

1 SUPPLEMENTAL METHODS

2 EXPOSURE DATING

3 To sample the shoreline we selected a site about 280 m west of the nearest inselberg
4 (Jebel Hawaja; Figure 2) to reduce the possibility of slope wash reaching the sample site.
5 Aerial images show slope wash halos around the inselbergs and a post-depositional alluvial
6 fan 2 m thick flanking the inselberg and extending over the adjacent shoreline to at least 100
7 m. A 3 m trench (Section S2 of Williams et al. (2010)) was excavated using a bulldozer and 6
8 samples were collected at 0.5 m intervals down the pit wall. Field observations and particle
9 size analysis at site S2 (Figure 2) show a decrease in gravel content and an increase in sand
10 from the base upwards, consistent with lake retreat, a decrease in wave energy and final
11 deposition of a sandy beach at the site. A well developed soil has subsequently formed on top
12 of the shoreline sediment.

13 The granite was crushed and a coarse quartz concentrate was prepared using heavy
14 liquids and magnetic separation. Only the coarse sand (>0.5 cm) and gravel fraction of the
15 shoreline sediment was used for ^{10}Be analysis. We purified quartz and extracted ^{10}Be using
16 standard methods (Barrows et al., 2001). The surface sample was run in duplicate. Isotopic
17 ratios ($^{10}\text{Be}/\text{Be}$) were measured by accelerator mass spectrometry on the 14UD accelerator at
18 the Australian National University.

19 To investigate the potential for sequential deposition of the shoreline and to quantify
20 inheritance of cosmogenic nuclides in the sand and gravel, we used a profile dating approach.
21 Prolonged exposure of a sediment column to cosmic rays produces a characteristic
22 cosmogenic nuclide depth profile that is a function of the production rate of the nuclide,
23 sediment density, the accumulation or erosion rate of the sediment, and inherited nuclides in
24 the sediment from previous exposure at the sediment source. Ideal conditions for exposure

25 dating a profile are rapid deposition, an absence of post-depositional accumulation or erosion
26 of the surface, and rapid erosion rates at the sediment source, which delivers a minimal
27 inherited nuclide component. The setting of the Jebelein shoreline satisfies these suitability
28 criteria. The ridge represents a high-energy depositional environment and therefore was
29 deposited in a brief period of time. Second, there is no geomorphic evidence for widespread
30 erosion of the surface in the form of gullying. Third, there appears to have been limited post-
31 depositional accumulation of sediment on the surface. However, inheritance is likely to be
32 significant because weathering rates are low on inselbergs in desert environments (e.g.,
33 Bierman and Caffee (2002)).

34 Cosmogenic nuclide production rates are a topic of active research and as yet no
35 consensus has emerged as to a ^{10}Be production rate and a single scaling scheme that is
36 globally applicable. The CRONUS calibration data set gives a ^{10}Be production rate by
37 spallation of $4.49 \pm 0.39 \text{ atom g}^{-1}\text{a}^{-1}$ at high latitude and sea level using the (Stone, 2000)
38 scaling scheme (Balco et al., 2008). However, recent calibrations indicate that this rate is too
39 high by at least 10% (Balco et al., 2009; Fenton et al., 2011; Kaplan et al., 2011; Putnam et
40 al., 2010). In the absence of a regional production rate calibration, we used the North
41 American regional production rate of Balco et al. (2009) which is at a similar elevation (<500
42 m) and within uncertainties is the same as the other recent Northern hemisphere calibration
43 (Fenton et al., 2011). It also has a longer calibration exposure age incorporating more
44 geomagnetic variability. To scale the production rates we use the scaling scheme of (Lifton et
45 al., 2005). Barrows et al. (2013) have shown that the Lifton et al. (2005) scheme produced the
46 best match between ^{10}Be and radiocarbon dates in New Zealand. The calculated ^{10}Be
47 concentrations, ages and weathering rates are given in Supplemental Table 2.

48 To determine the exposure age for the profile, we updated the coefficients of the
49 equations used in Granger and Smith (2000) for the ^{10}Be production rate of Balco et al.

50 (2009) from that of Stone (2000). Maximum likely inheritance on the gravel was determined
51 by processing three samples from the top of the adjacent Jebel Hawaja. The morphology of
52 the inselbergs at Jebelein, including exfoliation and solution pits, indicates a low long-term
53 weathering rate.

54 The uppermost part of the section appears to be reworked. OSL dating of fine sand in
55 the surface soil at a depth of 56 cm yielded an apparent age of 3.0-3.1 ka. Fine sand from 106
56 cm showed very high dispersion and had equivalent dose values at about 1.5 Gy and 120 Gy,
57 corresponding to ages between 2.0 ka and 158 ka (Williams et al., 2010). Additionally a
58 portable dolerite murhakah (lower grinding stone) was found within the soil after excavation,
59 but from an unknown depth. The age of this grindstone is difficult to determine because their
60 use extends from the upper Palaeolithic through to the recent. Dozens of grinding holes and
61 thick camp site hearth deposits of charcoal and broken bone line the edges of nearby
62 inselbergs adjacent to the Nile River and indicate an important site of Neolithic habitation
63 during a time when grassland was more widespread, probably before 4,500 ¹⁴C yr BP. We
64 therefore consider it likely that the grindstone was incorporated into the profile during the
65 mid Holocene and therefore the upper part of the sediment profile was probably reworked at
66 this time. It is also possible that some slope wash from the inselberg has reached the site and
67 lowered the mean ¹⁰Be concentration of the surface sediment. For these reasons we chose to
68 analyse the coarse sand and gravel which has a much lower potential for translocation both
69 up and down the profile. Our logging showed that original bedding of the coarse clasts is
70 intact below the surface of the soil (Fig. 2).

71 We used the CRONUS online calculator (Balco et al. 2008) to calculate the ¹⁰Be
72 production rate at the site. The exposure age (together with inheritance) was calculated from
73 a least squares fit of a curve that is a sum of exponentials representing spallation, slow muon

74 capture and fast muons (Granger and Smith, 2000). Because the relative errors are similar on
75 each sample, we did not weight the fit.

76 The density of the sediment varies through the profile and so it was not possible to
77 determine an integrated mean density for use in the calculations. The density of sand and
78 gravel mixtures ranges from 1.6-2.1 g/cm³, so we chose a density of 1.9 g/cm³. Variations of
79 0.1 g/cm³ in density change the exposure age by ~2 %. A density of 2.7 g/cm³ was used for
80 the granite samples. The quoted uncertainties are 68% confidence limits incorporating all
81 known sources of measurement error, and include fully propagated production rate errors.

82 **OPTICALLY STIMULATED LUMINESCENCE**

83 Seven samples were collected for optically stimulated luminescence (OSL) dating
84 from channel sands and overlying sediments, and one sample was collected from the modern
85 White Nile to assess the extent to which the OSL signal is reset at deposition in this
86 environment. Samples were collected by inserting an opaque tube into the sediment. This was
87 then sealed at both ends and wrapped in black plastic. In the laboratory the material from
88 both ends of the tube, which would have been exposed to light, was scraped out and used for
89 environmental dose rate measurements.

90 Samples were processed either at Aberystwyth University or at the University of
91 Adelaide. The comparability of the two sets of analysis has previously been demonstrated
92 (Macklin et al., 2013). Samples were prepared using similar procedures in both laboratories.
93 Samples were initially treated with hydrochloric acid to digest carbonates and hydrogen
94 peroxide to remove any small amounts of organic material. Grains in the ranges 180–212 µm
95 (Aberystwyth) and 125-180 µm (Adelaide) were selected by dry sieving. Solutions of sodium
96 polytungstate were used to isolate grains with a density between 2.62 and 2.70 g/cm³. Grains
97 were etched for 40 minutes in 40% hydrofluoric acid to remove feldspar and other non-quartz

98 material and to remove the outer 9 μm of the quartz grains. This eliminates the effect of alpha
99 particle contribution to the irradiation of the grains. The material was then re-sieved to
100 remove any remaining feldspar grains. All the laboratory procedures were carried out in
101 subdued red light.

102 In Adelaide, uranium (U), thorium (Th) and potassium (K) concentrations were
103 measured by Genalysis Laboratories, Perth, using Inductively Coupled Plasma Mass
104 Spectroscopy (ICPMS) for U and Th, and Inductively Coupled Plasma Atomic Emission
105 Spectrometry (ICPAES) for K. U and Th concentrations were also measured by thick source
106 alpha counting (TSAC) in the Adelaide laboratory, and K concentrations by CSIRO
107 laboratories, Adelaide, using X-ray fluorescence spectroscopy (XRS). The cosmic ray
108 contribution was obtained using the relationship between cosmic ray penetration, depth and
109 latitude determined by (Prescott and Hutton, 1994). The radionuclide concentrations obtained
110 by the various methods were averaged, and the dose rate to be used in the age calculation, as
111 well as the beta and gamma contributions to this dose rate, was determined using the AGE
112 program of Grün (see Grün (2009)).

113 In Aberystwyth, beta dose rate was determined using a Risø beta counter (Bøtter-
114 Jensen and Mejdahl, 1988), and the concentration of U, Th and K in the samples was
115 determined using geochemical methods. For each sample, a finely milled sub-sample was
116 fused using a sodium metaborate flux and put into solution. The resulting solutions were
117 analysed using ICP-OES to determine the concentration of K and ICP-MS to determine the
118 concentration of U and Th. The conversion factors of Adamiec and Aitken (1998) were then
119 used to calculate the gamma dose rate. The cosmic ray contribution was obtained using the
120 relationship between cosmic ray penetration, depth and latitude determined by Prescott and
121 Hutton (1994). Water content of the material as received was measured by weighing, drying
122 overnight at 180°C, and re-weighing. A value of $5 \pm 2\%$ was used as the best estimate of the

123 water content during the period of burial for all samples except the sample from the modern
124 White Nile where a value of $10 \pm 5\%$ was used (Note: an increase of 1% in the water content
125 causes an increase of approximately 1% in the measured age of the sample). The results of
126 these measurements, and the calculated dose rates, are given in Table S3.

127 Luminescence measurements for sample S10/5-2 were carried out at the University of
128 Adelaide, and the other six samples including the surface sample were measured at the
129 Aberystwyth Luminescence Research Laboratory (ALRL) at Aberystwyth University.

130 *Luminescence measurements in Adelaide*

131 In Adelaide luminescence measurements were made on small aliquots consisting of
132 between about 50 to 100 grains in the 125-180 μm range. Pilot studies had indicated that this
133 number of grains would enable individual differences between the aliquots to be detected,
134 while still providing measurable luminescence. Luminescence measurements were carried out
135 using a Risø TL/OSL-DA-20 reader. Radiation was applied using a calibrated $^{90}\text{Sr}/^{90}\text{Y}$ β
136 source delivering a dose rate of 0.135 Gy/sec. Luminescence was stimulated using blue light-
137 emitting diodes (wavelength 470 nm) and detected using an EMI 9235QB photomultiplier and
138 Hoya U340 filter. The measurement procedure followed the single aliquot regeneration
139 (SAR) protocol described by Murray and Wintle (2000). Four regeneration doses were used,
140 selected on the basis of a pilot study, plus a zero dose point to test for recuperation, a repeat
141 of the first dose to test for acceptable recycling and a second repeat of the first dose followed
142 by exposure to infra-red radiation to test for the absence of feldspars (IR depletion test). 24
143 aliquots were run, and those having recycling ratios greater than 0.85, recuperation less than
144 0.15 and IR depletion ratio less than 0.15 were accepted for analysis. 23 out of the 24 aliquots
145 passed the screening criteria. The equivalent dose D_e was determined for each aliquot by
146 interpolation of its dose response (luminescence vs dose) curve. The equivalent dose is the

147 laboratory-applied dose that produces luminescence equal to that of the natural sample and is
148 hence equal to the total burial dose.

149 The individual D_e values showed a fairly well constrained distribution with six clearly
150 outlying high values. The D_e values were combined, omitting the outliers, using the central
151 age model (CAM) to derive a final D_e value for age calculation.

152 *Luminescence measurements in Aberystwyth*

153 In Aberystwyth luminescence measurements were made on aliquots containing ~500
154 grains of quartz between 180-211 μm in diameter. A Risø TL/OSL reader model TL-DA-15
155 was used for measurements. Blue LEDs (470 Δ 20 nm) were used for optical stimulation, and
156 the resulting OSL was detected using an EMI 9635QA photomultiplier fitted with 7.5mm of
157 Hoya U-340 filter. The resulting OSL signals showed a rapid initial decrease in intensity (Fig.
158 S1a) as would be expected from samples where the OSL signal is dominated by the fast
159 component. As in Adelaide, the SAR method was used, and typically 6 to 7 regeneration
160 doses were used to characterise the dose response curve and calculate the equivalent dose
161 (D_e) for each aliquot (Fig. S1b). Uncertainties on individual D_e values were determined using
162 the approach outlined in Duller (2007) which incorporates the uncertainty due to counting
163 statistics on each individual OSL decay curve, the uncertainty in the curve fitting procedure,
164 and also a 1.5% uncertainty associated with the Risø instruments used for these
165 measurements. Twenty-four replicate measurements of D_e were made for each sample. Data
166 was screened using standard acceptance criteria, including (1) a recycling ratio within 10% of
167 unity, (2) an IR depletion ratio (Duller, 2003) within 10% of unity, (3) a net OSL signal from
168 the test dose that is at least three times larger than the standard deviation of the background,
169 (4) a dose response curve that increases monotonically and (5) a natural signal (and
170 associated uncertainty) that intersects with the dose response curve. This screening removed
171 very few aliquots, with between 21 and 24 aliquots passing all of these criteria.

172 The modern sample from the White Nile (S11/WN1) gave an age of 0 ± 21 years,
173 demonstrating that quartz is very effectively reset by exposure to daylight at the time of
174 deposition in this environment. One of the limitations of luminescence dating is the impact of
175 saturation of the OSL signal at high doses. The form of the dose response curve (Fig. S1b) is
176 either a single saturating exponential, or the sum of two saturating exponential functions. As
177 the radiation dose to which a sample is exposed increases, the rate of growth of the OSL
178 signal decreases, until at some point there is negligible growth. Beyond this point it is not
179 possible to resolve samples of different age. A number of the samples in this study have large
180 D_e values (> 100 Gy, Table S4), but no more than 2 aliquots from 24 measured for each
181 sample failed because the OSL signal was saturated. However, one feature of the samples
182 with a high D_e (> 100 Gy) is that they tend to have dose distributions (Fig. S1c) with
183 overdispersion values (25-30%) slightly higher than is normally encountered in younger
184 samples, and this is thought to result from curvature of the dose response curve. There is no
185 evidence that this overdispersion is related to incomplete resetting of the luminescence signal
186 at deposition, and the age of 0 ± 21 years for the modern sample supports this. Therefore, the
187 D_e values for each sample were combined using the central age model (CAM) to derive a
188 final D_e value for age calculation. The equivalent doses, dose rates and ages for all the
189 luminescence samples are summarised in Table S4.

190 **LAKE RECONSTRUCTION**

191 The SRTM90m digital elevation model (DEM) released by NASA has a 3 arc second
192 horizontal resolution and a ~ 5 m vertical precision in this region, and was used to map the
193 White Nile River valley in ARCGIS (Fig. 1). Palaeoshorelines are obvious in many areas and
194 were mapped directly from the DEM. The height of the shoreline drops by ~ 5 m over a
195 distance of 400 km, indicating a gradient of $\sim 1:80,000$. This is due to incomplete damming
196 of the lake. Contours were generated at 395 m in the lower catchment and 400 m in the upper

197 catchment, which showed the best fit to the top of the shorelines. This allowed for the
198 shoreline of the entire lake to be extrapolated to where shorelines were absent. The area and
199 volume of the palaeolake were calculated in GIS by filling the valley to the appropriate
200 height, whilst taking into account the underlying topography.

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273 **Supplemental Table S1. Site data**

| Sample | Longitude (°N) | Latitude (°E) | Altitude (m) | Horizon correction | Thickness/depth (cm) ¹ | Muon production rate (atoms g ⁻¹ yr ⁻¹) |
|---------------------------|-------------------|------------------|-----------------|-----------------------|--------------------------------------|---|
| <i>Jebel Hawaja</i> | | | | | | |
| JEB-01 | 12° 35' 3.3" | 32° 50' 8.1" | 426 | 1 | 3.0 | 0.210 |
| JEB-02 | 12° 35' 3.6" | 32° 50' 7.1" | 427 | 1 | 2.5 | 0.210 |
| JEB-03 | 12° 35' 4.0" | 32° 50' 6.8" | 425 | 1 | 3.0 | 0.210 |
| <i>Jebelein shoreline</i> | | | | | | |
| JEB-S06 | 12° 35' 7.5" | 32° 49' 55.9" | 402 | 1 | 2.5 | 0.209 |
| JEB-S05 | 12° 35' 7.5" | 32° 49' 55.9" | | 1 | 55 | |
| JEB-S04 | 12° 35' 7.5" | 32° 49' 55.9" | | 1 | 105 | |
| JEB-S03 | 12° 35' 7.5" | 32° 49' 55.9" | | 1 | 165 | |
| JEB-S02 | 12° 35' 7.5" | 32° 49' 55.9" | | 1 | 205 | |
| JEB-S01 | 12° 35' 7.5" | 32° 49' 55.9" | | 1 | 275 | |

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2761. Samples at Jebel Hawaja were exposed at the surface. Samples at the Jebelein shoreline are shown in depth below the surface (± 5 cm)

277 **Supplemental Table S2. ¹⁰Be results**

| Sample | Lab code | Production Rate (atoms g ⁻¹ yr ⁻¹) Muons | [¹⁰ Be] _c (x10 ⁵ g ⁻¹) | Weathering rate (m/Ma) ¹ |
|---------------------|-------------|---|---|--|
| <i>Jebel Hawaja</i> | | | | |
| JEB-01 | ANU-M392-02 | | 25.2 ± 1.70 | 0.73 ± 0.12 |
| JEB-02 | ANU-M392-03 | | 9.90 ± 0.70 | 2.64 ± 0.33 |
| JEB-03 | ANU-M392-04 | | 6.26 ± 0.46 | 4.62 ± 0.54 |
| <i>Section S2</i> | | | | |
| JEB-S2-06a | ANU-M392-05 | | 5.01 ± 0.21 | |
| JEB-S2-06b | ANU-M430-11 | | 5.56 ± 0.26 | |
| JEB-S2-05 | ANU-M430-10 | | 5.58 ± 0.24 | |
| JEB-S2-04 | ANU-M430-09 | | 4.17 ± 0.18 | |
| JEB-S2-03 | ANU-M430-08 | | 3.72 ± 0.18 | |
| JEB-S2-02 | ANU-M430-07 | | 3.88 ± 0.18 | |
| JEB-S2-01 | ANU-M392-06 | | 3.30 ± 0.14 | |

278 Data are normalised to NIST SRM 4325 assuming ¹⁰Be/⁹Be = 2.79 x 10⁻¹¹.279 Carrier ¹⁰Be/⁹Be = <1 x 10⁻¹⁵. ¹⁰Be decay constant = 4.998 x 10⁻⁷ yr⁻¹.280 Sample density for granite = 2.7 g cm⁻³

281 1. Calculated using CRONUS online calculator and Wrapper script: 2.2, Main calculator: 2.1, Objective function: 2.0, Constants: 2.2.1, and Muons: 1.1 versions

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Supplemental Table S3. Sample data for optical ages

| Site | Sample | Depth (m) | Aliquots | H ₂ O % | K (%) | U (ppm) | Th (ppm) |
|------|---------|-------------|----------|--------------------|-------------|-------------|-------------|
| B | S11/1-1 | 1.20 ± 0.10 | 22 | 5 ± 2 | 1.10 ± 0.06 | 0.76 ± 0.04 | 2.59 ± 0.13 |
| B | S11/1-2 | 1.50 ± 0.20 | 23 | 5 ± 2 | 1.09 ± 0.06 | 0.67 ± 0.03 | 2.00 ± 0.10 |
| B | S11/1-3 | 2.35 ± 0.50 | 21 | 5 ± 2 | 0.97 ± 0.05 | 0.67 ± 0.03 | 1.82 ± 0.09 |
| C | S10/5-2 | 1.10 ± 0.10 | 22 | 5 ± 2 | 0.87 ± 0.03 | 1.03 ± 0.07 | 3.72 ± 0.24 |
| C | S10/5-3 | 1.50 ± 0.10 | 21 | 5 ± 2 | 0.90 ± 0.05 | 0.64 ± 0.03 | 1.78 ± 0.09 |
| C | S10/5-4 | 1.65 ± 0.10 | 24 | 5 ± 2 | 0.92 ± 0.05 | 0.67 ± 0.03 | 1.33 ± 0.07 |
| D | S11/3-1 | 0.80 ± 0.50 | 22 | 5 ± 2 | 0.93 ± 0.05 | 1.03 ± 0.05 | 1.14 ± 0.06 |
| E | S11/WN1 | 0.00 ± 0.10 | 24 | 10 ± 5 | 0.97 ± 0.05 | 0.56 ± 0.03 | 0.78 ± 0.04 |

Supplemental Table S4. Dose and dose rates for optical ages

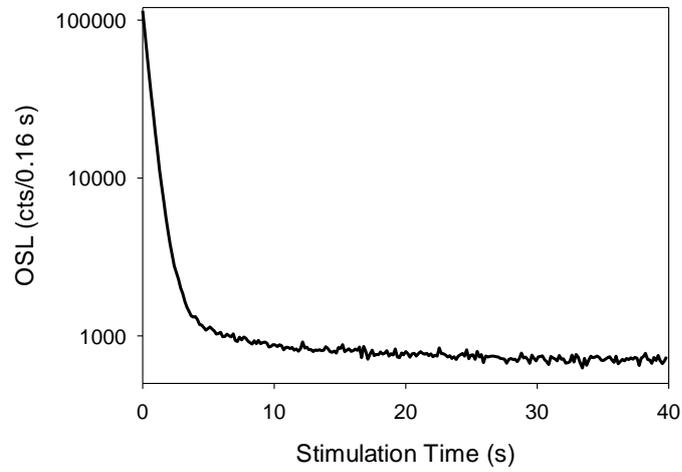
| Site | Sample | External β dose 'wet' (Gy/ka) | External γ dose 'wet' (Gy/ka) | Cosmic dose (Gy/ka) | Total dose rate (Gy/ka) | Equivalent dose (Gy) | Age (ka) |
|------|---------|---|--|---------------------------|----------------------------|-------------------------|-----------------|
| B | S11/1-1 | 0.99 ± 0.04 | 0.45 ± 0.02 | 0.18 ± 0.02 | 1.62 ± 0.05 | 112 ± 6.2 | 69 ± 4 |
| B | S11/1-2 | 0.90 ± 0.04 | 0.41 ± 0.02 | 0.17 ± 0.02 | 1.48 ± 0.04 | 130 ± 6.2 | 88 ± 5 |
| B | S11/1-3 | 0.83 ± 0.04 | 0.38 ± 0.02 | 0.16 ± 0.04 | 1.36 ± 0.06 | 140 ± 10 | 103 ± 9 |
| C | S10/5-2 | 0.79 ± 0.05 | 0.48 ± 0.05 | 0.19 ± 0.02 | 1.47 ± 0.04 | 15.3 ± 1.5 | 10.4 ± 1.1 |
| C | S10/5-3 | 0.71 ± 0.03 | 0.36 ± 0.02 | 0.17 ± 0.02 | 1.24 ± 0.03 | 141 ± 9.6 | 113 ± 8 |
| C | S10/5-4 | 0.85 ± 0.04 | 0.34 ± 0.01 | 0.17 ± 0.02 | 1.36 ± 0.04 | 156 ± 13 | 115 ± 10 |
| D | S11/3-1 | 0.80 ± 0.03 | 0.38 ± 0.02 | 0.19 ± 0.04 | 1.36 ± 0.05 | 142 ± 7.0 | 104 ± 6 |
| E | S11/WN1 | 0.65 ± 0.04 | 0.30 ± 0.02 | 0.30 ± 0.03 | 1.25 ± 0.06 | 0.00 ± 0.02 | 0.00 ± 0.02 |

Note:

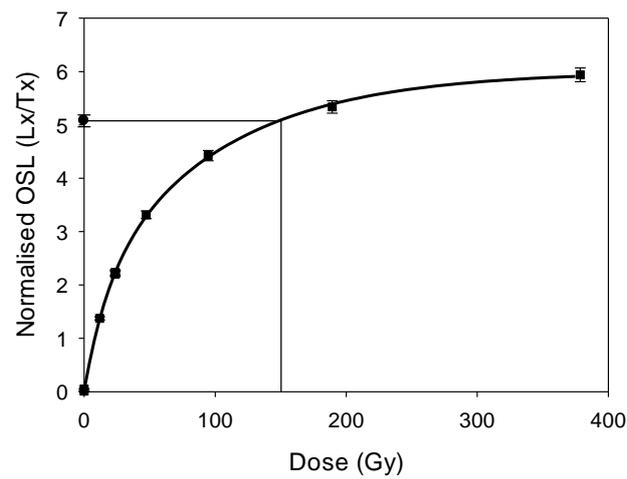
Grain Size 180-211 μ m for all samples except S10/5-2 for which grains 150-180 μ m in diameter were used.

Figure S1: Luminescence characteristics of a typical sample (S10/5-3). (a) Natural OSL decay curve. Note the log scale on the y-axis, (b) dose response curve and (c) radial plot showing the distribution of D_e values.

(a)



(b)



(c)

