_4 (middle curves), U_5 (right curves), and U_6 (left curves) with numerical results (solid) and formula (5.23) (dash), and (b) close view for the profile U_6, with amplifying radiative term (5.22) (dot), damping critical layer term (5.17) (dash) and combined effects, (5.23) (dash–dot).

For these profiles, modes are radiative for all speeds c_r. The first, linear profile was discussed by Knessl & Keller (1995), in terms of reflection and transmission of waves incident from infinity. We will provide WKBJ formula for the growth rate including the destabilisation due to radiation. The numerical results used for comparison were obtained using the pseudo-spectral code outlined in section 5.1.

If the presence of a critical point is ignored and the WKBJ expansion matched up across the two turning points, the resulting growth rate is

$$c_\mathrm{i} = -\tfrac{1}{4} e^{-2V_{12}} (\partial\Phi/\partial c)^{-1}, \qquad V_{12} = k \int_{y_{t1}}^{y_{t2}} \sqrt{\Delta_0}\, dy. \tag{5.22}$$

For the linear profile U_4, this is the complete formula, as $U_\mathrm{c}'' = 0$ and $\kappa = 0$, and comparison with numerics is given in figure 10(a) (middle curves). This shows good agreement in this case of outwards propagating radiated waves at infinity (positive group velocity) giving a positive growth rate (noting that $\partial\Phi/\partial c < 0$ for the modes considered).

For the nonlinear profiles U_5 or U_6, with $U_c'' \neq 0$, the effects of radiation and critical layer may be summed within our perturbative expansion, as in Parras & Le Dizès (2010), giving

$$c_\mathrm{i} = \left[-\tfrac{1}{2}\pi\kappa e^{-2V'_\mathrm{tc}} - \tfrac{1}{4} e^{-2V_{12}} \right] (\partial\Phi/\partial c)^{-1}, \tag{5.23}$$

with V'_tc defined in equation (5.18). The radiative contribution is always positive (for outward waves at infinity), whereas the critical layer term is positive when $U_\mathrm{c}'' > 0$ and negative when $U_\mathrm{c}'' < 0$. These can be considered the combined effect of the two overlapping branches of continuous spectrum given in (2.25, 2.26). A parallel may also be drawn between our study and a similar result from Parras & Le Dizès (2010) for compressible round jets: U_c'' can have a stabilising or destabilising effect. However note that for stratified vortices, the equivalent of U_c'', the radial derivative of the axial vorticity at the critical point, has always a stabilising effect on the radiative instability (Schecter & Montgomery, 2004; Le Dizès & Billant, 2009).

For all three profiles U_4, U_5 and U_6, we show in figure 10(a) the combined effects of a gradient of potential vorticity and the radiative instability with good agreement as k is increased. For all of these profiles and moderate k the origin of the instability

is mainly radiative. However the critical layer contribution tends to zero more slowly than the radiative term and becomes dominant for large k, which corresponds to large speed $c_{\rm r}$, when $y_{\rm c}$ is close to the boundary. To really test the theory developed, the crucial case is when the radiative and critical layer effects have a opposite signs. This is the case for U_6 and we zoom in on the growth rate as a function of k in figure 10(b). Here the radiative instability, dominant for moderate k is taken over by critical layer stabilisation at $k = k_0 \simeq 7.52$. The agreement between the numerical results (solid) and the combined formula (dash–dot) is clear. Below k_0 the branch corresponds to an unstable normal mode, which becomes a stable quasi-mode for $k > k_0$. (Note that unlike in section 4, no radiative quasi-mode associated with an incoming wave at infinity is presented in this section as these grow spatially as $y \to \infty$ and so cannot be obtained with the pseudo-spectral code.)

We could not find an equivalent of the branch of limit modes (or quasi-modes) seen for other profiles with more pronounced curvature, for example branches for U_1 and U_2 in figure 8. These modes are evanescent at infinity and for the present profiles all modes are radiative at infinity. However there seems no reason why one could not have a limit mode that is evanescent for small y, but as y increases becomes radiative after a turning point. Such modes would be strongly damped though, for profiles with small values of U'' such as U_5 and U_6, as found for U_3. We may also speculate as to the existence of radiative wave quasi-modes when $\kappa > 0$ as for the broken profiles with $\mu < 0$ (see figure 6); we have not seen these and suspect that if they exist they must be strongly damped.

6. Conclusion

We have given an analysis of the critical layer instability and radiative instability in shallow water fluid flows, for a representative Froude number $F = 3.5$ and varying the shape of the base flow velocity profile. We have considered both piecewise (linear and nonlinear) profiles and smooth profiles, as these give different perspectives on stability in a wide range of fluid flows; for example, a piecewise linear profile, although idealised, is a useful model for understanding instabilities in flows on scales greater than that of a concentrated vorticity gradient. Numerical results were obtained for a range of profiles, with growth rates in agreement with WKBJ analysis in the limit of large wavenumber k, and linked to arguments based on evolution of wave momentum. Across regions where all waves are evanescent, effects which are exponentially small in k can damp or destabilise what would otherwise be neutral modes. In this way, a wave-free or 'balanced' shear flow can spontaneously generate waves, that is become imbalanced, through exponentially small effects, a topic recently reviewed in Vanneste (2013).

For a piecewise linear profile (Satomura, 1981), instability occurs in the form of a resonance between surface gravity wave modes and a limit (or Rayleigh) mode, whose structure has been determined. In the absence of resonance this limit mode is neutral, and its presence is not only a result of the discontinuity in the vorticity, but also of the infinite extension of the evanescent domain which follows from a finite limit $U(\infty)$ of the velocity profile. For more general profiles the neutral limit mode becomes an unstable normal mode or a damped quasi-mode depending on the sign of the gradient of vorticity in the critical layer. The mode remains as a quasi-mode for the hyperbolic tangent profile (5.1) which does not present any discontinuity. For the profile (5.2) with algebraic fall-off, any limit mode would become strongly damped and we were unable to obtain it. In this case, the surface gravity modes remain unstable but the instability becomes much weaker.

Going now to the smooth profile we find a useful description of critical layer instability of surface gravity modes in this case, as the remnant of the resonance between such modes and the limit mode in the piecewise case. In both cases it is the presence of vorticity gradients (a delta function at the discontinuity in the piecewise case) that is linked to the destabilisation of the surface gravity modes. The description of critical layer instability as the remnant of the resonance between bounded waves and the limit mode (or Rayleigh wave) is reminiscent of the instability of a piecewise linear shear layer where the resonant interaction of two Rayleigh waves gives an instability of Kelvin–Helmholtz type.

However note that the description of critical layer instability as an interaction between discrete surface gravity wave modes and an isolated quasi-mode has limited general applicability. Indeed when we consider the profiles with unbounded velocity $U(y)$ as $y \to \infty$ in part 5.3, the growth rate of the radiative modes is modified by the critical layer term. For these profiles we cannot obtain an equivalent of the quasi-mode. Maybe an equivalent would be a surface wave in the distant fluid with evanescence before the critical point. If so, the radiative instability would be explained with two effects: (i) destabilisation because of radiation at the second turning point, and (ii) the interaction with a surface wave in the far distant flow, which gives the stabilising or destabilising effect of over-reflection at the critical point. Such a description would bring into agreement two independent views of radiative instability, namely as an over-reflection process in papers such as Le Dizès & Billant (2009), or as a wave/mean flow resonance in, for example, Schecter & Montgomery (2004).

Recent works on stratified vortices show similarities with the system we have studied. For the Rankine vortex presented in Billant & Le Dizès (2009), there is also an isolated mode living on the discontinuity in vorticity. However this case is very diferent as the critical point is at a greater radius than the discontinuity and the mode does not interact with the branches of bounded surface gravity waves. Moreover there are no unstable modes that are not radiative. In addition smoothing a Rankine vortex leads to stabilisation whereas for our piecewise linear profile smoothing can lead to amplification. Recent work by Yim & Billant (2013) shows that a bending, non-radiative instability can also exist for an isolated vortex in a stratified anticyclonic fluid, and that this instability is due to a critical layer. Finally, it would be interesting to investigate the existence of critical layer instablility for other types of flows, in particular for coherent vortices in shallow water, extending the study of Ford (1994) to smooth profiles with critical layers.

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Appendix A. Formulae for growth rates

In this appendix we derive the formulae (4.2) and (4.3), and show that (4.2) and (5.17) are equivalent. Recall the definition of M in (2.11): informally introducing time-dependence

with fields proportional to $\exp(-ikct)$, this satisfies

$$2\,\frac{dM}{dt} = -\int_0^\infty (vq^* + v^*q)\,dy - \Big[uv^* + u^*v\Big]_0^\infty \equiv I + B, \tag{A 1}$$

say. The wave momentum M can change by virtue of transport of potential vorticity through the integral term I, or by radiation through the boundary term B. We consider only the former process, taking a mode that is evanescent as $y \to \infty$. We may write the first, integral term I exactly as

$$I = \int_0^\infty ik^{-1}|v|^2 Q' \left[(U-c^*)^{-1} - (U-c)^{-1}\right] dy, \tag{A 2}$$

using (2.10). There is a contribution I_d to this integral from any discontinuity in Q that may be present, and a contribution I_c from integrating above or below a pole at $y = y_\mathrm{c}$.

For a contribution from integrating around the critical point, it is crucial as to whether the singular point in the complex y-plane, y_c, lies above or below the real axis. In the former case (upper sign below) we have $\mathrm{Im}\, y_\mathrm{c} \simeq U_\mathrm{c}' c_\mathrm{i} < 0$, and we must integrate above y_c. In the latter case (lower sign), $\mathrm{Im}\, y_\mathrm{c} \simeq U_\mathrm{c}' c_\mathrm{i} > 0$, and we integrate below y_c. We write

$$(U-c^*)^{-1} - (U-c)^{-1} = \frac{-2ic_\mathrm{i}}{(U-c_\mathrm{r})^2 + c_\mathrm{i}^2} \simeq \frac{-2ic_\mathrm{i}}{U_\mathrm{c}'^2\tilde{y}^2 + c_\mathrm{i}^2}\,, \tag{A 3}$$

with $\tilde{y} = y - y_\mathrm{r}$, where $U(y_\mathrm{r}) = c_\mathrm{r}$. The latter approximation follows from the assumption that c_i is small, so that the function is sharply peaked in the vicinity of $y = y_\mathrm{r}$. As other quantities in the integral vary slowly we simply integrate this expression (evaluating other quantities at $y = y_\mathrm{c}$ at leading order) to give

$$I_\mathrm{c} = \mp \frac{2\pi |v_\mathrm{c}|^2 Q_\mathrm{c}'}{kU_\mathrm{c}'}\,. \tag{A 4}$$

Putting this into (2.12) with appropriate attention to signs gives formula (4.2).

Another contribution I_d to I in (A 2) is obtained if Q has a step jump at a location y_d, for example for the profile U_1 in (4.1). In this case we have locally $Q' \simeq Q_\mathrm{d}\,\delta(y - y_\mathrm{d})$, and (assuming the discontinuity is not too close to the critical point) the integral I includes a contribution

$$I_\mathrm{d} = \frac{2Q_\mathrm{d}|v_\mathrm{d}|^2 c_\mathrm{i}}{k(U_\mathrm{d} - c_\mathrm{r})^2}\,. \tag{A 5}$$

If we have both effects we combine (A 4,A 5) with (2.12) to give the modified formula (4.3) (noting that $U_\mathrm{d} = 0$ for our example profiles).

Finally we consider a smooth profile, and link the formula (4.2) derived above for damping connected with the critical point, to the formula (5.17) based on the full WKBJ analysis and matching. We start by evaluating (4.2) in the WKBJ framework. We need M and v_c. First, we have, exactly

$$v = iF^{-2}k^{-1}(\partial_y g + U'(U-c)^{-1}g), \tag{A 6}$$

$$u = -U'(U-c)^{-1}F^{-2}k^{-2}(\partial_y g + U'(U-c)^{-1}g) - F^{-2}g. \tag{A 7}$$

Substituting these into M (2.11) and retaining only the leading order terms for the WKBJ approximation yields, with $c \simeq c_\mathrm{r}$,

$$M = -F^{-2}\int_0^\infty (U-c_\mathrm{r})|g|^2\,dy \simeq -\tfrac{1}{2}|A|^2F^{-2}\int_0^{y_\mathrm{t}} (U-c_\mathrm{r})(-\Delta_0)^{-1/2}\,dy, \tag{A 8}$$

using the leading form (3.7), the fact that the solution is exponentially small outside region I (in which $\Delta \simeq \Delta_0$ and that the average of the modulus squared of the cosine oscillations in (3.7) is one half. These approximations are good for large k. From this we may note from (3.9) and (2.14) that

$$M \simeq \tfrac{1}{2}F^{-4}|A|^2 k^{-1}\,\partial\Phi/\partial c. \tag{A 9}$$

This completes our evaluation of M in the WKBJ framework.

Next we need v_c: we substitute g from (5.8) with $G = 0$. We then take $\kappa = 0$ as we are evaluating the growth rate by perturbing about the case where there is no critical layer. This gives

$$g \simeq EW_{0,3/2}(z) = E(1 + 2z^{-1})e^{-z/2}, \qquad z \equiv 2k\tilde{y}. \tag{A 10}$$

Substituting into (A 6) yields

$$v \simeq 2iF^{-2}E(W'_{0,3/2} + z^{-1}W_{0,3/2}) = -iF^{-2}Ee^{-z/2}, \tag{A 11}$$

and so calculating at the critical layer with $\tilde{y} = 0$, $z = 0$ gives $v_\mathrm{c} = -iF^{-2}E$. Now we substitute v_c and M from (A 9) into (4.2) (noting that κ is given in (5.6)) to obtain

$$c_\mathrm{i} = \mp\frac{2\pi\kappa}{\partial\Phi/\partial c}\frac{|E|^2}{|A|^2}. \tag{A 12}$$

Finally, in going through the turning point from region I to region II we have $D = \frac{1}{2}A\cos\Phi$ and so $|D| = \frac{1}{2}|A|$ while in (5.14) we have D in terms of E. Putting these together yields precisely (5.17).

Appendix B. Connection formulae for Whittaker functions

In this appendix we consider connection formulae for Whittaker functions, which are multiple branched. We use formulae (13.14.13) of Olver (2010) for analytic continuation,

$$(-1)^{m+1}W_{\kappa,\mu}(ze^{2mi\pi}) = a_m W_{\kappa,\mu}(z) + b_m W_{-\kappa,\mu}(ze^{i\pi}), \tag{B 1}$$

where we do not give the general forms of $a_m(\kappa,\mu)$ and $b_m(\kappa,\mu)$ here. We note for $m = 1$,

$$W_{\kappa,\mu}(ze^{2i\pi}) = a_1 W_{\kappa,\mu}(z) + b_1 W_{-\kappa,\mu}(ze^{i\pi}), \tag{B 2}$$

which is equivalent to

$$W_{\kappa,\mu}(z) = a_1 W_{\kappa,\mu}(ze^{-2i\pi}) + b_1 W_{-\kappa,\mu}(ze^{-i\pi}), \tag{B 3}$$

and to

$$W_{\kappa,\mu}(z) = c_1 W_{\kappa,\mu}(ze^{2i\pi}) + d_1 W_{-\kappa,\mu}(ze^{i\pi}), \tag{B 4}$$

with $c_1 = 1/a_1$ and $d_1 = -b_1/a_1$. For our case $\mu = 3/2$, Olver (2010) gives

$$a_1 = c_1^* = e^{2i\pi\kappa}, \quad b_1 = d_1^* = \frac{2\pi i e^{i\pi\kappa}}{\Gamma(2-\kappa)\Gamma(-1-\kappa)}. \tag{B 5}$$

Now, consider a solution

$$g_1(z) = W_{\kappa,\mu}(z) \tag{B 6}$$

of the differential equation (5.7). We have the following asymptotic estimate as $z \to \infty$

$$g_1(z) \sim e^{-z/2}z^\kappa \equiv G_1(z) \qquad (-3\pi/2 < \arg z < 3\pi/2) \tag{B 7}$$

by (13.14.21) of Olver (2010). A second solution of the differential equation may be taken in either of the forms

$$g_2^{\pm}(z) = W_{-\kappa,\mu}(ze^{\pm i\pi}). \tag{B 8}$$

These second solutions have the asymptotic behaviours

$$g_2^{\pm}(z) \sim e^{z/2}z^{-\kappa}e^{\mp i\pi\kappa} \equiv G_2(z)e^{\mp i\pi\kappa} \qquad (-3\pi/2 \mp \pi < \arg z < 3\pi/2 \mp \pi). \tag{B 9}$$

We note the presence of Stokes lines $\arg z = m\pi$ where the exponential terms in the two asymptotic forms G_1 and G_2 are maximally disparate in modulus, and anti-Stokes lines $\arg z = m\pi + \pi/2$ where the two exponentials have purely imaginary arguments.

Now suppose a boundary condition requires exponential decay in the right-half plane given by RHP $= \{z : -\pi/2 \le \arg z \le \pi/2\}$ as is the case in section 5.2. The exact solution then is $g_1(z)$ in (B 6). Suppose furthermore that we are analysing a mode with $U_c' c_i < 0$. This means that the critical point y_c in the y-plane is pushed below the real axis, and if we are taking our integral along the real y axis in search of a normal mode, this corresponds to continuing the Whittaker function above the origin in terms of z. This means increasing $\arg z$, into the left half plane defined by LHP$^+ = \{z : \pi/2 \le \arg z \le 3\pi/2\}$. As we increase the argument of z we encounter Stokes phenomenon (Berry, 1989): the subdominant asymptotic term $G_2(z)$ in (B 9) becomes 'switched on' as we cross the Stokes line $\arg z = \pi$ and takes over the solution at the next anti-Stokes line $\arg z = 3\pi/2$. We then have

$$g_1(z) \sim G_1(z) + bG_2(z) \qquad (\pi < \arg z < 2\pi), \tag{B 10}$$

where b is a Stokes multiplier. To find b we use the formula (B 3). For z in LHP$^+$ we have also $-3\pi/2 \le \arg ze^{-2i\pi} \le -\pi/2$ and $-\pi/2 \le \arg ze^{-i\pi} \le \pi/2$ so we can use (B 7) for the two terms in (B 3) to give

$$W_{\kappa,\mu}(z) \sim a_1 e^{-z/2}(ze^{-2i\pi})^{\kappa} + b_1 e^{z/2}(ze^{-i\pi})^{-\kappa} = e^{-z/2}z^{\kappa} + b_1 e^{z/2}(ze^{-i\pi})^{-\kappa}. \tag{B 11}$$

As well as the exponential growing piece in LHP$^+$ we gain an exponentially decaying piece and the Stokes multiplier is $b = b_1$.

Note that c_i is so small that the z and y axes practically coincide provided one traverses the origin in the correct sense and that it is appropriate to use half of the Stokes multiplier for the solution on the Stokes line (Berry, 1989). The resulting connection formula is

$$\text{RHP} \qquad e^{-z/2}z^{\kappa} \quad \longrightarrow \quad e^{-z/2}z^{\kappa} + \tfrac{1}{2}be^{z/2}(ze^{-i\pi})^{-\kappa} \qquad (\arg z = \pi). \tag{B 12}$$

A similar argument for integration below a critical point, $U_c' c_i > 0$, using (B 4) yields

$$\text{RHP} \qquad e^{-z/2}z^{\kappa} \quad \longrightarrow \quad e^{-z/2}z^{\kappa} + \tfrac{1}{2}de^{z/2}(ze^{i\pi})^{-\kappa} \qquad (\arg z = -\pi). \tag{B 13}$$

Here $b = d^* = b_1$ are given in (B 5) for $\mu = 3/2$. To clarify the branch of $z^{\pm\kappa}$ to be taken, we rewrite these so that the power is taken of a positive real quantity,

$$\text{RHP} \quad e^{-z/2}z^{\kappa} \quad \longrightarrow \quad e^{-z/2}(ze^{-i\pi})^{\kappa}e^{i\pi\kappa} + \tfrac{1}{2}be^{z/2}(ze^{-i\pi})^{-\kappa} \qquad (\arg z = \pi), \tag{B 14}$$

$$\text{RHP} \quad e^{-z/2}z^{\kappa} \quad \longrightarrow \quad e^{-z/2}(ze^{i\pi})^{\kappa}e^{-i\pi\kappa} + \tfrac{1}{2}de^{z/2}(ze^{i\pi})^{-\kappa} \qquad (\arg z = -\pi). \tag{B 15}$$

We note that for small κ as in our analysis in section 5.2, the Stokes multipliers $b = b_1$, $d = d_1$ in (B 5) may be taken at leading order as simply $b = d^* = 2i\pi\kappa e^{i\pi\kappa}$. With the imposition of causality or weak viscosity for flows with $U' \le 0$, as in section 4 we always integrate above a critical point, and so the connection formula used is (5.10).

Appendix C. Matching Whittaker and WKBJ solutions

Here we undertake a more careful matching of the WKBJ solution in (3.10) and the Whittaker function approximation in (5.11), and so derive (5.18). First we keep the leading order effects of Δ_1 by expanding binomially, to write in place of (5.12),

$$I_1 = k\int_{y_t}^{y}\sqrt{\Delta}\,dy = k\int_{y_t}^{y}\sqrt{\Delta_0}\,dy + k^{-1}\int_{y_t}^{y}\tfrac{1}{2}\Delta_1\Delta_0^{-1/2}\,dy + \cdots . \tag{C 1}$$

Now the first integral on the right-hand side gives $V_{tc} + k\tilde{y} + \cdots$ as before in (5.12). The latter integral on the right-hand side becomes, using the form of Δ_1 in (5.4),

$$I_2 = k^{-1}\int_{y_t}^{y}\tfrac{1}{2}\Delta_1\Delta_0^{-1/2}\,dy = \int_{y_t}^{y}(-\kappa\tilde{y}^{-1} + k^{-1}\tilde{y}^{-2})\Delta_0^{-1/2}\,dy \simeq -\kappa\int_{y_t}^{y}\tilde{y}^{-1}\Delta_0^{-1/2}\,dy, \tag{C 2}$$

neglecting the small contribution from the $\tilde{y}^{-2}$ term to the integral. The remaining integrand has a singularity of the form $\tilde{y}^{-1}$ as y approaches y_c from below (bearing in mind that $\Delta_0 \to 1$). We may subtract this off by writing

$$I_2 \simeq \kappa\int_{y_t}^{y}\tilde{y}^{-1}(1-\Delta_0^{-1/2})\,dy - \kappa\int_{y_t}^{y}\tilde{y}^{-1}\,dy. \tag{C 3}$$

Allowing the upper limit to tend to y_c in the first integral and evaluating the second (noting that $\tilde{y}$ is negative in the appropriate range) gives

$$I_2 \simeq \kappa\int_{y_t}^{y_c}\tilde{y}^{-1}(1-\Delta_0^{-1/2})\,dy - \kappa\log(-\tilde{y}) + \kappa\log(y_c - y_t). \tag{C 4}$$

Assembling these results and redoing the matching process leads to equations (5.14–5.17) again but with V_{tc} replaced by V'_{tc} in (5.18). Note that the corrections in going from V_{tc} to V'_{tc} go to zero as $k^{-1}\log k$ for large k, thus justifying the form of V_{tc} as the leading order approximation. Nonetheless matching up the prefactors gives an improved approximation to c_i.

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